

1 An Improved Flexible Spatiotemporal DATA Fusion (IFSDAF) method for  
2 producing high spatiotemporal resolution NDVI time series

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21 **ABSTRACT**

22       The Normalized Difference Vegetation Index (NDVI) is one of the mostly used  
23 vegetation index for ecosystem dynamics monitoring and biosphere process modeling.  
24 However, global NDVI products are usually provided with relatively coarse spatial  
25 resolutions, being short of important spatial details. Producing NDVI time-series data  
26 with a high spatiotemporal resolution is thus indispensable for monitoring land  
27 surface and ecosystem changes, especially in heterogeneous areas. An Improved  
28 Flexible Spatiotemporal DATA Fusion (IFSDAF) method is developed in this study to  
29 overcome the existing issues. In accordance with the distinctive characteristics of  
30 NDVI with large data variance and high spatial autocorrelation compared with raw  
31 reflectance bands, the IFSDAF method first produces temporal increment with linear  
32 unmixing and spatial-dependent increment by thin plate spline (TPS) interpolation,  
33 and then obtains final prediction from the optimal integration of two increments by  
34 Constrained Least Square (CLS) theory. Moreover, IFSDAF is developed with a  
35 capacity of employing all available and partially contaminated fine images. Coarse  
36 spatial resolution NDVI (MODIS) and fine spatial resolution NDVI images (Landsat  
37 and Sentinel) in areas with great spatial heterogeneity and significant land cover  
38 changes were used to test the performance of the new method. The promising results  
39 (RMSE 0.0884, rRMSE 22.12% in heterogeneous areas, RMSE 0.0546, rRMSE 25.77%  
40 in land cover change areas) demonstrate the strengths and robustness of the proposed  
41 method in providing reliable high spatial and temporal resolution NDVI datasets to

42 support research on land surface processes. The proposed IFSDAF method can be  
43 further simplified by only using spatial-dependent increment to improve the efficiency  
44 to a great extent. It will make IFSDAF a feasible method for applications in large  
45 geographical area and has the potential for global studies.

46

47 **Keywords:** Normalized Difference Vegetation Index (NDVI), Spatiotemporal Data  
48 Fusion, High Spatial and Temporal Resolution, Constrained Least Square (CLS)  
49 method, Weighted Integration

## 50 **1. Introduction**

51 The Normalized Difference Vegetation Index (NDVI) enhances the absorptive  
52 and reflective features of vegetation and provides a proxy for measuring canopy  
53 greenness and vigor (Rouse et al., 1974; Huete et al., 2002). Accordingly, NDVI  
54 time-series data derived from spaceborne sensors are widely employed in ecosystem  
55 dynamics monitoring and biosphere process modeling, helping to understand  
56 responses of ecosystems to climate change (Pettorelli et al., 2005). As the most  
57 significant constraint of the available NDVI time-series products (e.g., GIMMS,  
58 MODIS, SPOT VGT), coarse spatial resolutions ranging from 250 m to 8 km prevent  
59 these products from capturing spatial details necessary for monitoring land surface  
60 and ecosystem changes, especially in geographically heterogeneous areas (Gao et al.,  
61 2006; Rao et al., 2015). Producing NDVI time-series data with both high spatial and  
62 high temporal resolutions is thus critically required for such applications, raising the  
63 need for developing spatiotemporal fusion methods by blending the high frequent but  
64 low spatial resolution images (e.g., MODIS images, hereinafter referred to as coarse  
65 images) with the high spatial resolution but low frequent images (e.g., Landsat images,  
66 hereinafter referred to as fine images) (Zhu et al., 2018). Recently with emerging  
67 constellations of CubeSats and new satellite systems (e.g. Sentinel 2 with 5 day NDVI  
68 at 10 m resolution observations), new opportunities to alleviate the issue of the  
69 classical trade-off between spatial and temporal resolution is becoming hoped,  
70 however, spatiotemporal fusion is still necessary for long time series analysis as such

71 data are unavailable before 2015.

72 When using spatiotemporal fusion technology to produce NDVI data with high  
73 spatial and temporal resolutions, users need to solve the two puzzles: (I) selecting an  
74 appropriate blending strategy: Blend-then-Index (BI) or Index-then-Blend (IB), and  
75 (II) selecting a suitable and accurate spatiotemporal fusion method. For the first  
76 puzzle, recent studies (Chen et al., 2018; Jarihani et al., 2014; Tian et al., 2013) have  
77 demonstrated that the IB strategy consistently yields better or comparable results than  
78 the BI, mainly because the IB method has these advantages compared with the BI: (i)  
79 less error propagation in the blending process; (ii) less computationally expensive;  
80 and (iii) easier to clean the noises (e.g., cloud effects) on NDVI than the raw  
81 reflectance bands by the advanced filters (e.g., Chen et al., 2004). Consequently, IB is  
82 generally recommended and becomes the dominant blending strategy for producing  
83 fused NDVI products.

84 Regarding the second puzzle, a number of spatiotemporal fusion methods have  
85 been proposed and validated over past years (Zhu et al., 2018). These methods need at  
86 least one pair of cloud-free fine and coarse NDVI images at a base date and a series of  
87 coarse NDVI images at the prediction dates as the input. However, the consensus  
88 regarding the most suitable method for producing high spatiotemporal resolution  
89 NDVI data has not been reached. Generally, as a band combination index for feature  
90 enhancement, NDVI enlarges the contrast between vegetated and non-vegetated  
91 pixels and therefore displays larger spatial and temporal variance (i.e., larger

92 heterogeneity) than the raw reflectance in most satellite images. Accordingly, a  
93 suitable spatiotemporal fusion method for fusing NDVI product is supposed to satisfy  
94 the following criteria in practice: (i) obtaining good prediction in areas with large  
95 spatial and temporal variance; (ii) requiring only one pair of clear fine and coarse  
96 NDVI image at a base date, ensuring its applicability in areas with frequent cloud  
97 contamination; (iii) having a capacity to handle land cover change, such as  
98 urbanization, deforestation/reforestation, wildfires, floods and land cover transitions  
99 caused by other forces. Among the existing spatiotemporal fusion methods, the  
100 Flexible Spatiotemporal DATA Fusion method (FSDAF) (Zhu et al., 2016) is the one  
101 meeting these criteria and can be considered a potential candidate, while other  
102 existing methods fail in at least one criterion, especially the third criterion. For  
103 example, the spatial and temporal adaptive reflectance fusion model (STARFM, Gao  
104 et al., 2006), the enhanced STARFM (ESTARFM, Zhu et al., 2010), the spatial and  
105 temporal adaptive vegetation index fusion model (STAVFM, Meng et al., 2013),  
106 unmixing-based spatiotemporal reflectance fusion model (U-STFM, Huang and  
107 Zhang, 2014), NDVI linear mixing growth model (NDVI-LMGM, Rao et al., 2015),  
108 and spatial and temporal reflectance unmixing model (STRUM, Gevaert and  
109 Garcia-Haro, 2015) cannot handle land cover changes occurring between base date  
110 and prediction date. The learning-based methods, such as Sparse-representation-based  
111 spatiotemporal reflectance fusion model (SPSTFM, Huang and Song, 2012; Song and  
112 Huang, 2013), an error-bound-regularized semi-coupled dictionary learning model

113 (EBSCDM, Wu et al., 2015) and an extreme learning machine based fusion method  
114 (Liu et al., 2016) can better capture land cover change but their learning step is time  
115 consuming, and the accuracy decreases when the spatial heterogeneity is high and  
116 scale differences between coarse and fine images are large (Zhu et al., 2016).

117 FSDAF is based on the spectral unmixing analysis and further introduces thin  
118 plate spline (TPS) interpolation to capture land cover change if the change is  
119 detectable in coarse images (Zhu et al., 2016). Compared with two widely used  
120 spatiotemporal fusion methods, STAFRM algorithm (Gao et al., 2006) and an  
121 unmixing-based data fusion (UBDF) algorithm (Zurita-Milla et al., 2008), FSDAF  
122 needs the same input data as these two methods but is superior in producing more  
123 accurate predictions especially in the NIR band of heterogeneous landscapes (Table 3  
124 and Table 4 in Zhu et al., 2016). Like NDVI, the NIR band has larger spatial and  
125 temporal variances than red band, because the reflectance in NIR band generally has  
126 larger difference among different land covers than red band, and it has more significant  
127 temporal changes than red band during vegetation growth cycles. Moreover, FSDAF  
128 can capture both the gradual and abrupt land cover changes, which is an existing issue  
129 with current spatiotemporal fusion methods. Considering many advantages of FSDAF,  
130 it could be the appropriate method for producing high spatiotemporal resolution  
131 NDVI data. However, there is still space to further improve the FSDAF method. It  
132 should be noted that the FSDAF method only relies on the result of TPS interpolation  
133 to distribute residuals ( $\epsilon$ ) between prediction and true values under an assumption that

134 errors mainly depend on the landscape homogeneity. Such an assumption is very  
135 empirical and has not been demonstrated by theoretical analysis. It may be not an  
136 optimal way to distribute residuals for different scenarios. Furthermore, in practice,  
137 many available fine images are partially contaminated by clouds. Clear pixels on  
138 these partially contaminated fine images can provide significant information of  
139 temporal changes, demonstrated by STAIR method proposed by Luo et al. (2018)  
140 with better result of producing daily surface reflectance than STARFM. Consequently,  
141 using cloud-free fine images together with partially contaminated fine images will  
142 benefit spatiotemporal NDVI fusion and expand its applicability to clouded regions.  
143 Unfortunately, the FSDAF method falls short in such a capacity and is not applicable  
144 in clouded regions.

145 To address the abovementioned limitations, we propose an Improved Flexible  
146 Spatiotemporal DAta Fusion (IFSDAF) method for producing high spatiotemporal  
147 resolution NDVI time series. The IFSDAF incorporates Constrained Least Square  
148 (CLS) theory into FSDAF method, by which temporal prediction derived from  
149 unmixing procedure and spatial prediction derived from TPS interpolation are  
150 combined, thus ensuring final prediction obtained from the optimal integration of  
151 temporal and spatial predictions. Moreover, IFSDAF was developed with the capacity  
152 of employing all available and partially contaminated fine images (e.g. maximum  
153 cloud coverage is less than 70%). To validate the effectiveness of the proposed  
154 method, comparison with three popular NDVI fusion methods (i.e., NDVI-LMGM,

155 STARFM and FSDAF) under the IB strategy were performed in several experiment  
156 areas, including a site with a heterogeneous landscape, a site with abrupt land cover  
157 changes, and a site where satellite images contain a lot of clouds.

## 158 **2. Methodology**

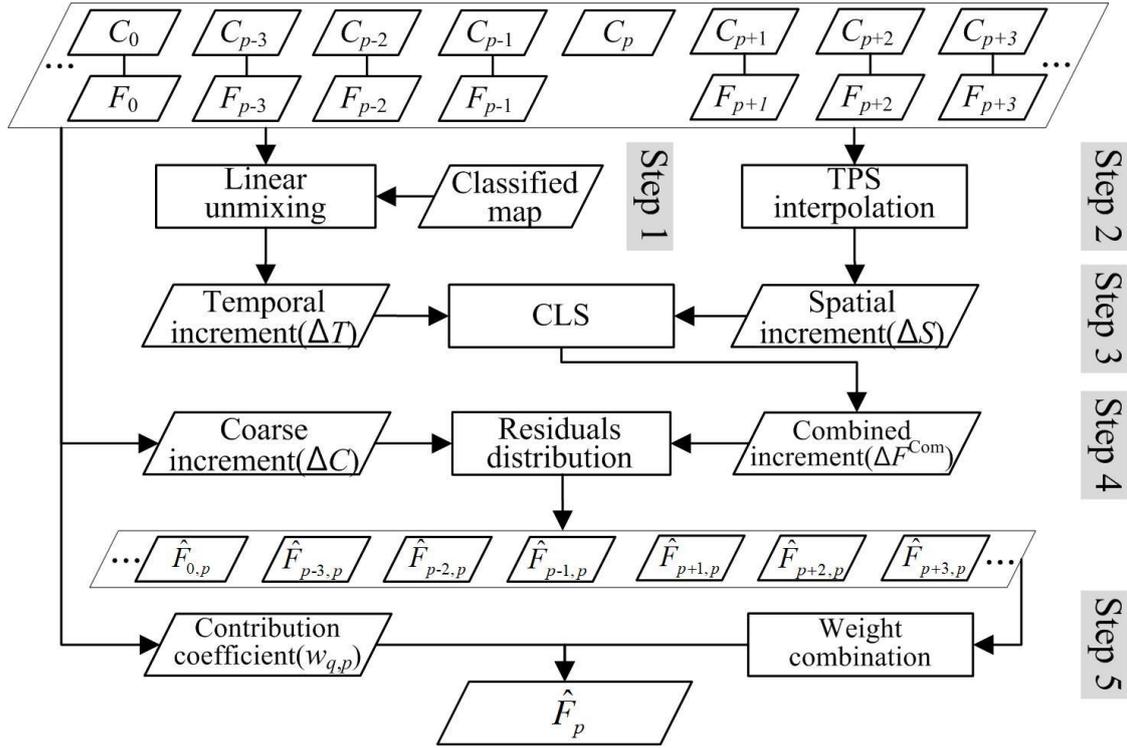
159 Although the principles of existing spatiotemporal fusion methods have a great  
160 variety, the main idea can be framed by Eq. (1), in which the fine increment of NDVI  
161 ( $\Delta F$ ) between the predicting date ( $t_p$ ) and the base date ( $t_0$ ) is firstly estimated, and then  
162 fine NDVI values ( $F_p$ ) on the predicting date ( $t_p$ ) are predicted as the sum of the base  
163 fine NDVI value ( $F_0$ ) and the increment ( $\Delta F$ ), plus the residuals  $\varepsilon$ .

$$164 \quad F_p = F_0 + \Delta F + \varepsilon \quad (1)$$

165 Given that  $F_0$  is known, IFSDAF also follows this unified equation but estimates the  
166 increment in two ways: the temporal increments using (i) unmixing analysis and  
167 spatial-dependent increments using (ii) the Thin Plate Spline (TPS) interpolation  
168 method, and then combine the two increments to obtain final  $\Delta F$  through a Constrained  
169 Least Square (CLS) method. The CLS method adopted here purifies the original  
170 FSDAF because it can adaptively combine the two increments, allowing the final  $\Delta F$   
171 approaching the one with higher accuracy.

