

1 Sensitivity of six typical spatiotemporal fusion methods to differ-
2 ent influential factors: a comparative study for a normalized dif-
3 ference vegetation index time series reconstruction

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24 **Abstract**

25 Dozens of spatiotemporal fusion methods have been developed to reconstruct veg-
26 etation index time-series data with both high spatial resolution and frequent coverage
27 for monitoring land surface dynamics. Although several comparison studies among dif-
28 ferent fusion methods have been conducted, selecting suitable fusion methods is still
29 challenging as inevitable influential factors tend to be neglected. To address this prob-
30 lem, this study compared six typical spatiotemporal fusion methods, including the Un-
31 mixing-Based Data Fusion (UBDF), Linear Mixing Growth Model (LMGM), Spatial
32 and Temporal Adaptive Reflectance Fusion Model (STARFM), Fit-FC (regression
33 model Fitting, spatial Filtering and residual Compensation), One Pair Dictionary-
34 Learning method (OPDL), and Flexible Spatiotemporal DAta Fusion (FSDAF), based
35 on simulation experiments and theoretical analysis with the consideration of three in-
36 fluent factors between sensors, including geometric misregistration, radiometric in-
37 consistency, and spatial resolution ratio. The results indicated that Fit-FC achieved the
38 best performance with the strongest tolerance to geometric misregistration when radio-
39 metric inconsistency was negligible; thus, it is the first recommended algorithm for
40 blending normalized difference vegetation index (NDVI) imagery. FSDAF could gen-
41 erate satisfactory results with resistance to radiometric inconsistency as well. These
42 findings could help users to determine which method is appropriate for different remote

43 sensing datasets and provide guidelines for developers in the future development of

44 novel methods.

45 **Keywords:** Spatiotemporal fusion; Normalized difference vegetation index (NDVI);

46 Geometric misregistration; Radiometric inconsistency; Spatial resolution ratio

47

48 **1. Introduction**

49 Time series of vegetation indices (e.g., Normalized Difference Vegetation Index,
50 NDVI) produced by satellite sensors play a unique role in various environmental appli-
51 cations as important data sources, such as cropland mapping (Chang et al., 2007; Ward-
52 low et al., 2007), vegetation phenology monitoring (Bradley et al., 2007; Cao et al.,
53 2015; Zhang et al., 2003), and disturbance detection (Verbesselt et al., 2012). However,
54 most of the sensors onboard the launched satellites cannot acquire data with both high
55 spatial and temporal resolutions simultaneously, due to hardware technology or budget
56 limitations. For example, the data from sensors with dense temporal coverage usually
57 hold coarse spatial resolution (e.g. MODIS, hereafter referred to as coarse images), im-
58 posing restrictions on capturing enough spatial details in heterogeneous areas. On the
59 other hand, the data from sensors with fine spatial resolution (e.g., Landsat TM or
60 ETM+, hereafter referred to as fine images) have their drawback due to a long revisit
61 cycle (e.g., 16 days), limiting their potential in time-series analyses. Consequently, var-
62 ious spatiotemporal fusion methods that combine the merits of two such kinds of data
63 have been developed and were used to produce NDVI data with high spatial and tem-
64 poral resolutions (Chen et al., 2018; Liao et al., 2017; Liu et al., 2019; Maselli et al.,
65 2019; Rao et al., 2015). Furthermore, they have been successfully applied in various
66 fields, such as crop growth progress monitoring (Gao et al., 2017), land cover classifi-
67 cation (Chen et al., 2017; Jia et al., 2014), biomass estimation (Zhang et al., 2016), and
68 disturbance detection (Hilker et al., 2009).