172 The flowchart of the proposed IFSDAF is shown in Fig. 1. The input data for  
173 IFSDAF include coarse NDVI time series images and all available fine NDVI images  
174 within the same period. In these images, coarse NDVI and fine NDVI images acquired  
175 at the same dates are named as one pair. The pair with minimal cloud contaminations is

176 selected as the base images ( $C_0$  and  $F_0$ ) and its acquisition date is the base date  $t_0$ . The  
177 dates of other pairs are denoted as  $\dots, p-3, p-2, p-1, p+1, p+2, p+3, \dots$ . The coarse and  
178 fine NDVI images of these pairs are denoted as  $(\dots, C_{p-3}, C_{p-2}, C_{p-1}, C_{p+1}, C_{p+2}, C_{p+3}, \dots)$   
179 and  $(\dots, F_{p-3}, F_{p-2}, F_{p-1}, F_{p+1}, F_{p+2}, F_{p+3}, \dots)$  respectively. The task of IFSDAF is to  
180 predict fine NDVI images at any dates whenever a coarse NDVI image is available, e.g.,  
181 the date of  $t_p$ . In IFSDAF, the input fine NDVI images are not required to be cloud free  
182 except  $F_0$ . Like other spatiotemporal fusion methods, all the coarse and fine NDVI  
183 images need to be geo-registered and cropped to become the same image size. Besides,  
184 coarse NDVI time-series images need to be smoothed by an algorithm based on  
185 Savitzky-Golay filter (Chen et al., 2004), which was designed to reconstruct a  
186 high-quality NDVI time-series data by keeping the clear-sky values and interpolating  
187 clouded values. And cloud pixels in partially cloud-contaminated fine NDVI images  
188 are also masked by the Fmask algorithm (Zhu and Woodcock, 2012). A land cover  
189 classification map at a fine resolution, which can be derived from either existing land  
190 cover products (e.g. Globeland30, Chen et al., 2015) or the classification result of the  
191 input clear fine images, is needed to provide fractional cover for the unmixing process.  
192 The output of IFSDAF is synthetic fine NDVI images ( $\hat{F}_p$ ) on the prediction date  $t_p$   
193 ( $p=1, 2, 3, \dots$ ). More detailed description for each implementation step of IFSDAF is  
194 given below and a list of notations and explanation is given in Appendix.



195

196 **Fig.1.** Flowchart of the Improved Flexible Spatiotemporal DATA Fusion method

197 (IFSDAF)

## 198 2.1 Generation of temporal increment by unmixing method

199 Following the linear spectral mixing theory, the temporal NDVI change

200 (increment) of a coarse pixel can be considered as the linear combination of NDVI

201 increments of fine pixels within the coarse pixel during a short period (Rao et al.,

202 2015). Accordingly, a linear mixture model is used to unmix the increment of coarse

203 pixels from the base date  $t_0$  to the prediction date  $t_p$ , assuming that fine pixels

204 belonging to the same land cover class have a similar increment within the local region

205 (Busetto et al., 2008; Rao et al., 2015). Neighboring coarse pixels within a moving

206 window centered by coarse pixel  $(x, y)$  are used to establish a linear equation system, as

207 shown in Eq. (2).

$$\begin{matrix}
208 \\
209
\end{matrix}
\begin{bmatrix}
\Delta C(1,1) \\
\mathbf{M} \\
\Delta C(x,y) \\
\mathbf{M} \\
\Delta C(n,n)
\end{bmatrix}
=
\begin{bmatrix}
f_1(1,1) & f_2(1,1) & \mathbf{L} & f_l(1,1) \\
\mathbf{M} & \mathbf{M} & & \mathbf{M} \\
f_1(x,y) & f_2(x,y) & \mathbf{L} & f_l(x,y) \\
\mathbf{M} & \mathbf{M} & & \mathbf{M} \\
f_1(n,n) & f_2(n,n) & \mathbf{L} & f_l(n,n)
\end{bmatrix}
\begin{bmatrix}
\Delta F_1 \\
\mathbf{M} \\
\Delta F_c \\
\mathbf{M} \\
\Delta F_l
\end{bmatrix}, \quad (2)$$

$$\begin{matrix}
210
\end{matrix}
\text{with } s.t. \min(\Delta C_{\text{window}}) - \text{std}(\Delta C_{\text{window}}) \leq \Delta F_c \leq \max(\Delta C_{\text{window}}) + \text{std}(\Delta C_{\text{window}})$$

211 where  $n$  is the number of coarse pixels and  $l$  is the number of land cover classes within  
212 the moving window.  $\Delta C(x, y)$  is the NDVI increment of the coarse pixel  $(x, y)$  that can  
213 be obtained directly from coarse NDVI time series images.  $\Delta F_c$  is the fine NDVI  
214 increment of class  $c$  within the window.  $f_l(x, y)$  is the fraction of class  $l$  within the coarse  
215 pixel  $(x, y)$ , which can be obtained from the land cover map at a fine resolution.  
216  $\Delta C_{\text{window}}$  is the set of all coarse NDVI increments in the window.  $\min(\Delta C_{\text{window}})$ ,  
217  $\max(\Delta C_{\text{window}})$ , and  $\text{std}(\Delta C_{\text{window}})$  are the minimum value, maximum value and  
218 standard deviation of  $\Delta C_{\text{window}}$ , respectively. A moving window sized at a  $7 \times 7$  coarse  
219 pixel is recommended because the number of coarse pixels in the window, 49, is  
220 commonly much larger than the number of land cover classes. This choice of window  
221 size ensures the abovementioned overdetermined linear equations less influenced by  
222 collinearity and land cover changes. By solving the linear equations, the temporal  
223 NDVI increment of each class ( $\Delta F_c$ ) in the moving window can be acquired. Then, the  
224 fine temporal increment  $\Delta T(x_j, y_j)$ , where  $(x_j, y_j)$  devotes  $j$ th fine pixel in the coarse  
225 pixel  $(x, y)$ , is defined by Eq. (3), as following,

$$\begin{matrix}
226
\end{matrix}
\Delta T(x_j, y_j) = \Delta F_c \text{ if fine pixel } (x_j, y_j) \text{ belongs to class } c. \quad (3)$$

227           The fine-resolution land cover map used to compute the class fractions can be  
228 an available land cover product or classification of a cloud-free fine image. In practice,  
229 to make the fusion process automatic, existing fusion methods often use unsupervised  
230 classifiers (e.g. K-means and ISODATA) to obtain spectral classes rather than real  
231 land cover classes (Rao et.al 2015, Zhu et.al, 2017). Users need to set the number of  
232 classes in unsupervised classification. According to previous studies, the number of  
233 classes ranging from 3 to 6 could get satisfied results for most situations (Rao et.al  
234 2015, Zhu et.al, 2017). Accuracy assessment of the classification map is not included  
235 in the fusion process because: (1) aggregation of fine-scale class to coarse-scale  
236 fraction will average out some errors in classification so it may not cause large  
237 problem in solving Eq. (2); (2) temporal change assigned to a pixel with wrong class  
238 labels using Eq. (3) will be compensated by the spatial-dependent increment  
239 introduced in the next section; and (3) reference samples selection for accuracy  
240 assessment introduces more human-computer interaction. Although the proposed  
241 method is not sensitive to classification accuracy, including more accurate and robust  
242 classification methods in IFSDAF could further improve its performance.

## 243 2.2 Generation of spatial-dependent increment by TPS interpolation

244 Coarse NDVI image on  $t_p$  contains signals of land cover changes when changes  
245 are significant enough to be shown in coarse pixels. Therefore, spatial interpolation of  
246 coarse NDVI to fine resolution will retain useful information of land cover changes.  
247 Accordingly, coarse spatial resolution NDVI images on  $t_p$  and  $t_0$  are interpolated to fine  
248 spatial resolution respectively, through Thin Plate Spline (TPS) interpolation method  
249 (Chen et al., 2014; Zhu et al., 2016). TPS as a spatial interpolation technique for point  
250 data based on spatial dependence (Dubrule, 1984), is employed to obtain interpolation  
251 result thanks to its high accuracy. Then, another increment from the difference between  
252 interpolation results on  $t_p$  and  $t_0$  can be acquired. As this increment only uses spatial  
253 dependence among coarse pixels, it can be referred to as the spatial-dependent  
254 increment  $\Delta S(x_j, y_j)$ , as shown in Eq. (4), where  $F_p^{\text{TPS}}(x_j, y_j)$  and  $F_0^{\text{TPS}}(x_j, y_j)$  are TPS  
255 interpolated values on  $t_p$  and  $t_0$  respectively, and  $(x_j, y_j)$  is the  $j$ th fine pixel within the  
256 coarse pixel  $(x, y)$ .

$$257 \quad \Delta S(x_j, y_j) = F_p^{\text{TPS}}(x_j, y_j) - F_0^{\text{TPS}}(x_j, y_j) \quad (4)$$

258 Compared with the temporal increment, spatial-dependent increment has two  
259 advantages. First, coarse NDVI image on date  $t_p$  contains signals of land cover changes  
260 if the changes are significant enough to be recorded. By TPS interpolation, such land  
261 cover change information can be directly captured at a fine resolution. Second,  
262 spatial-dependent increment is independent of classification map and unmixing  
263 procedure, thus it has the potential to justify errors in the temporal increment resulted

264 from classification or unmixing. In this study, TPS is used to estimate spatial-dependent  
 265 increment rather than estimate the NDVI value on  $t_p$ , a strategy used in FADAF,  
 266 because the increment reveals the changes of NDVI directly. Zhang et al. (2015) also  
 267 suggested that using increment yields higher accuracy than predicting the value directly  
 268 at  $t_p$ . The use of this spatial-dependent increment will be further discussed in **Section 5**.

### 269 **2.3 Combination of two increments by CLS**

270 The abovementioned two increments can be considered to be two independent  
 271 predictions by two different models. Due to the distinct features used by the two  
 272 predictions, the former uses the information of temporal changes of NDVI, and the later  
 273 mainly utilizes the spatial dependence. Their prediction accuracies should be different  
 274 under different scenarios and spatial-dependencies. Therefore, it is natural to expect  
 275 that a reasonable combination of the two increments can improve the performance and  
 276 robustness of the fusion method.

277 The simplest and most effective way of combining temporal increment ( $\Delta T$ ) and  
 278 spatial-dependent increment ( $\Delta S$ ) should be summing them by reasonable weights.  
 279 Moreover, an ideal combination should be as close to the true fine NDVI increment ( $\Delta F$ )  
 280 as possible. Thus, an objective function of weighted combination can be written as,

$$281 \quad (\hat{w}_S, \hat{w}_T) = \arg \min_{(w_S, w_T) \in (0,1)} \sum_k (w_S \Delta S_k + w_T \Delta T_k - \Delta F_k)^2, \quad (5)$$

282 where  $\Delta S_k$ ,  $\Delta T_k$ , and  $\Delta F_k$  are the spatial-dependent increment, the temporal increment,  
 283 and the true increment of the  $k$ th fine pixel, respectively.  $w_S$  and  $w_T$  are weights of the  
 284 spatial-dependent increment and the temporal increment, respectively. Eq. (5) can be

285 solved by the Constrained Least Square (CLS) method, with constraints of  $w_S$  and  $w_T$   
 286 being nonnegative and summing up to one.

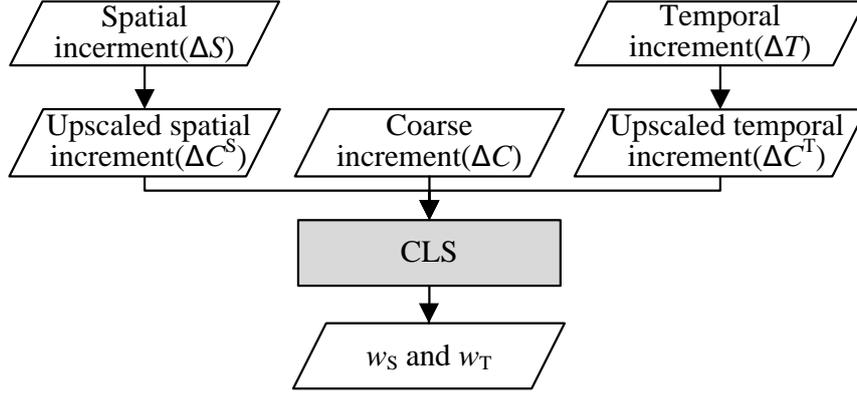
287 However, as fine NDVI values on  $t_p$  are unknown, it is impossible to obtain the  
 288 true fine increment ( $\Delta F$ ). Fortunately, a real NDVI increment of a coarse pixel ( $\Delta C$ )  
 289 from  $t_0$  to  $t_p$  is available because coarse observations are available on two dates.  
 290 Therefore, both the temporal increment and the spatial-dependent increment are  
 291 up-scaled to the resolution of coarse pixel ( $\Delta C^T$  and  $\Delta C^S$ ), as shown in Fig. 2. Then,  
 292  $w_S$  and  $w_T$  in Eq. (5) can be obtained by solving Eq. (6) alternatively:

$$293 \quad (\hat{w}_S, \hat{w}_T) = \arg \min_{(w_S, w_T) \in (0,1)} \sum_k \left( w_S \Delta C_k^S + w_T \Delta C_k^T - \Delta C_k \right)^2 \quad (6)$$

294 where  $\Delta C_k^S$ ,  $\Delta C_k^T$  and  $\Delta C_k$  are up-scaled spatial-dependent increment, up-scaled  
 295 temporal increment and true increment of  $k$ th coarse pixel, respectively. Here, the  
 296 average value of all fine NDVI pixels within the coarse pixel are used to produce  
 297 up-scaled spatial increment ( $\Delta C^S$ ) and up-scaled temporal increment ( $\Delta C^T$ ), and  $\Delta C_k$  is  
 298 calculated as difference between coarse NDVI values on prediction date  $t_p$  and  $t_0$ .  
 299 Considering that the weights  $w_S$  and  $w_T$  are spatially-dependent, Eq. (6) is solved in a  
 300  $7 \times 7$  moving window at a coarse resolution corresponding to the window size of the  
 301 unmixing process. Then, with the estimated  $w_S$  and  $w_T$ , the final fine increment can be  
 302 calculated as following:

$$303 \quad \Delta F^{\text{Com}}(x_j, y_j) = w_S \times \Delta S(x_j, y_j) + w_T \times \Delta T(x_j, y_j), \quad (7)$$

304 where  $\Delta F^{\text{Com}}(x_j, y_j)$  is the combined increment of fine pixel  $(x_j, y_j)$ .  $w_S$  and  $w_T$  are  
 305 supposed to be scale-invariant and its rationality will be discussed **Section 5**.



306

307 **Fig. 2.** Illustration of weighted calibration based on Constrained Least Square (CLS)

308 method.

### 309 2.4 Distribution of residuals

310 After CLS optimization, the combined increment can capture most of the fine  
 311 NDVI increment. However, residuals are inevitable even though they are minimal. The  
 312 residuals can be mathematically expressed as Eq. (8),

$$313 \quad R(x, y) = \Delta C(x, y) - \frac{1}{m} \sum_{j=1}^m \Delta F^{\text{Com}}(x_j, y_j), \quad (8)$$

314 where  $R(x, y)$  is the residual within a coarse pixel  $(x, y)$  and  $m$  is the number of fine  
 315 pixels within the coarse pixel. In order to further improve the accuracy of the combined  
 316 increment, residual derived above needs to be allocated to each fine pixel  $(x_j, y_j)$  within  
 317 the coarse pixel  $(x, y)$ . Because the residuals are minimal after the weighted  
 318 combination of two increments, they can be distributed equally (Chen et al., 2014) as  
 319 Eq. (9).

$$320 \quad \hat{F}_{0,p}(x_j, y_j) = F_0(x_j, y_j) + \Delta F^{\text{Com}}(x_j, y_j) + R(x, y), \quad (9)$$

321 where  $F_0(x_j, y_j)$  is fine NDVI of pixel  $(x_j, y_j)$  on date  $t_0$  and  $\hat{F}_{0,p}(x_j, y_j)$  is the predicted  
 322 fine NDVI on date  $t_p$ . After the residuals distribution, a smoothing process based on

323 similar pixels (Zhu et al., 2016) is applied to remove block effects in the fused image.

## 324 2.5 Combination of multi-time predictions

325 Through Sections 2.1 to 2.4, a prediction  $\hat{F}_{0,p}$  for date  $t_p$  based on the fine  
 326 NDVI on  $t_0$  can be acquired. In the same way, there will be several NDVI predictions,  
 327 such as ...,  $\hat{F}_{p-3,p}$ ,  $\hat{F}_{p-2,p}$ ,  $\hat{F}_{p-1,p}$ ,  $\hat{F}_{p+1,p}$ ,  $\hat{F}_{p+2,p}$ ,  $\hat{F}_{p+3,p}$ , ..., for date  $t_p$  based on  
 328 clear observations at  $p+i$  ( $i=\dots, -3, -2, -1, 1, 2, 3, \dots$ ) in other partially clouded fine  
 329 NDVI images. Recognition of a pixel is either clear or clouded can be performed  
 330 based on the Fmask algorithm (Zhu and Woodcock, 2012). Generally, the predictions  
 331 with a base date too far from  $t_p$  are excluded considering that the base NDVI images  
 332 hold weak relationship with the NDVI image on date  $t_p$ . Operationally, the maximum  
 333 interval between the base date and the prediction date is set as two months. Then, the  
 334 NDVI difference of coarse pixels between the base date and the prediction date is  
 335 used to calculate the contribution of each prediction, as shown in Eq. (10).

$$336 \quad w_{q,p}(x, y) = \frac{1}{\sum_{i=1}^9 |C_q^i(x, y) - C_p^i(x, y)|} \quad (10)$$

337 where  $C_q^i(x, y)$  and  $C_p^i(x, y)$  are coarse NDVI values of the  $i$ th pixel on base date  
 338  $q$  and the prediction date  $t_p$  in the  $3 \times 3$  moving window centered by coarse pixel  $(x, y)$ .  
 339  $w_{q,p}(x, y)$  is the contribution coefficient of predicted fine NDVI value  $\hat{F}_{q,p}(x_j, y_j)$   
 340 within the center coarse pixel  $(x, y)$ . Based on the contribution coefficient, the  
 341 combined prediction of a fine pixel  $(x_j, y_j)$  on date  $t_p$  is,

$$342 \quad \hat{F}_p(x_j, y_j) = \sum_q [w_{q,p}(x, y) \times \hat{F}_{q,p}(x_j, y_j)] / \sum_q w_{q,p}(x, y), \quad (11)$$

343 If  $C_q^i(x, y)$  equals  $C_p^i(x, y)$ ,  $\hat{F}_p(x_j, y_j)$  will be set as  $\hat{F}_{q,p}(x_j, y_j)$  since  $w_{q,p}(x, y)$  is

344 infinite under this situation. Finally, for each prediction date in the time series, a final  
345 prediction in Eq. (11) can be obtained using the same routine described in **Sections**  
346 **2.1 through 2.5.**

347 To assess the performance of the new method, four accuracy indices, Root Mean  
348 Square Error (RMSE), relative RMSE (RMSE divided by averaged observation value),  
349 Correlation Coefficient ( $r$ ) and Average Difference (AD) were used. These indices have  
350 been widely used to assess the accuracy of fused images in previous studies (e.g. Gao et  
351 al., 2006; Rao et al., 2015; Zhu et al., 2016).

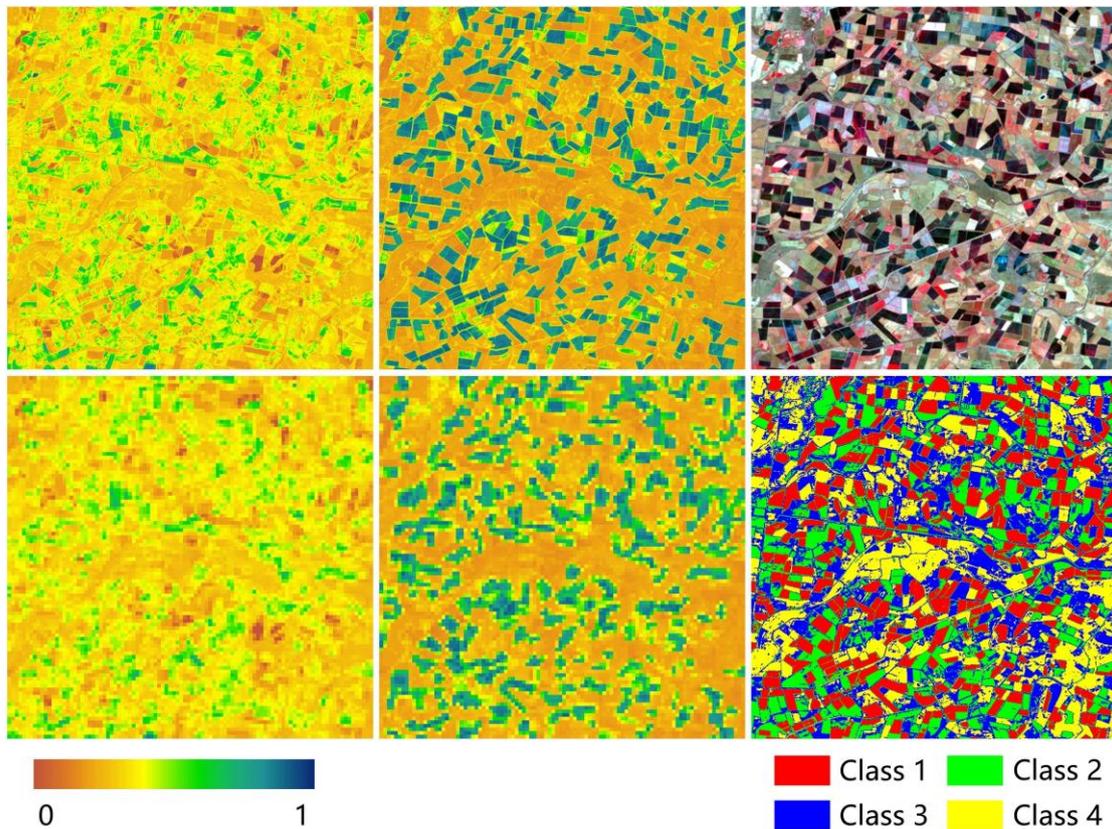
### 352 **3. Data**

#### 353 **3.1 Data for experiments using single cloud-free fine image**

354 We used Landsat images without clouded pixels to evaluate the performance of  
355 the proposed IFSDAF model at two sites with different land-cover characteristics.  
356 Considering that the performance of the existing spatiotemporal fusion methods  
357 generally perform well in homogeneous areas (Zhu et al., 2018), the study only tests  
358 the performance of the new method in cases with relative complexities (i.e., a  
359 heterogeneous site and a site with significant land cover changes.) The Landsat  
360 images covering the two sites were shared by Emelyanova et al. (2013) and were also  
361 used to test the NDVI-LMGM and FSDAF algorithms (Rao et al., 2015; Zhu et al.,  
362 2016).

363 This first site is located in the Coleambally irrigated area (34°54'S and 145°57'E),  
364 characterized by great heterogeneity in landscape with many small patches of farm

365 land and rapid phenological changes (Fig. 3). Two Landsat ETM+ images (800×800  
366 pixels), acquired on November 25<sup>th</sup>, 2001 ( $t_0$ ) and January 12<sup>th</sup>, 2002 ( $t_p$ ) during the  
367 growing season, were upscaled by the ratio of 8:1 to synthesize MODIS images. In  
368 this test, the synthesized MODIS image instead of the real MODIS image was used,  
369 because the synthesized MODIS image can exclude the co-registration error (Gevaert  
370 and Garcia-Haro, 2015; Wang and Atkinson, 2018; Zhu et al., 2016). This exclusion  
371 ensures a fair comparison of different algorithms. The NDVI data were then derived  
372 from corresponding reflectance images. Then, the land cover classification map was  
373 obtained by the Iterative Self-Organizing Data Analysis Technique (ISODATA)  
374 method based on the Landsat image acquired on November 25<sup>th</sup>, 2001 ( $t_0$ ).

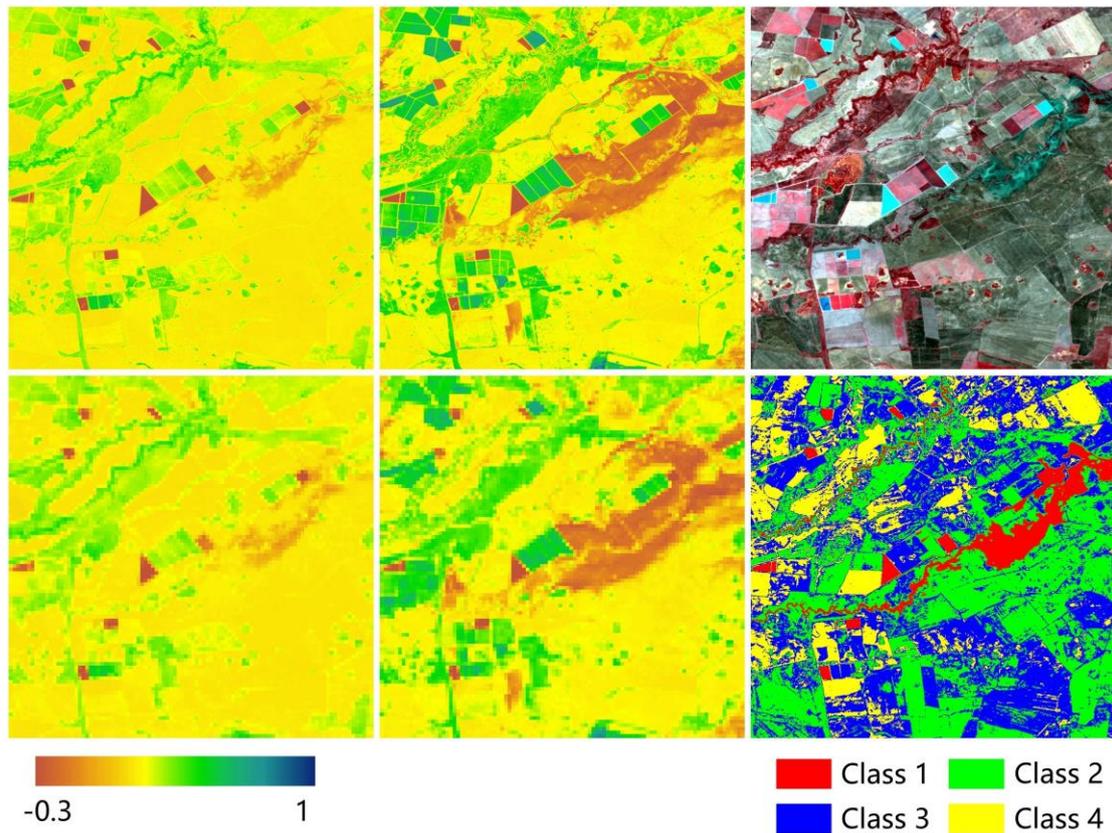


**Fig.3.** Test data of the heterogeneous site in Coleambally irrigation area: Landsat

377 NDVI on November 25<sup>th</sup>, 2001 (a) and January 12<sup>th</sup>, 2002 (b); false-color-composite  
378 Landsat image on November 25<sup>th</sup>, 2001 (c); MODIS NDVI on November 25<sup>th</sup>, 2001  
379 (d) and January 12<sup>th</sup>, 2002 (e); and land cover map on November 25<sup>th</sup>, 2001 by  
380 ISODATA (f).

381

382 The second site is located in the Gwydir area (29°07'S and 149°04'E) with a  
383 flood event occurred in December 2004. Two Landsat TM images (800×800 pixels)  
384 on November 26<sup>th</sup>, 2004 ( $t_0$ ) and December 12<sup>th</sup>, 2004 ( $t_p$ ) were used at this site (Fig.  
385 4). Abrupt land cover changes can be observed in these two images due to the flood  
386 (Emelyanova et al., 2013). Two Landsat images were also upscaled by the ratio of 8:1  
387 to synthesize the MODIS images. Then, the NDVI data were derived from all the  
388 original images. A land cover classification map was obtained based on Landsat image  
389 on November 26<sup>th</sup>, 2004 ( $t_0$ ) by the ISODATA method.