69 Zhu et al. (2018) grouped the published spatiotemporal fusion methods into the

70 following five categories according to technique principles: unmixing-based, weight
71 function-based, learning-based, Bayesian-based, and hybrid methods. Unmixing-based
72 methods downscale coarse pixel to fine resolution based on the linear spectral mixing
73 theory (Rao et al., 2015; Zhukov et al., 1999; Zurita-Milla et al., 2008). Weight func-
74 tion-based methods estimate target pixel through combining neighborhood pixels with
75 empirically designed weight functions of spectral similarity, spatial distance or other
76 related measurements (Gao et al., 2006; Wang et al., 2018; Zhu et al., 2010). Learning-
77 based methods are relatively new, which use machine learning methods to model the
78 relationship between coarse and fine images (Huang and Song, 2012; Liu et al., 2016;
79 Song and Huang, 2013; Song et al., 2018). Bayesian-based methods described spatio-
80 temporal fusions as a Maximum A Posterior (MAP) problem based on Bayesian frame-
81 work (Huang et al., 2013; Liao et al., 2016; Shen et al., 2016). Hybrid methods attempt
82 to integrate two or more methods mentioned above to improve the performance of spa-
83 tiotemporal fusion (Li et al., 2020; Liu et al., 2019; Quan et al., 2018; Zhu et al., 2016).

84 Although the technique principles are diverse, each developed method was
85 claimed by its original study to have unique advantages in terms of prediction accuracy,
86 computation efficiency, or input data requirements. However, as these studies used dif-
87 ferent datasets in their method comparison, it was difficult to reach a consensus on
88 which method outperforms all the others. Thus, it is necessary to assess the applicability
89 of these methods to different application scenarios. Accordingly, several cross-compar-
90 ison studies had been conducted to explore the advantages and weaknesses of different
91 methods based on time-series data (Chen et al. 2015; Emelyanova et al., 2013; Liu et

92 al., 2019). As a general conclusion, the performances of different fusion methods
93 mainly depend on the sensitivity to spatial heterogeneity and temporal variations of the
94 used data.

95 However, these comparison studies have neglected the influence of inevitable
96 noise in real applications, including geometric misregistration and radiometric incon-
97 sistency. In spite of large efforts devoted on the inter-calibration and geometric regis-
98 tration among different sensors, adequate elimination on such inherent noises is still
99 challenging (Chander et al. 2013a, 2013b; Claverie et al., 2018; Yan et al., 2016). Thus,
100 numerous studies have focused on quantifying the impact of geometric misregistration
101 error and radiometric uncertainty on land cover change detection (Dai and Khorram,
102 1998; Chen et al., 2014; Roy et al., 2000) and vegetation dynamic monitoring (Fan and
103 Liu, 2018; Skakun et al., 2018; Sulla-Menashe et al., 2016). Considering the potential
104 impacts of these noises on spatiotemporal fusion methods (Belgiu and Stein, 2019; Zhu
105 et al., 2018) and the lack of corresponding comparative research, it is still difficult for
106 users to choose appropriate methods for their applications. A recent study has shown an
107 encouraging desire to address these issues by quantifying the influence of geometric
108 errors on the fusion methods (Tang et al., 2020). However, only two algorithms were
109 explored in this study, which is not enough for most users.

110 Recently, open source data from sensors with fine spatial resolution have made a
111 progress in enhancing temporal frequency, such as Sentinel-2, but they are still not suf-
112 ficient in many applications due to cloud contamination (Wang et al., 2018). Thus, Sen-

113 tinel-3 at 300 m resolution is also launched to supply daily observations, which is cru-
114 cial for monitoring land surface dynamics. Moreover, long-term data analysis is also
115 very important for many applications and spatiotemporal fusion plays important role
116 for the study period without rich data. Therefore, spatiotemporal fusion methods are not
117 only used for MODIS and Landsat images, but also images from other satellite sensors
118 with different spatial resolutions (e.g., AMSR, ASTER, Sentinel-2/3, GF-1, Worldview,
119 and Planet) in recent studies (Kong et al., 2016; Kwan et al., 2018; Mizuochi et al.,
120 2017; Li et al., 2017; Wang and Atkinson, 2018). And it has been aware that the input
121 images with different resolution ratios could lead to significant variations of different
122 method performances (Yokoya et al., 2017). Unfortunately, to our knowledge, there are
123 no comparative studies for evaluating the performances of spatiotemporal fusion meth-
124 ods based on data with different spatial resolution ratios of sensors.