**Fig.4.** Test data of a site experienced land cover change in Gwydir area: Landsat NDVI on November 26<sup>th</sup>, 2004 (a) and December 12<sup>th</sup>, 2004 (b); false-color-composite Landsat image on November 26<sup>th</sup>, 2004 (c); MODIS NDVI on November 26<sup>th</sup>, 2004 (d) and December 12<sup>th</sup>, 2004 (e); and classification map on November 26<sup>th</sup>, 2004 by ISODATA (f).

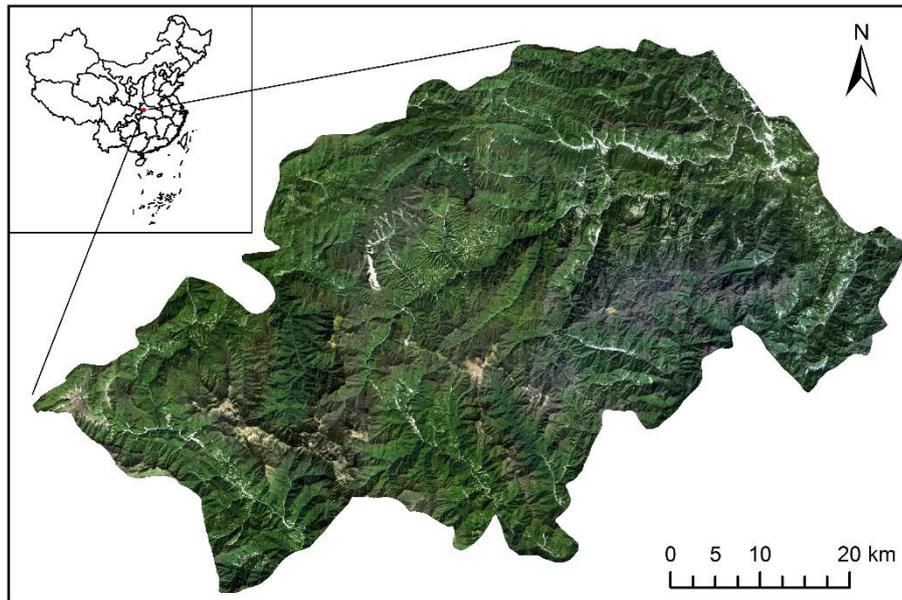
For these two sites, NDVI-LMGM (Rao et al., 2015), STARFM (Gao et al., 2006) and FSDAF (Zhu et al., 2016) were also applied to the same data set for comparison.

### 3.2 Data for experiments using multiple clouded fine images

Experiments using multiple cloudy fine images were implemented to assess the performance of the proposed IFSDAF method for predicting the NDVI time series when the input fine images which were partially contaminated by clouds. To test the

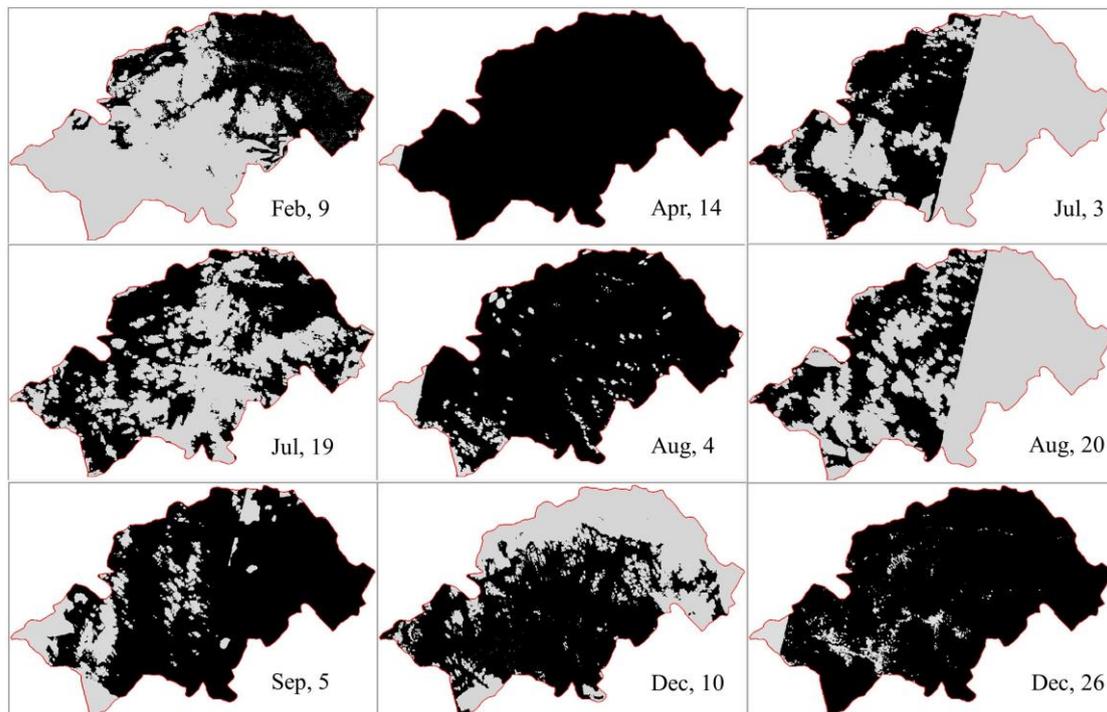
402 applicability of IFSDAF for fusing images from diverse sensors, we fused both  
403 Landsat and Sentinel-2 images with MODIS in two cloudy sites respectively.

404 The first site is Shennongjia Forestry District (109°59'–110°58' E, 31°15'–31°57'  
405 N) located in the western part of Hubei Province, central China (Fig. 5). This area  
406 belongs to subtropical monsoon climate; and its elevation ranges from 398 to 3105 m.  
407 The vegetation distribution of this area is very heterogeneous with the types of  
408 evergreen broadleaf forest, deciduous broadleaf forests, and evergreen coniferous  
409 forest. There are also farmlands and artificial surfaces in this area (Wang et al., 2018;  
410 Zhao et al., 2005). The 250 m (resampled to 240 m) 16-day composite MODIS NDVI  
411 products (MOD13Q1) covering this site in 2015 were acquired from NASA  
412 (<https://ladsweb.nascom.nasa.gov/search/>) and then resampled to the resolution of 240  
413 m. Landsat 8 level 2A surface reflectance products and their cloud masks by Fmasks  
414 in 2015 were downloaded from the USGS (<https://espa.cr.usgs.gov/ordering/new/>).  
415 All Landsat images were co-registered to MODIS images. Mosaic of two adjacent  
416 Landsat-8 scenes can cover the whole area of this site. When mosaicking two Landsat  
417 8 images with close acquisition dates, pixels in the overlapped part have two NDVI  
418 values and the higher one is kept because higher NDVI is less likely affected by poor  
419 atmospheric condition. Those Landsat images with clouds, shadows, and snow more  
420 than 70% were discarded. Finally, one clear Landsat image ( $t_0$ , on October 14<sup>th</sup>, 2015)  
421 and nine partially contaminated Landsat images (Fig. 6) were selected as the input of  
422 fine-resolution NDVI images for data fusion.



423

424 **Fig. 5.** Shennongjia Forestry District in Hubei Province, central China. The image is a  
 425 true-color-composite Landsat 8 OLI image acquired on the day of year 287, October  
 426 14<sup>th</sup>, 2015.

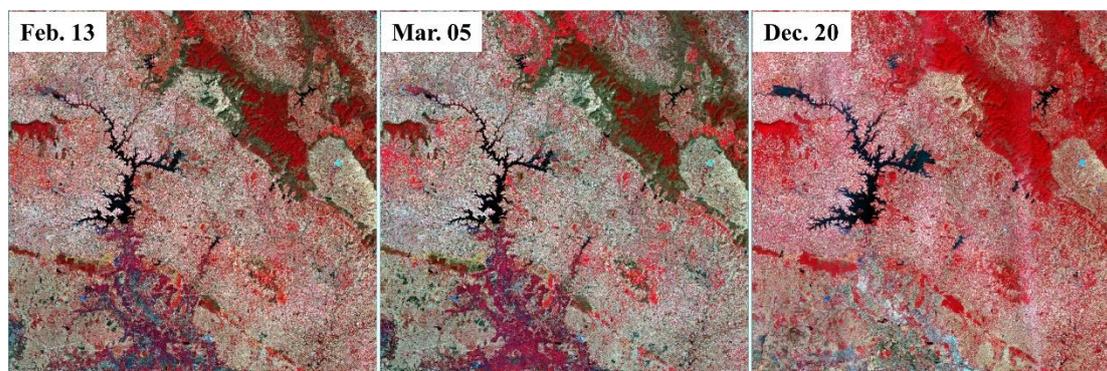


427

428 **Fig. 6.** Cloud masks of nine partially contaminated Landsat 8 images in Shennongjia  
 429 Forestry District by Fmask method, where gray color indicates pixels contaminated by  
 430 clouds and cloud shadows, and black color represents clear pixels.

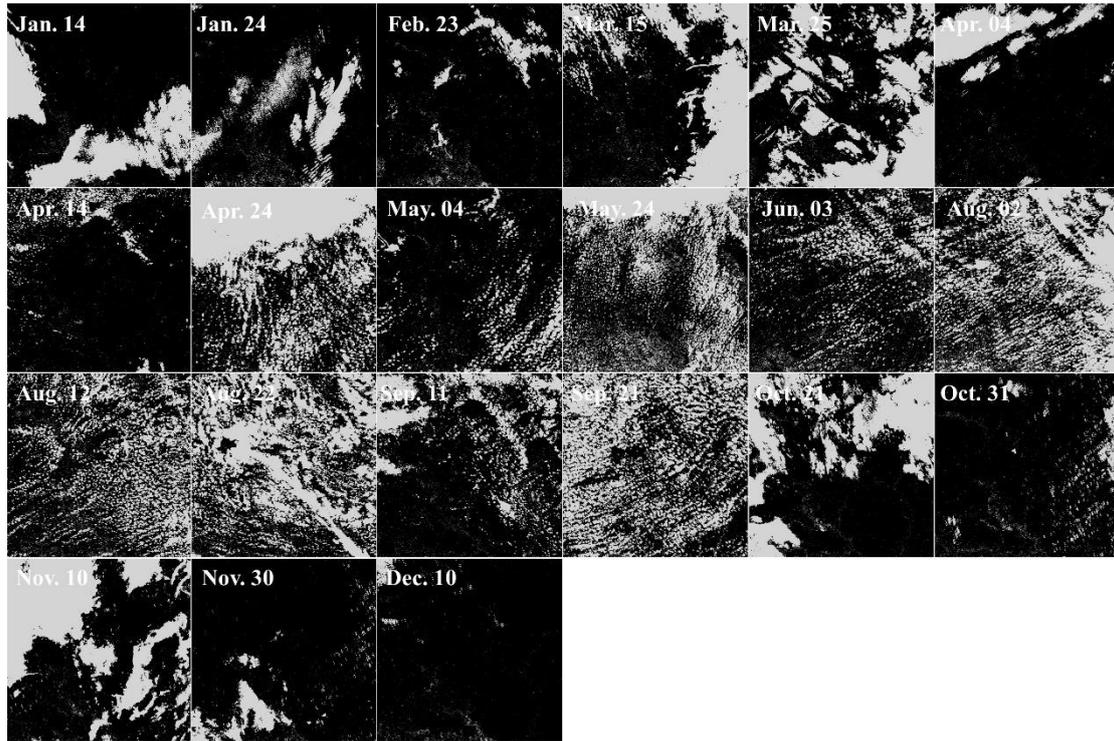
431

432 The second site is in Southeast Asia which has complex landscapes with  
433 croplands, water, forest and urbans. This site is covered by one Sentinel-2A scene  
434 (size 10980×10980 10m pixels) with the tile number of T48QUD. We acquired  
435 Sentinel-2A satellite level 1C products from EarthExplorer  
436 (<https://earthexplorer.usgs.gov/>) and 8-day composite MODIS surface reflectance  
437 products (MOD09Q1) from NASA. Both images were acquired in 2017. Atmospheric  
438 correction of Sentinel-2A images was done with the tool provided by European Space  
439 Agency, Sen2Cor (<http://step.esa.int/main/third-party-plugins-2/sen2cor/>). Cloud  
440 masks of sentinel-2A images were produced by the Fmask software  
441 (<https://github.com/gersl/fmask>) and images with cloud cover more than 70% were  
442 discarded. Finally, we obtained three clear Sentinel-2A images on Feb. 13<sup>th</sup>, Mar. 5<sup>th</sup>  
443 and Dec. 20<sup>th</sup> in 2017 respectively (Fig. 7), and 21 partially cloud contaminated  
444 images (Fig. 8). Sentinel-2A NDVI and MODIS NDVI images were then calculated  
445 from the surface reflectance data.



446

447 **Fig. 7.** False-color-composite of clear Sentinel 2A images on Feb. 13<sup>th</sup>, Mar. 05<sup>th</sup> and  
448 Nov. 20<sup>th</sup> in 2017 with the tile number of T48QUD, respectively.



449

450 **Fig. 8.** Cloud masks of 21 partially cloud contaminated Sentinel-2A images with the  
 451 tile number of T48QUD by Fmask method, where black color is clear pixel and gray  
 452 color is cloudy pixel.

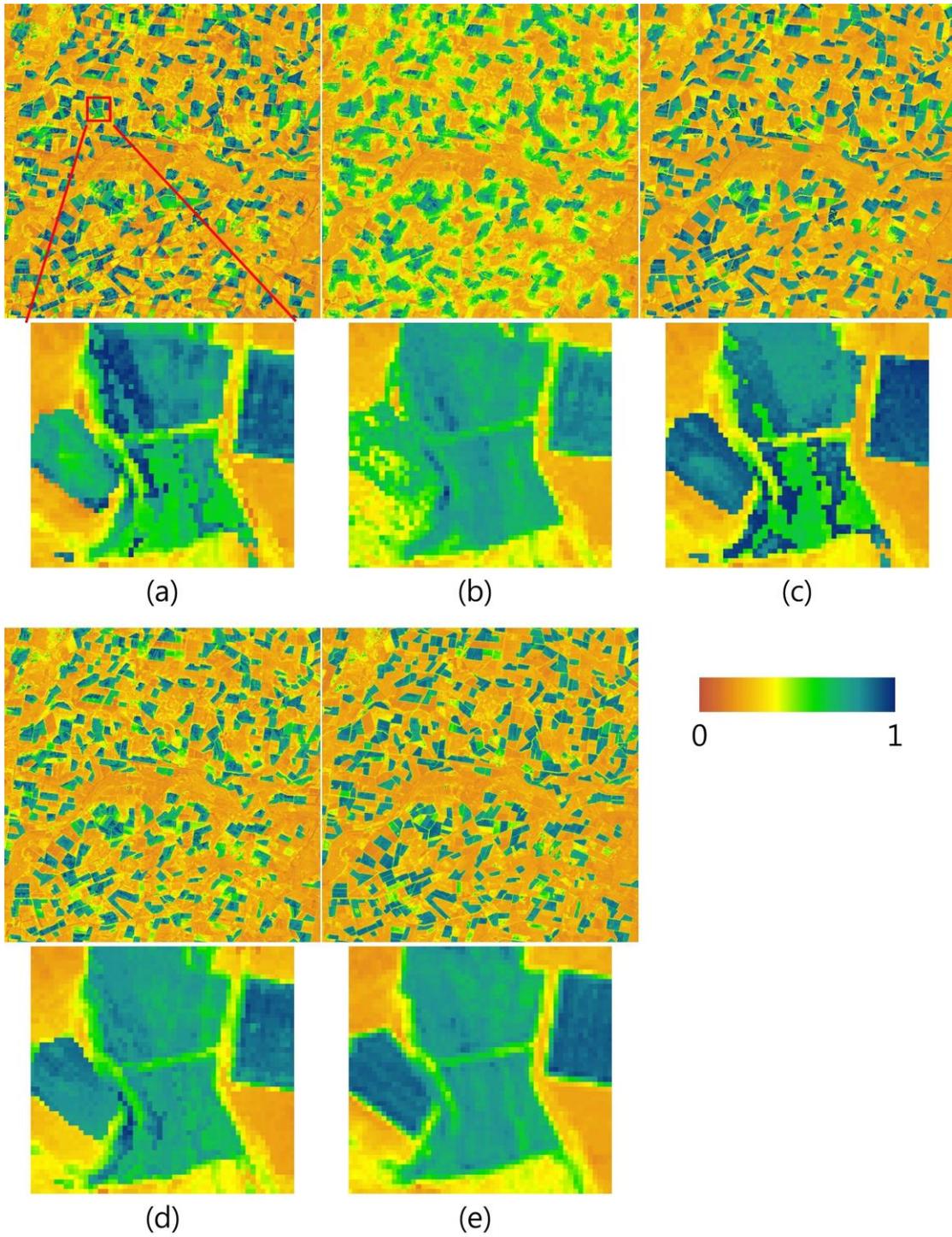
#### 453 **4. Results**

##### 454 **4.1 Fusion using single cloud-free fine image**

455 Fig. 9 shows the visual comparison of predicted Landsat NDVI by IFSDAF and  
 456 the three existing methods with observed Landsat NDVI on January 12<sup>th</sup>, 2002 ( $t_p$ ) for  
 457 the farmland site with great heterogeneity and rapid phenological changes. Compared  
 458 with the other three methods, the fused image by IFSDAF (Fig. 9d) is more similar to  
 459 the actual NDVI image (Fig. 9e) (e.g., the zoomed-in sub-region). On the contrary, the  
 460 NDVI-LMGM (Fig.9a) and FSDAF (Fig.9c) methods yield large errors in some  
 461 pixels leading to discontinuity in the fused images; and STARFM (Fig.9b) leads to an

462 unsatisfactory blurring effect for small objects. Scatter plots (Fig. 10) and quantitative  
463 assessment also confirms that the proposed method obtains the highest accuracy (Table  
464 1). The IFSDAF has the lowest RMSE (0.0884), lowest rRMSE (22.12%) and highest  $r$   
465 (0.9376) for the whole image. Furthermore, the AD (-0.0001) of the newly proposed  
466 method is closer to zero, indicating less biased. In addition, the accuracies of the  
467 NDVI-LMGM (RMSE= 0.1300 and rRMSE= 32.54%) and STARFM (RMSE= 0.1646  
468 and rRMSE= 41.19%) are much lower for the whole image compared with the FSDAF  
469 (RMSE= 0.1002 and rRMSE= 25.06%).

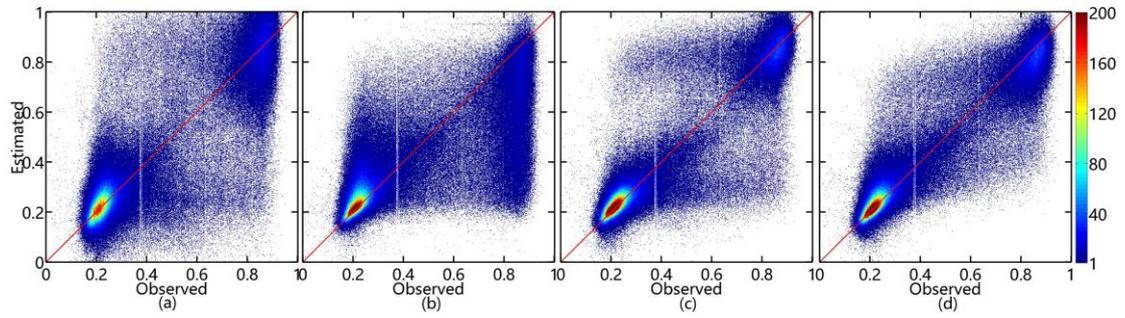
470 As we mentioned, the NDVI normally has a larger variance than raw reflectance  
471 bands. Fig. 11a shows histogram distributions of bands Red, NIR, and corresponding  
472 NDVI from Landsat image on January 12<sup>th</sup>, 2002 ( $t_p$ ). The NDVI displays two peaks  
473 with significantly a greater variance than that of band Red or band NIR due to the  
474 amplification of vegetation signals and the suppression of non-vegetation signals. To  
475 further investigate the performance of the proposed method in sub-regions with  
476 different NDVI magnitudes, based on histogram distribution of NDVI (Fig.11a), the  
477 whole image is thus divided into three parts: low NDVI (< 0.4), medium NDVI  
478 (0.4-0.7), and high NDVI (>0.7). It can be seen that NDVI-LMGM and STARFM have  
479 relatively lower accuracies compared with IFSDAF and FSDAF. And IFSDAF has a  
480 better performance than FSDAF in medium NDVI and high NDVI sections.



481

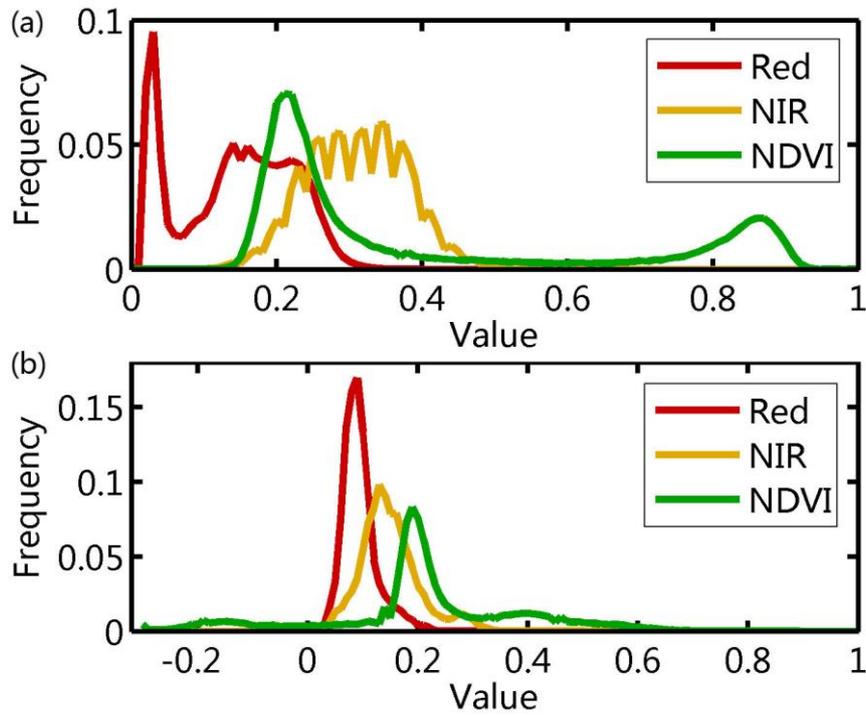
482 **Fig. 9.** Landsat NDVI on January 12th, 2002: predictions by NDVI-LMGM (a),

483 STARFM (b), FSDAF (c), IFSDAF (d) and the actual NDVI (e).



484

485 **Fig. 10.** Scatter plots of estimated results compared with observed value of Landsat  
 486 NDVI on January 12th, 2002: NDVI-LMGM (a), STARFM (b), FSDAF (c) and  
 487 IFSDAF (d).



488

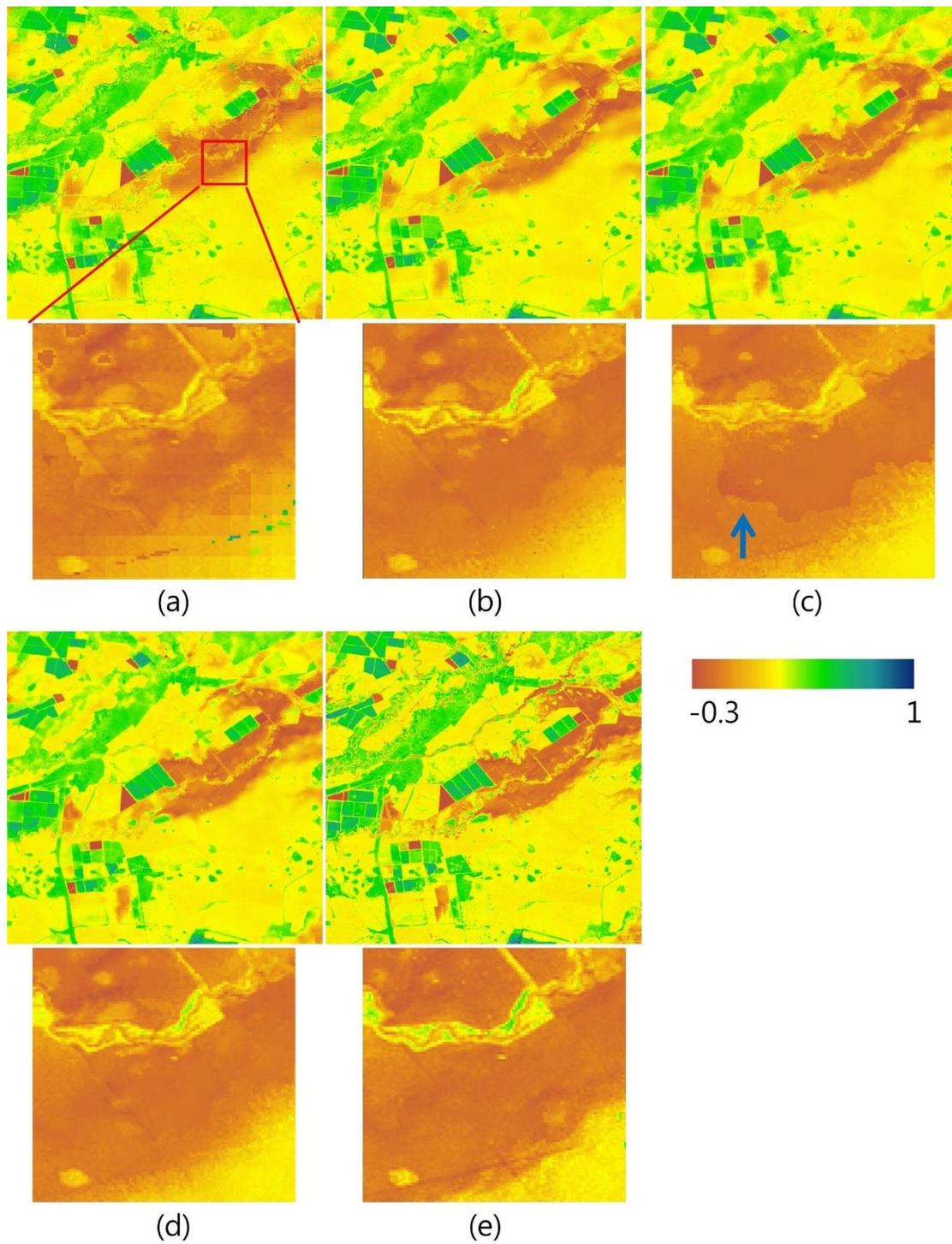
489 **Fig. 11.** Histograms of band Red, band NIR and NDVI in Coleambally irrigation area  
 490 on January 12<sup>th</sup>, 2002 (a) and the Gwydir area on December 12<sup>th</sup>, 2004 (b).

491 **Table.1.** RMSE, rRMSE,  $r$  and AD between predicted NDVI and observed NDVI of NDVI-LMGM, STARFM, FSDAF and the IFSDAF method in the  
 492 Coleambally irrigation area.

Method	NDVI-LMGM				STARFM				FSDAF				IFSDAF			
	RMSE	rRMSE	$r$	AD	RMSE	rRMSE	$r$	AD	RMSE	rRMSE	$r$	AD	RMSE	rRMSE	$r$	AD
Low NDVI	0.0916	38.85%	0.4476	<b>0.0148</b>	0.1068	45.29%	0.5214	0.0564	0.0669	28.36%	0.5576	0.0160	<b>0.0664***</b>	<b>28.16%</b>	<b>0.6334</b>	0.0170
Medium NDVI	0.2415	45.11%	0.2917	-0.0476	0.1482	27.69%	0.2805	-0.0175	0.1962	36.64%	0.3740	-0.0205	<b>0.1473***</b>	<b>27.51%</b>	<b>0.4328</b>	<b>-0.0060</b>
High NDVI	0.1589	19.03%	0.3171	-0.0477	0.2130	25.52%	0.2983	-0.1656	0.1221	14.62%	0.3926	-0.0426	<b>0.1116***</b>	<b>13.36%</b>	<b>0.4493</b>	<b>-0.0408</b>
Whole image	0.1300	32.54%	0.8744	-0.0053	0.1646	41.19%	0.7778	-0.0295	0.1002	25.06%	0.9238	-0.0012	<b>0.0884***</b>	<b>22.12%</b>	<b>0.9376</b>	<b>-0.0001</b>

493 Note: for t test, \* means  $p < 0.05$ ; \*\* means  $p < 0.01$ ; \*\*\* means  $p < 0.001$  compared with results of FSDAF.