125 To fill the gap in the previous comparison studies, we conducted comparison ex-
126 periments and theoretical analyses on the spatiotemporal fusion of NDVI time-series
127 data with considering various influential factors, including geometric misregistration,
128 radiometric inconsistency, and spatial resolution ratio. Six typical spatiotemporal fusion
129 methods requiring only one fine image and two coarse images as input, including the
130 UBDF (Unmixing-Based Data Fusion), LMGM (Linear Mixing Growth Model),
131 STARFM (Spatial and Temporal Adaptive Reflectance Fusion Model), Fit-FC (regres-
132 sion model Fitting, spatial Filtering and residual Compensation), OPDL (One Pair Dic-
133 tionary-Learning method), and FSDAF (Flexible Spatiotemporal DAta Fusion), were
134 selected for comparison. The six methods were selected in this study considering their

135 unique contributions in their own categories and the availability of source codes. More-
136 over, the performances were evaluated on time-series data instead of individual images
137 to better satisfy the application requirement. In general, the goal of this study is to ex-
138 plore the sensitivity of the six fusion methods to three influential factors and, thus, pro-
139 vide useful guidelines for method selection and future method design to users and de-
140 velopers.

141 **2. Methods and datasets**

142 **2.1 Experiment design**

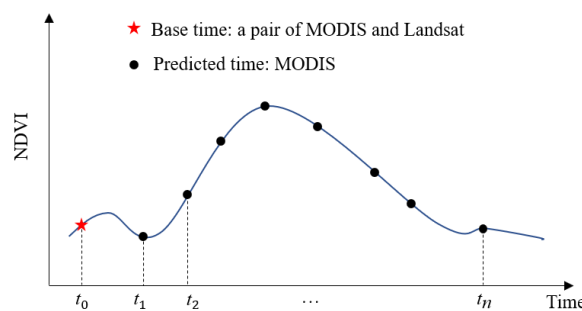
143 To explore the sensitivity of spatiotemporal fusion methods to various influencing
144 factors for NDVI time series reconstruction, experiments were specifically designed in
145 terms of geometric misregistration, radiometric inconsistency, and different spatial res-
146 olution ratios. Like in previous studies (Gevaert and García-Haro, 2015; Liu et al., 2019;
147 Zhu et al., 2016), the time series of cloud-free Landsat imagery and simulated coarse
148 resolution imagery aggregated from Landsat data were used for a spatiotemporal fusion
149 experiment and validation. The standard experiment is based on ideal simulated data
150 without any errors; it is used as the reference for the later simulation experiments. To
151 explore the effect of various influencing factors, three additional fusion experiments
152 were designed based on the simulated data with geometric misregistration, radiometric
153 inconsistency, and different spatial resolution ratios. In addition, fusion experiments
154 based on actual Landsat and MODIS data were also conducted.

155 NDVI fusion is the main object of this study considering the widely application of

156 NDVI time series. As the time series of surface reflectance also receive attentions (Her-
157 mosilla et al., 2015; Xiao et al., 2016), similar fusion experiments were also conducted
158 on reflectance data (green, red, near infrared bands) for a comparison.

159 **2.1.1 Standard fusion experiment based on ideal simulated data**

160 This experiment followed the experimental settings of previous studies (Gevaert
161 and García-Haro, 2015; Zhu et al., 2016). Coarse images were simulated by the aggre-
162 gation of Landsat images to avoid misregistration and radiometric inconsistency be-
163 tween fine and coarse images. In the standard experiment, we aggregate 8×8 pixels for
164 NDVI and 16×16 pixels for reflectance to simulate the MODIS NDVI at 250m resolu-
165 tion and MODIS reflectance data at 500m resolution. The first fine image in the time
166 series and corresponding simulated coarse image were used as the base-paired image
167 input for the fusion experiment. The other simulated coarse images were then
168 downsampled to a fine spatial resolution by different fusion methods (Fig. 1. Schematic
169 diagram of the standard experiment.).



170

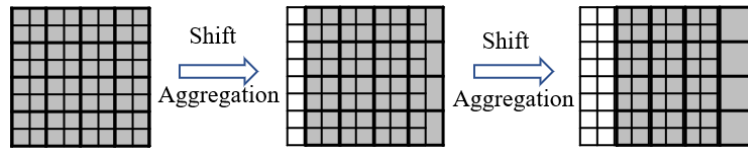
171

Fig. 1. Schematic diagram of the standard experiment.

172 **2.1.2 Fusion experiment based on simulated data with geometric misregistration**

173 The settings of this experiment are similar to the standard one, except that misreg-
174 istration error was simulated when aggregating the Landsat images to the coarse images.

175 Specifically, pixel shifting is one of the most serious consequences caused by geometric
 176 distortions. Therefore, similar to the previous study (Tang et al., 2020), the fine images
 177 were shifted 2, 4, 6, or 8 pixels before aggregation, thus different degrees of misregis-
 178 tration error were generated for the simulated coarse images (Fig. 2). This experiment
 179 compares the robustness of different methods to the geometric error. For experiments
 180 of reflectance, the fine images were shifted 4, 8, 12, or 16 pixels before aggregation
 181 considering the resolution of coarse reflectance images was doubled as that of NDVI
 182 image.



183

184 **Fig. 2.** Schematic diagram of the experiment with geometric misregistration.

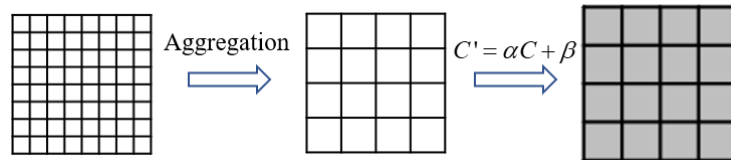
185 2.1.3 Fusion experiment based on simulated data with radiometric inconsistency

186 The special experiment setting of this experiment, which is the only difference
 187 from the standard one, is that a linear stretch was conducted on the aggregated coarse
 188 image to simulate the radiometric inconsistency between fine and coarse sensors (Fig.
 189 3):

$$C' = \alpha C + \beta, \quad (1)$$

190 where C and C' are the ideal and the stretched NDVI of simulated coarse pixels, re-
 191 spectively; α and β are the linear stretch parameters. The parameters were referenced
 192 from an intercalibration study of vegetation indices derived from different sensors (Ste-
 193 ven et al., 2003), in which the linear relationships of TM and MODIS, TM and AVHRR,

194 ETM+ and MODIS, POLDER and ASTR2, as well as QuickBird and ASTR2 were in-
 195 vestigated (Table 1). And Table 2 presents the linear relationships used in the reflec-
 196 tance experiments. With such a simulation, the sensitivity of different methods to the
 197 radiometric inconsistency could be explored.



198

199 **Fig. 3.** Schematic diagram of the experiment with radiometric inconsistency

200 **Table 1**

201 Coefficients of linear stretches for simulated radiometric inconsistency in the NDVI
 202 experiments between sensors (Steven et al., 2003).

Satellite Sensors	Slope (α)	Intercept (β)
TM-MODIS	1.002	-0.012
TM-AVHRR	1.106	-0.007
ETM+-MODIS	1.023	-0.013
POLDER-ASTR2	1.008	-0.110
QuickBird-ASTR2	0.928	-0.105

203 **Table 2**

204 Coefficients of linear stretches for simulated radiometric inconsistency in the reflec-
 205 tance experiments between sensors.

Simulations	Slope (α)	Intercept (β)
Simu1	0.9	0.0
Simu2	1.1	0.0
Simu3	1.0	-0.05
Simu4	1.0	0.05

206 **2.1.4 Fusion experiment based on simulated data with different spatial resolution**
 207 **ratios**

208 To explore the applicability of six fusion algorithms to various satellite products

209 with different spatial resolutions, this experiment compares the sensitivity of these
210 methods to different spatial resolution ratios of coarse and fine images. Coarse images
211 are simulated at 4 levels of spatial resolution ratios (4, 8, 16, and 32; Fig. 4). Other
212 experiment settings are similar to those in the standard one (i.e., without any geometric
213 error or radiometric inconsistency).



214

215 **Fig. 4.** Schematic diagram of the experiment with different spatial resolution ratios.