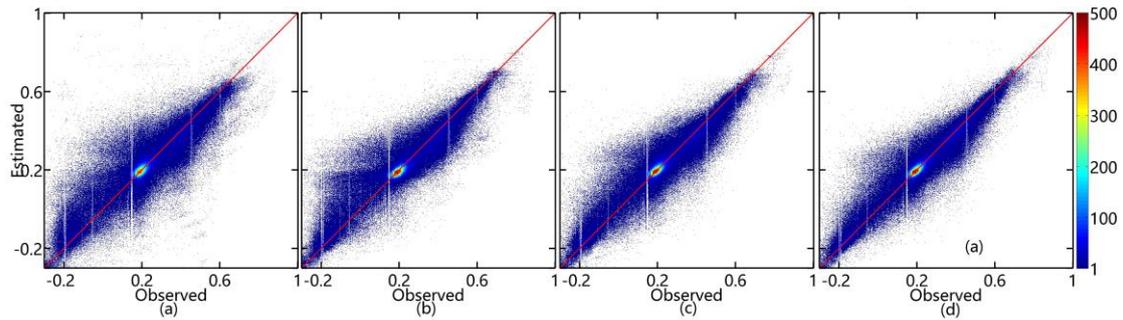
494 For the Gwydir site where a flood event occurred, as shown in Fig. 12, the fusion  
495 result of IFSDAF (RMSE = 0.0546) captures the change (Fig. 12d), being more  
496 similar to the actual NDVI pattern than the other three methods. NDVI-LMGM gets  
497 fused image (RMSE = 0.0794) with significant block effects (Fig. 12a). The result of  
498 STARFM (RMSE = 0.0686) is generally similar to the actual NDVI image as shown  
499 in Fig. 12b. FSDAF also has high accuracy (RMSE= 0.0617) in fusion but it has  
500 abnormal predictions for some pixels shown in the selected area (Fig. 12c). The blue  
501 arrow in Fig. 12c indicates the error edges produced by FSDAF. In fact, before the big  
502 flood, there is a small river in the zoomed-in area, resulting in the edge (marked by  
503 the blue arrow) between water and barren land. However, after the flood, the river  
504 overflowed and covered nearby farmland. Thus, the original edge of the river  
505 disappeared as shown in the actual Landsat NDVI image of Fig. 12e. IFSDAF is the  
506 only among the four methods which can capture this phenomenon. Scatter plots in Fig.  
507 13a-d show no obvious bias of these four methods; but points of FSDAF and  
508 IFSDAF are closer to the 1:1 line than the other two methods which reveal the  
509 comparable capacity of both IFSDAF and FSDAF in capturing land cover changes.



510

511 **Fig. 12.** Landsat NDVI on December 12<sup>th</sup>, 2004: predictions by NDVI-LMGM (a),

512 STARFM (b), FSDAF (c), IFSDAF (d) and the actual NDVI (e).



513

514 **Fig. 13.** Scatter plots of estimated results compared with observed value of Landsat  
 515 NDVI on December 12<sup>th</sup>, 2004: NDVI-LMGM (a), STARFM (b), FSDAF (c) and  
 516 IFSDAF (d).

517 Fig. 11b shows histogram distributions of band Red, band NIR, and NDVI on  
 518 December 12<sup>th</sup>, 2004 ( $t_p$ ) respectively. The variance of NDVI is also significantly  
 519 higher than that of band Red or band NIR. Moreover, due to the flood event, there are  
 520 many negative values in NDVI, causing three peaks around NDVI = -0.2, NDVI = 0.2  
 521 and NDVI = 0.4 in the histogram distribution of NDVI. The whole image is divided  
 522 into three parts (low NDVI < 0, medium NDVI 0-0.3, and high NDVI > 0.3) to  
 523 quantitatively assess the accuracy (Table 2). It is clear that the new method yields  
 524 higher accuracy with lower RMSE of 0.0546, higher  $r$  of 0.9527 among the whole  
 525 image than the other three methods. In the three separate NDVI parts, the new method  
 526 also displays higher accuracy with RMSE = 0.0798, 0.0467, and 0.0584, respectively.

527 **Table.2.** RMSE, rRMSE,  $r$  and AD between predicted NDVI and observed NDVI of NDVI-LMGM, STARFM, FSDAF and the IFSDAF method in the Gwydir  
 528 area.

Methods	NDVI-LMGM				STARFM				FSDAF				IFSDAF			
	RMSE	rRMSE	$r$	AD	RMSE	rRMSE	$r$	AD	RMSE	rRMSE	$r$	AD	RMSE	rRMSE	$r$	AD
Low NDVI	0.1267	-89.59%	0.6406	0.0708	0.0964	-68.16%	0.7364	0.0497	0.0906	-64.03%	0.7521	0.0431	<b>0.0798***</b>	<b>-56.42%</b>	<b>0.7821</b>	<b>0.0337</b>
Medium NDVI	0.0636	33.46%	0.4872	0.0061	0.0526	27.65%	0.5794	<b>0.0042</b>	0.0516	27.11%	0.6040	0.0066	<b>0.0467***</b>	<b>24.55%</b>	<b>0.6524</b>	0.0045
High NDVI	0.0858	19.67%	0.7425	-0.0442	0.0654	15.00%	0.8410	-0.0341	0.0679	15.55%	0.8302	-0.0368	<b>0.0584***</b>	<b>13.38%</b>	<b>0.8658</b>	<b>-0.0270</b>
Whole image	0.0794	37.45%	0.8970	0.0013	0.0686	29.53%	0.9250	0.0028	0.0617	29.09%	0.9395	0.0002	<b>0.0546</b>	<b>25.77%</b>	<b>0.9527</b>	<b>0.0001</b>

529 Note: for t test, \* means  $p < 0.05$ ; \*\* means  $p < 0.01$ ; \*\*\* means  $p < 0.001$  compared with results of FSDAF.

#### 530 **4.2 Fusion using multiple fine images partially covered by clouds**

531 In the site of Shennongjia, each of the four Landsat NDVI images captured on  
532 Apr 14<sup>th</sup>, Jul 3<sup>th</sup>, Sep 5<sup>th</sup>, and Dec 10<sup>th</sup> in 2015 serves as reference data for the  
533 independent validation. For example, Apr 14<sup>th</sup> was predicted by IFSDAF using all  
534 other eight partially contaminated fine NDVI images as input, and then clear pixels in  
535 the true Apr 14<sup>th</sup> image were used to assess the accuracy of predicted Apr 14<sup>th</sup> image.  
536 For the comparison, FSDAF only used one clear Landsat NDVI image on Oct. 14<sup>th</sup> of  
537 2015 to predict the above four Landsat NDVI images. The accuracies of fusion results  
538 of the four images are summarized in Table 3 and the predictions are shown in Fig. 14.  
539 For the purpose of simplification, results of NDVI-LMGM and STARFM are not  
540 shown in this experiment because they yielded lower accurate results than FSDAF.

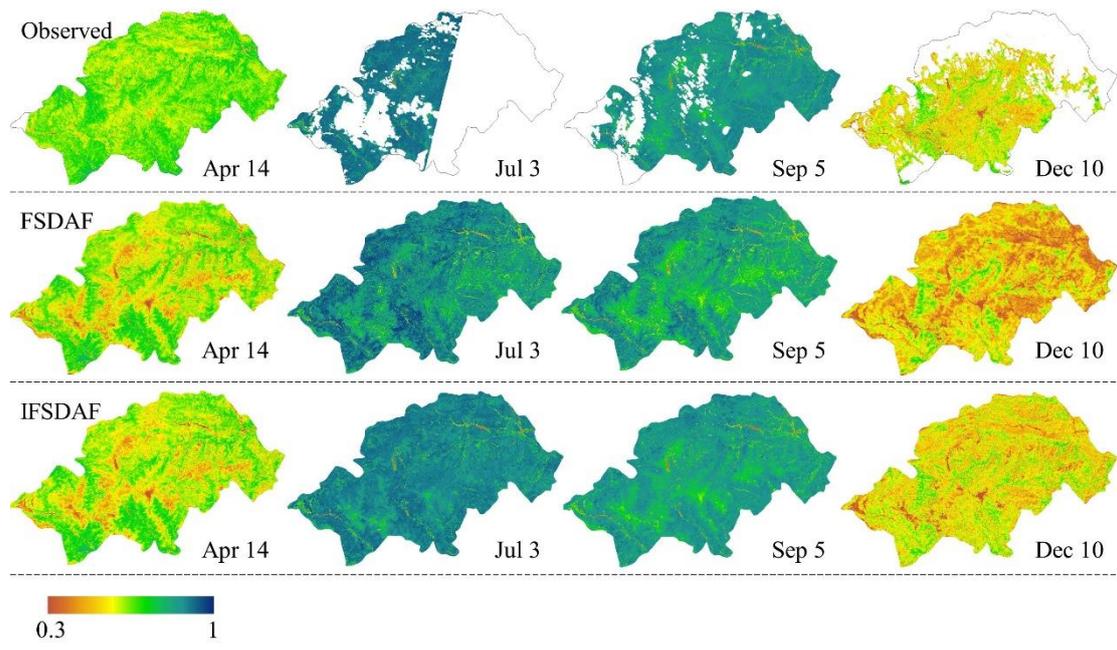
541 It is evident from Fig. 14 that IFSDAF can produce fused images more similar  
542 with real Landsat NDVI than FSDAF. In Table 3, RMSE values of IFSDAF on all  
543 dates are lower than that of FSDAF. These improvements of accuracy are mainly  
544 attributed to the extra information provided by the partially contaminated Landsat  
545 images, which can be well used in IFSDAF but not in FSDAF. On the contrary,  
546 FSDAF only used one fine image on Oct 14<sup>th</sup> in 2015 which is far away from some  
547 prediction dates, leading to low accuracy on these prediction dates. More important,  
548 the improvement of IFSDAF on Jul 3<sup>th</sup> and Sep 5<sup>th</sup> during the peak stage of vegetation  
549 growth is more significant than other two dates, indicating that IFSDAF may be more  
550 effective for fusing images with medium to high NDVI values. This result is similar to

551 the experiment in the Coleambally irrigation area.

552 **Table.3.** RMSE, rRMSE,  $r$  and AD between the predicted NDVI and observed  
 553 partially contaminated fine NDVI on Apr 14<sup>th</sup>, Jul 3<sup>th</sup>, Sep 5<sup>th</sup> and Dec 10<sup>th</sup> in year  
 554 2015, in the Shennongjia Forestry District.

Date	Methods	RMSE	rRMSE	$r$	AD
Apr 14 <sup>th</sup>	FSDAF	0.0873	13.33%	0.6319	-0.0481
	IFSDAF	<b>0.0819***</b>	<b>12.51%</b>	<b>0.6620</b>	<b>-0.0475</b>
Jul 3 <sup>th</sup>	FSDAF	0.0578	6.44%	0.6504	-0.0138
	IFSDAF	<b>0.0368***</b>	<b>4.09%</b>	<b>0.8508</b>	<b>-0.0137</b>
Sep 5 <sup>th</sup>	FSDAF	0.0671	7.86%	0.7279	-0.0306
	IFSDAF	<b>0.0393***</b>	<b>4.61%</b>	<b>0.8615</b>	<b>-0.0173</b>
Dec 10 <sup>th</sup>	FSDAF	0.1246	21.24%	0.6516	-0.0729
	IFSDAF	<b>0.0913***</b>	<b>15.57%</b>	<b>0.7768</b>	<b>-0.0366</b>

555 Note: for t test, \* means  $p < 0.05$ ; \*\* means  $p < 0.01$ ; \*\*\* means  $p < 0.001$  compared  
 556 with results of FSDAF.



557  
 558 **Fig. 14.** Landsat 8 NDVI in Shennongjia Forestry District on Apr 14<sup>th</sup>, Jul 3<sup>th</sup>, Sep 5<sup>th</sup>,  
 559 and Dec 10<sup>th</sup> in year 2015 predicted by FSDAF and IFSDAF, respectively.

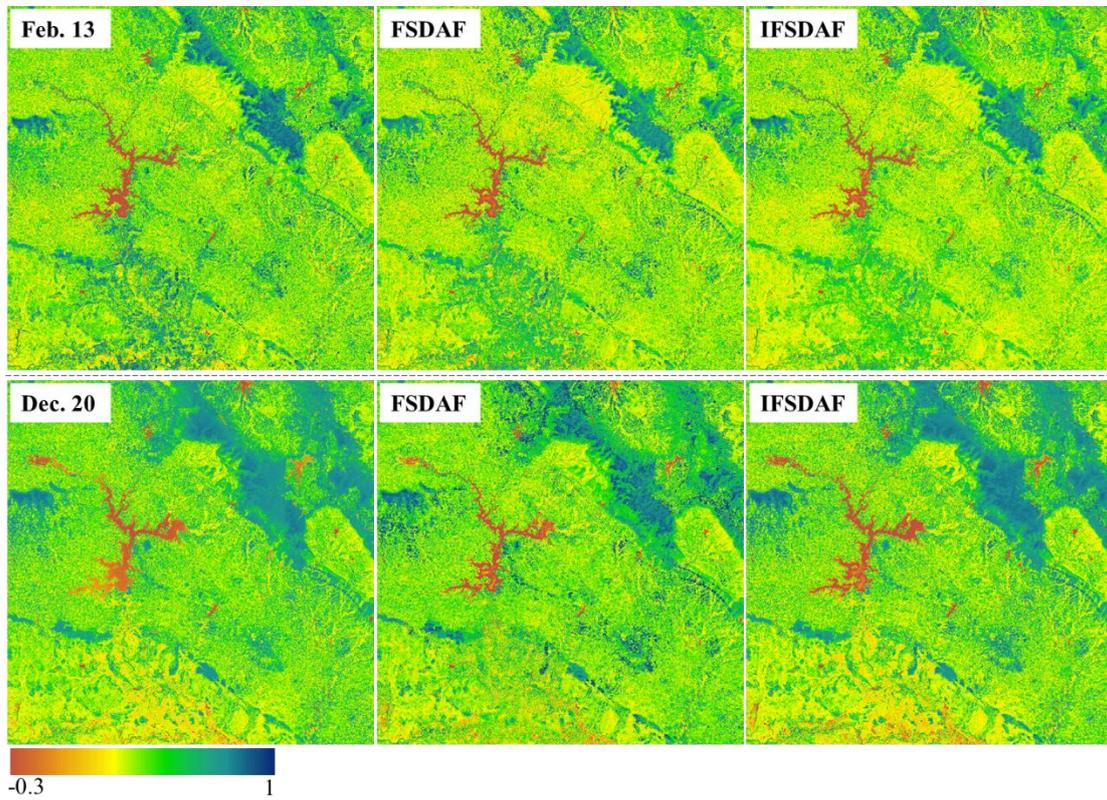
560 In the South Asia site, clear Sentinel-2A NDVI image on Mar 5<sup>th</sup>, 2017 was  
 561 selected as the base image. The other two clear Sentinel NDVI images (Feb. 12<sup>th</sup> and  
 562 Dec. 20<sup>th</sup>) were used as reference data to assess the accuracy of IFSDAF and FSDAF.  
 563 The base fine spatial resolution NDVI image and the 21 partially cloud contaminated  
 564 fine NDVI images were used as input for IFSDAF, while only the base fine NDVI  
 565 image was input to FSDAF. Results in Table. 4 shows that IFSDAF produces more  
 566 accurate predictions with lower RMSE in both dates (0.0863 and 0.0740) compared  
 567 with the results by using FSDAF (0.0999 and 0.1469).