216 **2.1.5 Fusion experiment based on actual MODIS images**

217 The actual MODIS images were used for this fusion experiment. In addition, sim-
218 ulated MODIS images with certain geometric error and radiometric inconsistency were
219 also used for comparison. Although the geolocation accuracy of MODIS achieved 50
220 m at nadir (Wolfe et al., 2002), the large scan angle and procedures of reprojection and
221 resampling could further enlarge the geolocation error. Thus, it should be reasonable to
222 assume an averaged misregistration error of approximately a half-pixel size (120 m),
223 corresponding to four pixel shifting of fine images. The parameters of linear stretches
224 for TM-MODIS (Table 1) were used for simulating the radiometric inconsistency. This
225 experiment was conducted to illustrate how much of the fusion error of the results using
226 the actual MODIS data could be accounted for by the results based on the simulated
227 MODIS data with simulated geometric and radiometric errors.

228 **2.1.6 Accuracy indices for evaluation**

229 Two accuracy indices, root mean square error (RMSE) and, correlation coefficient

230 (r) were used to evaluate the performance of different fusion methods. The RMSE was
231 calculated using all pairs of the predicted and true images throughout the time series.
232 The correlation coefficient (r) was calculated between the predicted and the true NDVI
233 time-series for each fine pixel. Then, an averaged r of the whole image was used to
234 represent the overall accuracy of the predicted time-series data. Different aspects of
235 fusion results were assessed. The image-based RMSE evaluates the average pixel-wise
236 prediction errors, which has drawn the attention of quantitative remote sensing studies.
237 The coefficient r is the similarity between the predicted temporal profile and true tem-
238 poral profile, which will benefit dynamic monitoring research. With the above two in-
239 dices, the overall performances on the time-series instead of the individual images were
240 evaluated for different fusion methods under different experimental scenarios.

241 In addition, to further explore the relationship between the fusion accuracy and the
242 temporal variation of the input data, an absolute relative difference index (ARDI), was
243 calculated to represent temporal change between base and predicted time.

$$244 \quad \text{ARDI} = |F_2 - F_1|/F_1, \quad (2)$$

245 where F_1 and F_2 denote the NDVI or reflectance of fine images at based and pre-
246 dicted time.

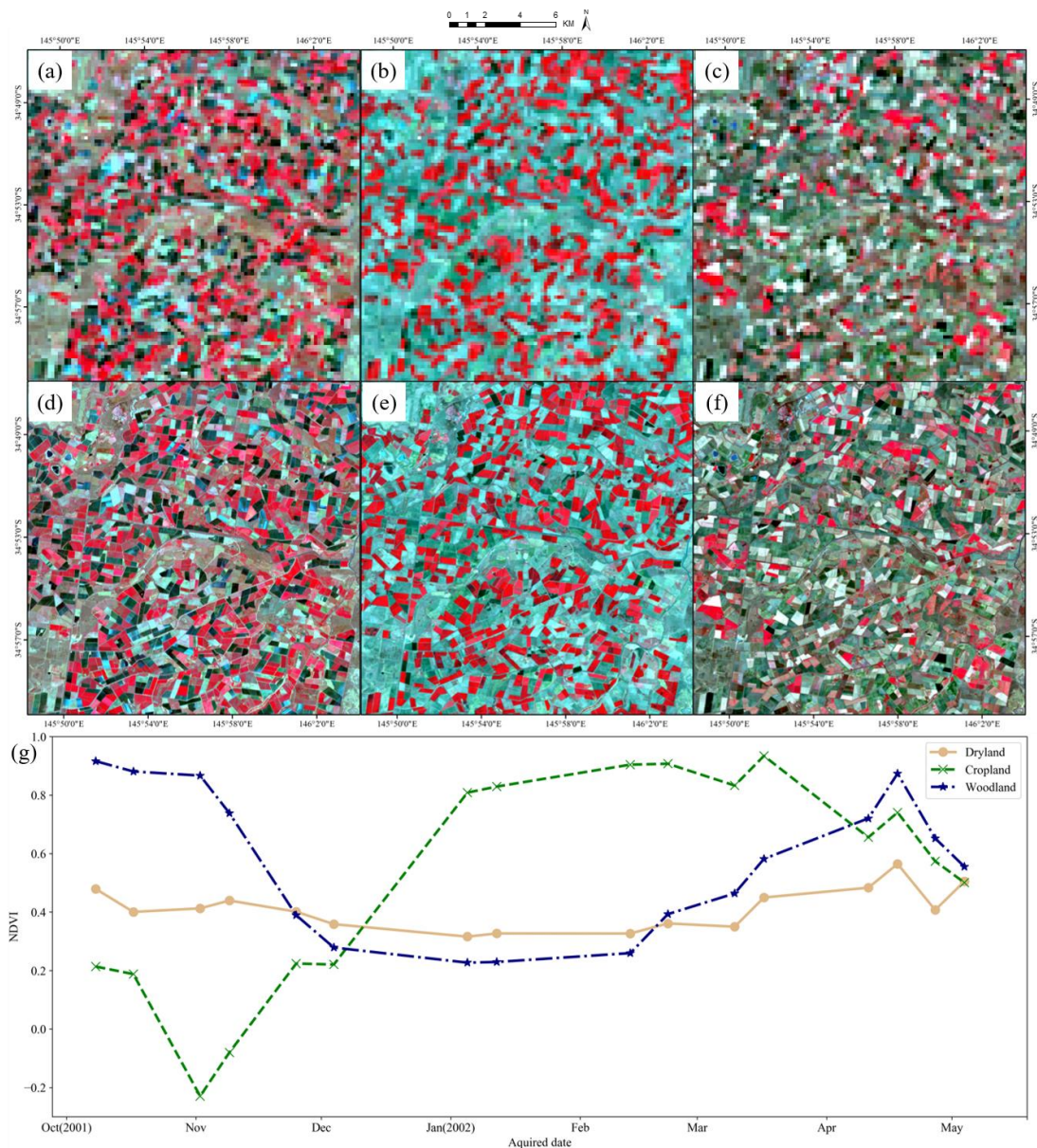
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248 **2.2 Experimental Datasets**