568 **Table. 4.** RMSE, rRMSE, *r* and AD between the predicted NDVI and observed fine  
 569 NDVI on Feb. 12<sup>th</sup> and Dec. 20<sup>th</sup>, 2017 with Sentinel-2A data.

Date	Methods	RMSE	rRMSE	<i>r</i>	AD
------	---------	------	-------	----------	----

Feb 13 <sup>th</sup>	FSDAF	0.0999	24.04%	0.8885	-0.0463
	IFSDAF	<b>0.0863***</b>	<b>20.76%</b>	<b>0.9305</b>	<b>-0.0427</b>
Dec 20 <sup>th</sup>	FSDAF	0.1469	33.69%	0.7401	<b>-0.0082</b>
	IFSDAF	<b>0.0740***</b>	<b>17.69%</b>	<b>0.9584</b>	-0.0141

570 Note: for t test, \* means  $p < 0.05$ ; \*\* means  $p < 0.01$ ; \*\*\* means  $p < 0.001$  compared  
571 with results of FSDAF.



572  
573 **Fig. 15.** Sentinel-2A NDVI images on Feb. 13<sup>th</sup> and Dec. 20<sup>th</sup> (left) and the results  
574 predicted by FSDAF (middle) and IFSDAF (right), respectively.

## 575 5. Discussion

### 576 5.1 The way of deriving spatial-dependent increment

577 In this study, the spatial-dependent increment ( $\Delta S$ ) is acquired based on  
578 difference between interpolation results of coarse NDVI on date  $t_p$  and date  $t_0$ ,

579 respectively, as shown in Eq. (4). However, there is also another way of obtaining  $\Delta S$   
 580 in Eq. (12), since the  $F_0$  is available on date  $t_0$ .

$$581 \quad \Delta S(x_j, y_j) = F_p^{\text{TPS}}(x_j, y_j) - F_0(x_j, y_j) \quad (12)$$

582 where  $F_0(x_j, y_j)$  is fine NDVI value of pixel  $(x_j, y_j)$  on date  $t_0$ . However,  $\Delta S$  derived from  
 583 Eq. (4) is a better indicator than that from Eq. (12). A theoretical comparison of these  
 584 two types of  $\Delta S$  is explained as below. Eq. (9) can be simplified as shown in Equation  
 585 (13), where residuals  $R$  are ignored as they are small.

$$586 \quad \hat{F}_{0,p} = F_0 + \Delta F^{\text{Com}} = F_0 + w_S \Delta S + w_T \Delta T \quad (13)$$

587 For simplification, the notation  $(x_i, y_i)$  is removed by replacing Eq. (9) with Eq. (13).

588 And then, replacing  $F_0$  by  $w_S F_0 + w_T F_0$ , as  $w_S + w_T = 1$ , specifically,

$$589 \quad \hat{F}_{0,p} = w_S (F_0 + \Delta S) + w_T (F_0 + \Delta T) \quad (14)$$

590 Based on  $\Delta S$  in IFSDAF as Eq. (4), Eq. (14) can be written as below,

$$591 \quad \hat{F}_{0,p} = w_S (F_0 + F_p^{\text{TPS}} - F_0^{\text{TPS}}) + w_T (F_0 + \Delta T) \quad (15)$$

592 Based on  $\Delta S$  in Eq. (12), Eq. (14) can also be written as below,

$$593 \quad \begin{aligned} \hat{F}_{0,p} &= w_S (F_0 + F_p^{\text{TPS}} - F_0) + w_T (F_0 + \Delta T) \\ &= w_S (F_p^{\text{TPS}}) + w_T (F_0 + \Delta T) \end{aligned} \quad (16)$$

594 Difference between Eq. (15) and Eq. (16) is the term  $F_0 - F_0^{\text{TPS}}$  in Eq. (15). TPS

595 prediction is a spatially smoothed prediction which loses spatial details to some

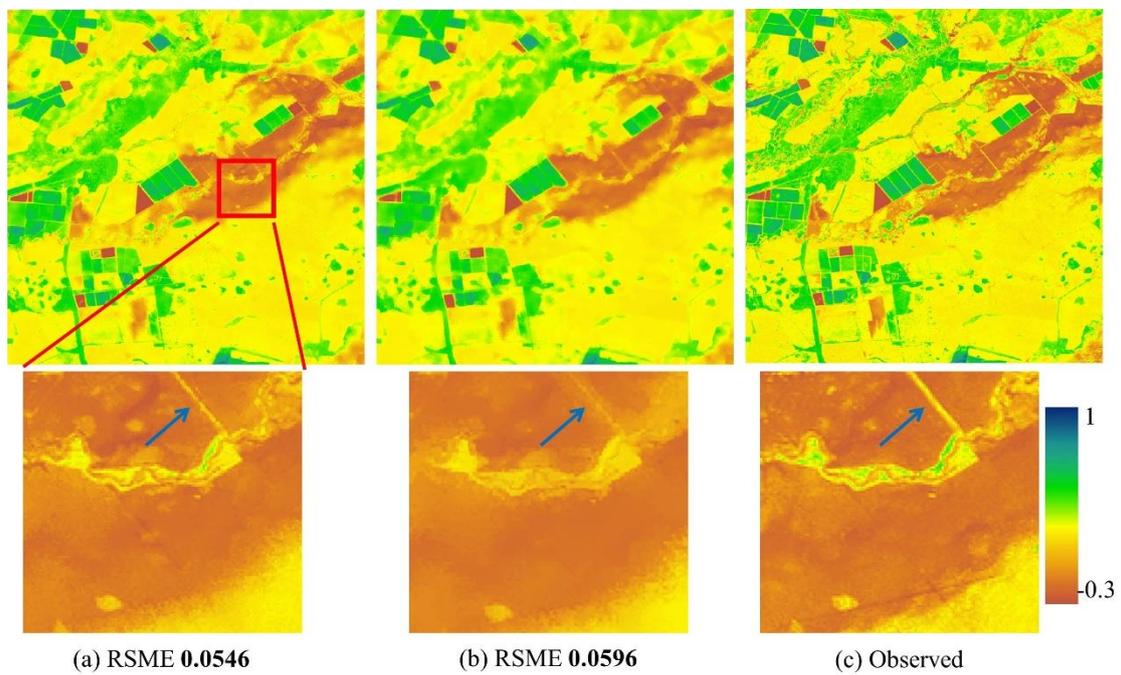
596 degree. As a result,  $F_0 - F_0^{\text{TPS}}$  functions similarly as a high-pass modulation to model

597 the spatial contrast at  $t_0$ . This spatial contrast is assumed relatively stable from the  $t_0$

598 to  $t_p$  in several fusion models (Song and Huang, 2013; Luo et al., 2018), so  $F_0 - F_0^{\text{TPS}}$

599 in Eq. (15) can better capture spatial details in the fused image. To demonstrate the

600 abovementioned theoretical analysis, an experiment based on these two types of  $\Delta S$   
 601 was conducted in the Gwydir area. Result shows that the prediction based on  $\Delta S$   
 602 derived from Eq. (4) (Fig. 16a) is more accurate than that from Eq. (12) with RMSEs  
 603 being 0.0546 and 0.0596, respectively. Moreover, as the zoomed-in pictures in Fig. 16  
 604 illustrate, the prediction result using Eq. (4) (Fig. 16a) contains more spatial details  
 605 (e.g., the road marked by blue arrows), providing corroborative support to the  
 606 analysis.



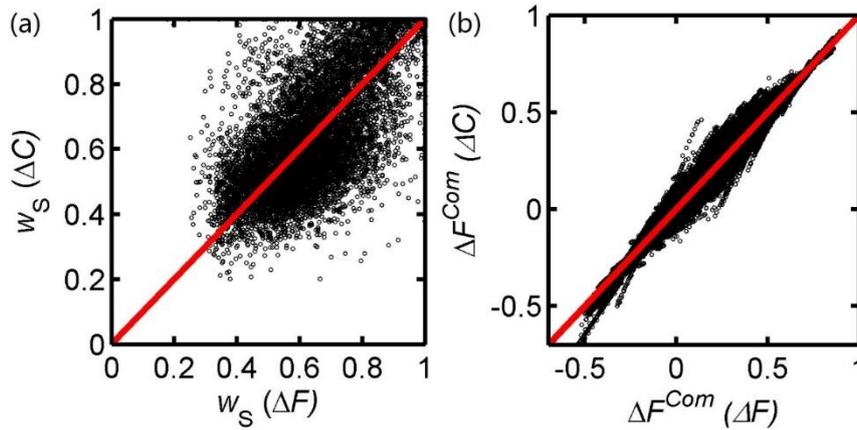
607 (a) RSME **0.0546** (b) RSME **0.0596** (c) Observed

608 **Fig. 16.** Comparison of NDVI value on  $t_p$  in the Gwydir area predicted from Eq. (4) (a)  
 609 Eq. (12) (b), and the real Landsat NDVI image(c).

## 610 5.2 The weights scale-invariant assumption

611 In the proposed IFSDAF method, spatial-dependent increment and temporal  
 612 increment are combined by optimized weights. As the fine NDVI image on  $t_p$  is  
 613 unknown in the real-world application, coarse NDVI increment ( $\Delta C$ ), upscaled

614 spatial-dependent increment ( $\Delta C^S$ ), and upscaled temporal increment ( $\Delta C^T$ ) are used  
615 to derive  $w_S$  and  $w_T$  in Eq. (5). Such operation assumes that weight is scale-invariant.  
616 To verify the assumption, an experiment was conducted in the Coleambally irrigation  
617 area, a moving window of  $7 \times 7$  at a coarse resolution was used to calculate the weights  
618 ( $w_S$  and  $w_T$ ) for two increments for the center coarse pixel. Because the fine NDVI  
619 image  $F_p$  actually exists at the site, the two weights at a fine resolution can also be  
620 derived based on fine increment ( $\Delta F = F_p - F_0$ ) using the CLS method. Fig. 17a displays  
621 the scatter plot of weights derived from the two approaches. All points are close to the  
622 1:1 line, where  $x$ -axis represents the weight of the spatial-dependent increment at the  
623 fine resolution and  $y$ -axis represents the same weight at the coarse resolution,  
624 suggesting that the weights derived from both the coarse and fine images are  
625 substitutable. Then, combined increments calculated using the two types of weight are  
626 very similar (Fig. 17b). RMSE values of combined increments based on weights from  
627 the coarse resolution and fine resolution are 0.0941 and 0.0934, respectively.  $t$ -test  
628 shows that there is no significant difference between the two combined increments.  
629 Consequently, it can be concluded that the scale effect on the derived weights is  
630 minimal, and it will not cause significant errors on the combined increment. Thus, the  
631 assumption that weights ( $w_S$  and  $w_T$ ) are scale-invariant is reasonable.



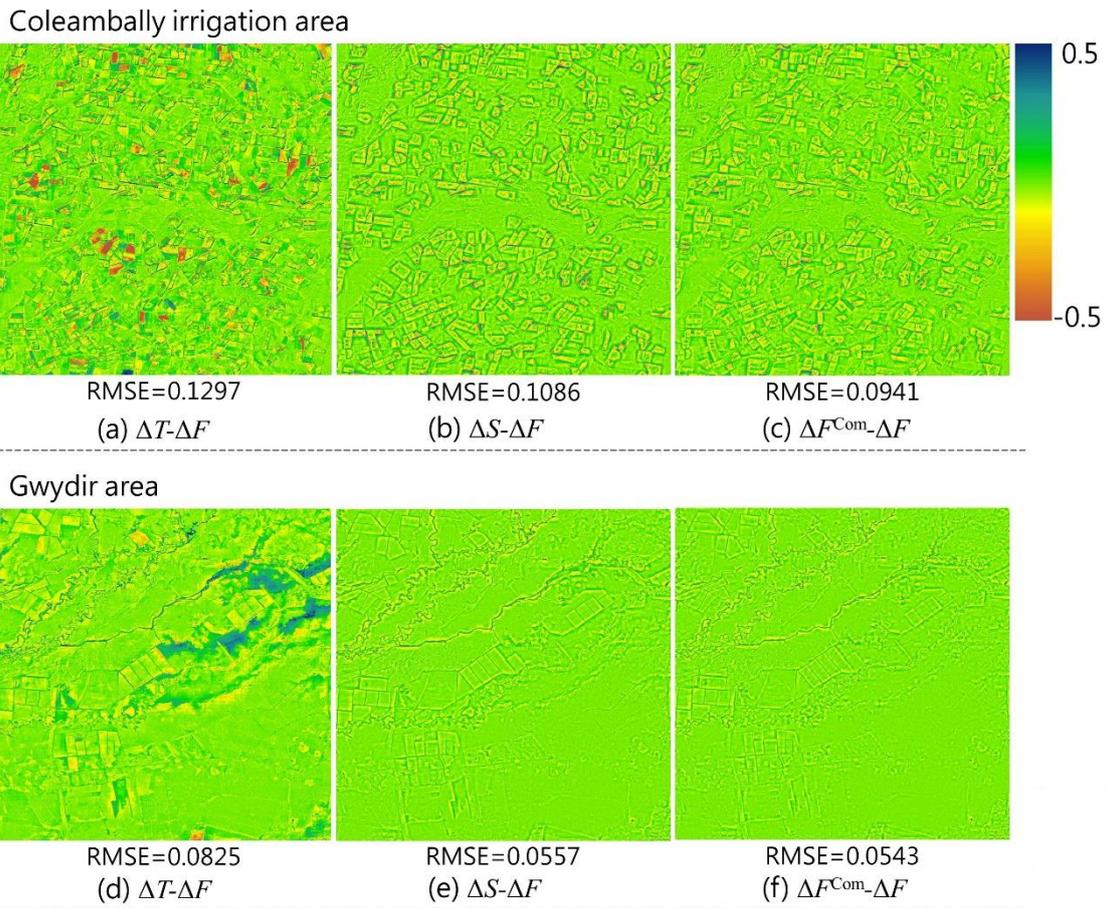
632

633 **Fig.17.** Scatter plot of weights of spatial-dependent increment based on (a) coarse  
 634 resolution increment ( $\Delta C$ ) and fine resolution increment ( $\Delta F$ ); (b) comparison of  
 635 combined increments using weights derived from coarse and fine images.

### 636 5.3 Combination of temporal and spatial-dependent increments

637 Temporal increment and spatial-dependent increment are combined in IFSDAF  
 638 by CLS method in moving windows. Such a combination is based on the assumption  
 639 that the accuracies of two increment estimations are different under different scenarios;  
 640 thus the weighted combination is able to improve the accuracy of NDVI prediction  
 641 through balancing biases in the estimate of two increments. We verified the  
 642 assumption by comparing the temporal increment, the spatial-dependent increment  
 643 and the combined increment with the real increment (Fig.18), in which RMSE was  
 644 used to represent the error of estimation. Theoretically, a good combination is  
 645 expected to obtain smaller RMSE value than either of the two increment estimations.  
 646 As shown in Fig.18, the performance of CLS-based combination agrees with our  
 647 expectation with decreased RMSE values at both study sites, demonstrating the  
 648 necessity of combining two increments. Moreover, the residual of spatial-dependent

649 increment ( $\Delta S - \Delta F$ ) is much more similar to the residual of the combined increment  
 650 ( $\Delta F^{\text{Com}} - \Delta F$ ) than that of the temporal increment ( $\Delta T - \Delta F$ ), suggesting that the  
 651 spatial-dependent increment contributes more to the combined increment than the  
 652 temporal increment at these two sites.



653  
 654 **Fig. 18.** Difference between predicted increment and observed increment: difference  
 655 between (a) temporal increment  $\Delta T$ , (b) spatial-dependent increment  $\Delta S$ , (c) combined  
 656 increment  $\Delta F^{\text{Com}}$  and the observed increment  $\Delta F$  in the Coleambally irrigation area;  
 657 difference between (d) temporal increment  $\Delta T$ , (e) spatial-dependent increment  $\Delta S$ , (f)  
 658 combined increment  $\Delta F^{\text{Com}}$  and the observed increment  $\Delta F$  in the Gwydir area.  
 659

#### 660 **5.4 Improvements of IFSDAF compared with FSDAF**

661 Compared with FSDAF, IFSDAF has improved in the following aspects. First,  
662 the increment estimation in FSDAF mainly produced by the unmixing process, and  
663 the TPS interpolation result is only used to guide the distribution of residuals rather  
664 than producing spatial-dependent increment. However, as shown in Fig.18, the  
665 spatial-dependent increment estimated by the TPS interpolation may be more accurate  
666 than the temporal increment by the unmixing process. FSDAF underestimates the  
667 contribution of the TPS interpolation to some extent. The reason why the  
668 spatial-dependent increment is superior to the temporal increment can be found from  
669 Table 5, where we calculated the global Moran's I index of the coarse images for band  
670 Red, band NIR, and NDVI on the base date  $t_0$  and prediction date  $t_p$ , respectively. The  
671 global Moran's I index was used here to measure the spatial autocorrelation of the  
672 image, i.e., the relationship of pixel values between neighboring pixels. Larger  
673 Moran's I index indicates higher spatial autocorrelation. Table 5 shows that the spatial  
674 autocorrelation of NDVI represented by Moran's I index is greater than both the Red  
675 and NIR bands, because NDVI, as a feature-enhancing index, can enlarge the data  
676 variance compared with the Red and NIR bands (Fig.11a-b). As well known, greater  
677 spatial autocorrelation can yield more accurate result in spatial interpolation.  
678 Accordingly, the spatial-dependent increment estimated by the TPS interpolation for  
679 NDVI should be more accurate than that of Red and NIR bands. Therefore,  
680 spatial-dependent increment is more important for fusing NDVI than the raw bands,

681 which greatly benefits the NDVI fusion in IFSDAF. This study implies that without  
682 combined with temporal increment, spatial-dependent increment itself may obtain  
683 acceptable fusion results. This solution can greatly simplify the fusion process and  
684 reduce the computing cost. The simplified fusion model is more effective for the  
685 applications in larger areas and when the scale difference between coarse and fine  
686 images are not too large, which ensures adequate number of sample points (the center  
687 of coarse pixel) to obtain accurate prediction of the fine image by TPS interpolation.

688 **Table. 5.** Moran’s I of band Red, band NIR, and NDVI on the base date  $t_0$  and  
689 prediction date  $t_p$  in both the Coleambally irrigation area and Gwydir area at the  
690 coarse resolution.

Band	Base date $t_0$			Prediction date $t_p$		
	Red	NIR	NDVI	Red	NIR	NDVI
Coleambally irrigation area	0.5048	0.5225	0.6439	0.5677	0.4764	0.6840
Gwydir area	0.5867	0.7069	0.7881	0.6401	0.7568	0.8584

691

692 Second, IFSDAF uses a better way to combine two increments while FSDAF  
693 uses only one increment. As we know, the collinearity effect impacts the accuracy of  
694 unmixing for temporal increment estimation. Moreover, errors in the classification  
695 map and change of land cover also cause uncertainties of temporal increment. In order  
696 to correct the potential errors in the temporal increment, FSDAF introduces a  
697 homogeneity index  $HI(x_j, y_j)$  (over the range of 0-1, derived from the classification

698 map at  $t_0$ ) to help allocate residuals  $R(x, y)$  within the coarse pixel. However, when  
699 there are land cover changes and misclassification,  $HI(x_j, y_j)$  calculated from the  
700 classification map at  $t_0$  will not be suitable for allocating residuals on date  $t_p$ . Under  
701 this circumstance, the effectiveness of residuals distribution in FSDAF is restricted.  
702 Unlike FSDAF, the IFSDAF employs CLS method in moving windows and avoids  
703 the use of the homogeneity index; moreover, it allows the final increment estimation  
704 with local and adaptive capacity to better combine temporal and spatial-dependent  
705 increment.

706 The third improvement of IFSDAF is that it can employ fine NDVI images  
707 partially contaminated by clouds, in which clear pixels also provide valuable  
708 information. In IFSDAF, the clear pixels in those fine images are also used as base  
709 date to estimate the fine NDVI values at the prediction date respectively, and all  
710 predictions on date  $t_p$  are finally integrated by weights based on the temporal change  
711 magnitude in NDVI between base and prediction dates. This weighted prediction can  
712 reduce the critical dependency to the clear fine NDVI image and alleviate prediction  
713 uncertainties if date of the clear fine NDVI image is far from the prediction date. Of  
714 course, the better use of partially contaminated images needs accurate cloud labeling  
715 method (e.g., Fmask method). If there are mistakes in cloud labels, estimation results  
716 of IFSDAF will be impacted. For instance, land surface with high reflectance (e.g.,  
717 sand or snow) is possibly misidentified as clouds (Chen et al., 2016). Moreover,  
718 Fmask sometimes omits thin clouds, resulting in cloudy pixels being used in the

719 process of data fusion. Fortunately, effect of the errors in cloud mask can be  
720 minimized in IFSDAF because of the weighted combination of predictions from  
721 multiple dates. Moreover, with the advance of cloud screening methods (Zhu and  
722 Helmer, 2018), the issue can be greatly alleviated. Besides, like other existing  
723 spatiotemporal fusion methods, a pure clear image guarantees all pixels to be  
724 predicted by IFSDAF. However, in some areas such as tropical areas (e.g., Amazonia),  
725 it is difficult to obtain such clear fine image during a long period. Under this condition,  
726 using all partially contaminated fine images instead of a purely clear image is a  
727 practical choice in IFSDAF although it may leave some pixels not predicted in the  
728 fused images if these pixels do not have any one cloud-free observation in the time  
729 series.

### 730 **5.5 Applications to other remote sensing products**

731 Although IFSDAF is designed for the spatiotemporal fusion of NDVI time series,  
732 it can also be applied to fuse other vegetation indices like Enhanced Vegetation Index  
733 (EVI) and other products such as surface reflectance. To test the applicability of  
734 IFSDAF to other products, we assess the performance of IFSDAF in fusion of EVI,  
735 Red and NIR bands in Coleambally Irrigation area where have great heterogeneity.  
736 RMSEs of the fused images on Jan. 12<sup>th</sup>, 2002 (Table. 6) suggest that IFSDAF  
737 produces higher accuracy than FSDAF when fusing EVI, while when predicting  
738 surface reflectance (Red and NIR bands) the accuracy of IFSDAF and FSDAF does  
739 not differ too much. These results confirm that IFSDAF is more suitable than the

740 original FSDAF model for fusing remote sensing products with high spatial  
 741 autocorrelation.

742 **Table 6.** RMSE (rRMSE) of predictions by IFSDAF and FSDAF on EVI, Red and  
 743 NIR bands.

index/surface reflectance	EVI	Red	NIR
FSDAF	0.0755 (29.11%)	0.0271 (19.03%)	0.0341 (11.02%)
IFSDAF	0.0650 (25.09%)	0.0245 (17.21%)	0.0337 (10.91%)

## 744 6. Conclusions

745 In this study, we proposed an improved FSDAF method specifically for  
 746 producing NDVI time series with a high spatiotemporal resolution. Coarse NDVI  
 747 (MODIS) and fine NDVI images (Landsat and Sentinel) were used to test the  
 748 performance of the new method for different sensors. Experiments show that the  
 749 fused NDVI images by IFSDAF is more accurate than FSDAF as well as other two  
 750 existing methods (NDVI-LMGM and STARFM) in areas with a great degree of  
 751 spatial heterogeneity and with significant land cover changes. The better performance  
 752 of IFSDAF can be attributed to producing spatial-dependent increment by the TPS  
 753 interpolation, employing CLS method in moving windows to adaptively combine the  
 754 temporal increment and the spatial-dependent increment, as well as the better use of  
 755 partially contaminated fine images. Such significant improvements are made in  
 756 accordance with the characteristics of NDVI with larger data variance and spatial  
 757 autocorrelation compared with raw reflectance bands. Considering the significant

758 contribution of spatial-dependent increment by the TPS interpolation, when the scale  
759 difference between coarse and fine images is not very large, the proposed IFSDAF  
760 method can be further simplified by only using spatial-dependent increment to  
761 improve the efficiency. This result of the study also supports the IFSDAF to be a  
762 feasible method for applications in a large area and different sensors. Moreover, it is  
763 also applicable to other vegetation index data. We call for more testing of the new  
764 method by using other satellite data (e.g. Sentinel and VIIRS data) and in other areas.

#### 765 **Acknowledgement**

766 This study was supported by National Key Research and Development Program of  
767 China (No. 2017YFD0300201), the Research Grants Council of Hong Kong (project  
768 no.25222717), and the National Natural Science Foundation of China (project  
769 no.41701378), and CERESEA Oversea Joint Research Program, Chiba University  
770 (No.CI17-103).

771

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886 **Appendix**

887 Useful notations.

$t_0$	Base date	$f_l(x, y)$	Fraction of class $l$ within coarse pixel $(x, y)$
$t_p$	Prediction date	$\Delta F_c$	Fine spatial resolution increment of class $c$ within the moving window
$(x, y)$	Location of coarse spatial resolution pixel $(x, y)$	$F_0^{\text{TPS}}$	Result of TPS interpolation based on coarse NDVI on $t_0$
$(x_j, y_j)$	Location of $j$ th fine spatial resolution pixel within coarse pixel $(x, y)$	$F_p^{\text{TPS}}$	Result of TPS interpolation based on coarse NDVI on $t_p$
$F_0$	Fine spatial resolution NDVI on $t_0$	$w_s$	Weight of spatial-dependent increment
$F_p$	Fine spatial resolution NDVI on $t_p$	$w_T$	Weight of temporal increment
$C_0$	Coarse spatial resolution NDVI on $t_0$	$\hat{F}_p$	Fine spatial resolution prediction on date $t_p$
$C_p$	Coarse spatial resolution NDVI on $t_p$	$\hat{F}_{0,p}$	Fine spatial resolution prediction on date $t_p$ based on fine NDVI on date $t_0$
$\Delta F$	Fine spatial resolution NDVI increment	$\hat{F}_{p+1,p}$	Fine spatial resolution prediction on date $t_p$ based on fine NDVI on date $p+1$

$\Delta C$	Coarse spatial resolution NDVI increment	$\Delta F^{Com}$	Combined fine spatial resolution increment based on $\Delta T$ and $\Delta S$
$\Delta T$	Fine spatial resolution temporal increment	$R(x, y)$	Residual within the coarse pixel ( $x, y$ )
$\Delta S$	Fine spatial resolution spatial-dependent increment	$C_q^i(x, y)$	$i$ th coarse pixel in the moving window centered by coarse pixel ( $x, y$ ) on date $q$
$\Delta C^T$	Upscaled fine spatial resolution temporal increment	$C_p^i(x, y)$	$i$ th coarse pixel in the moving window centered by coarse pixel ( $x, y$ ) on date $t_p$
$\Delta C^S$	Upscaled fine spatial resolution spatial-dependent increment	$w_{q,p}(x, y)$	Contribution coefficient of fine spatial resolution pixels on date $q$ to the final prediction on $t_p$ within coarse pixel ( $x, y$ )