249 For a unified comparison, the typical datasets in previous spatiotemporal fusion
250 studies (Emelyanova et al., 2013; Jun et al., 2020), Coleambally irrigated area (CIA) in

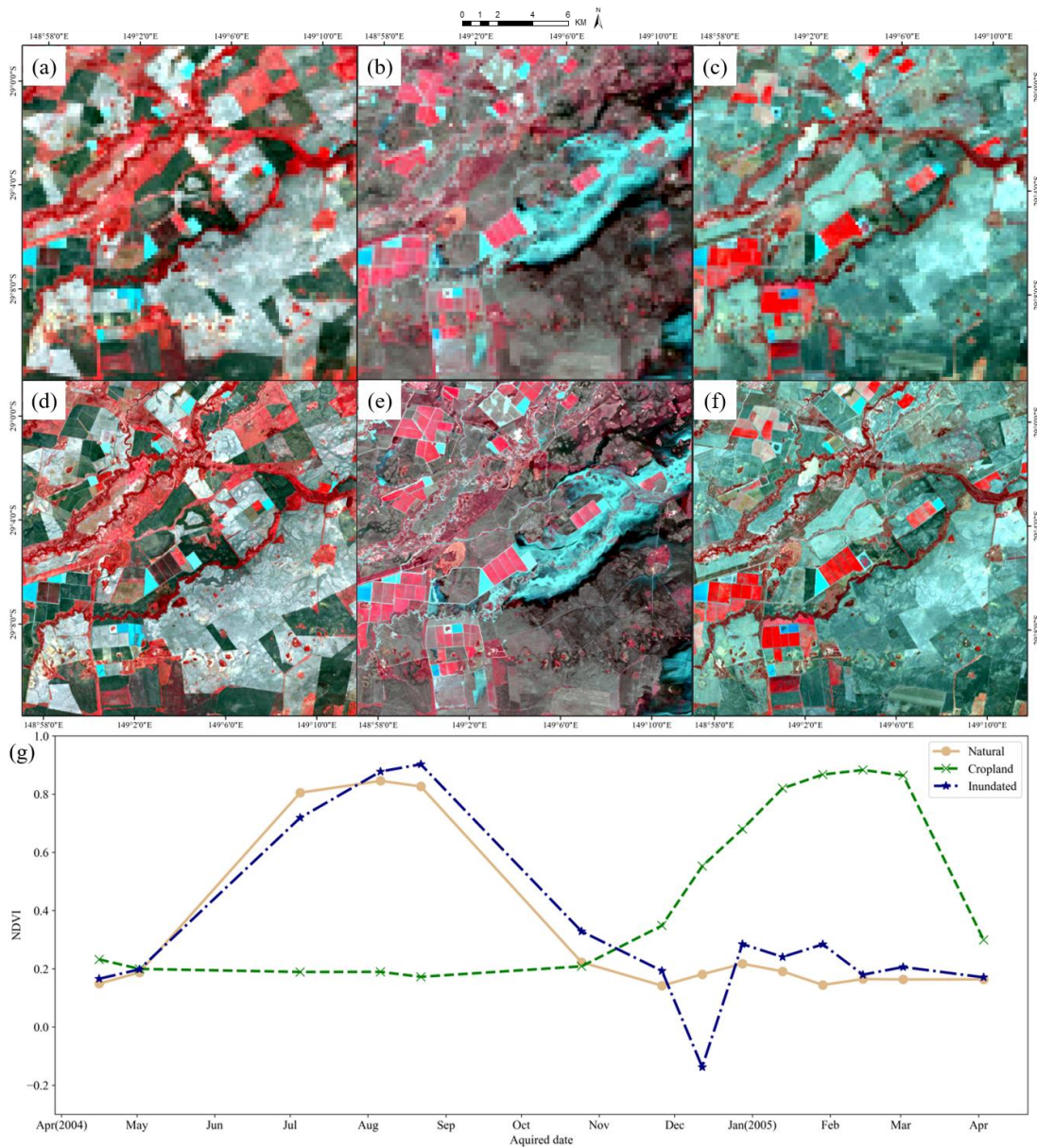
251 southern New South Wales (145.10°E, 34.05°S), Gwydir Catchment (GWY) in north-
252 ern New South Wales (149.63°E, 29.77°S) and Tianjin in northern China (117.20°E,
253 39.30°N) were used in this study. The CIA site was dominated by woodlands, cropland,
254 and dryland land cover types. A total of 16 cloud-free pairs of Landsat-7 ETM+
255 (800×800 pixels at 30 m spatial resolution) data were collected in this area from Octo-
256 ber 2001 to May 2002. As shown in Fig. 5(a) and (b), there are fragmented cropland
257 and woodlands parcels in this area, resulting in a heterogeneous landscape. In addition,
258 woodlands, croplands, and drylands show distinctive NDVI profiles during this period
259 (Fig. 5(c)). The main purpose using this dataset with the high heterogeneity and com-
260 plex NDVI seasonality is to compare the performance of fusion methods for the moni-
261 toring phenology changes in fragmented cropland landscape. The GWY site was dom-
262 inated by winter crops and natural vegetation. A total of 14 cloud-free pairs of Landsat-
263 5 TM (800×800 pixels at 30 m spatial resolution) data were collected in this area from
264 April 2004 to April 2005. This site was relatively homogeneous, displaying relatively
265 large parcels of crop fields and natural vegetation (Fig. 6(a) and (b)). However, a flood
266 occurred in December 2004, leading to a sudden drop in the NDVI of the inundated
267 areas (Fig. 6(c)). Thus, this dataset is employed to test the performance of fusion meth-
268 ods for capturing abrupt land cover change. As for the third site, Tianjin, the main land
269 cover were impervious surface, cropland and waterbody. Many small impervious sur-
270 faces (e.g., buildings, and roads) were distributed in this site, resulting in a heterogenous
271 landscape. There were 11 cloud-free pairs of Landsat-8 OLI (800×800 pixels at 30 m
272 spatial resolution) collected for Tianjin Site. As shown in Fig. Fig. 7, each land cover

273 had the unique NDVI temporal profile. The main purpose using this dataset is to test
 274 the accuracy of fusion methods for the detecting phenology changes in urban land-
 275 scapes. For all three sites, true MODIS surface reflectance (MODIS Terra MOD09GQ
 276 collection 6, resampled to 240 m spatial resolution) acquired in the corresponding pe-
 277 riods were also downloaded for comparison.

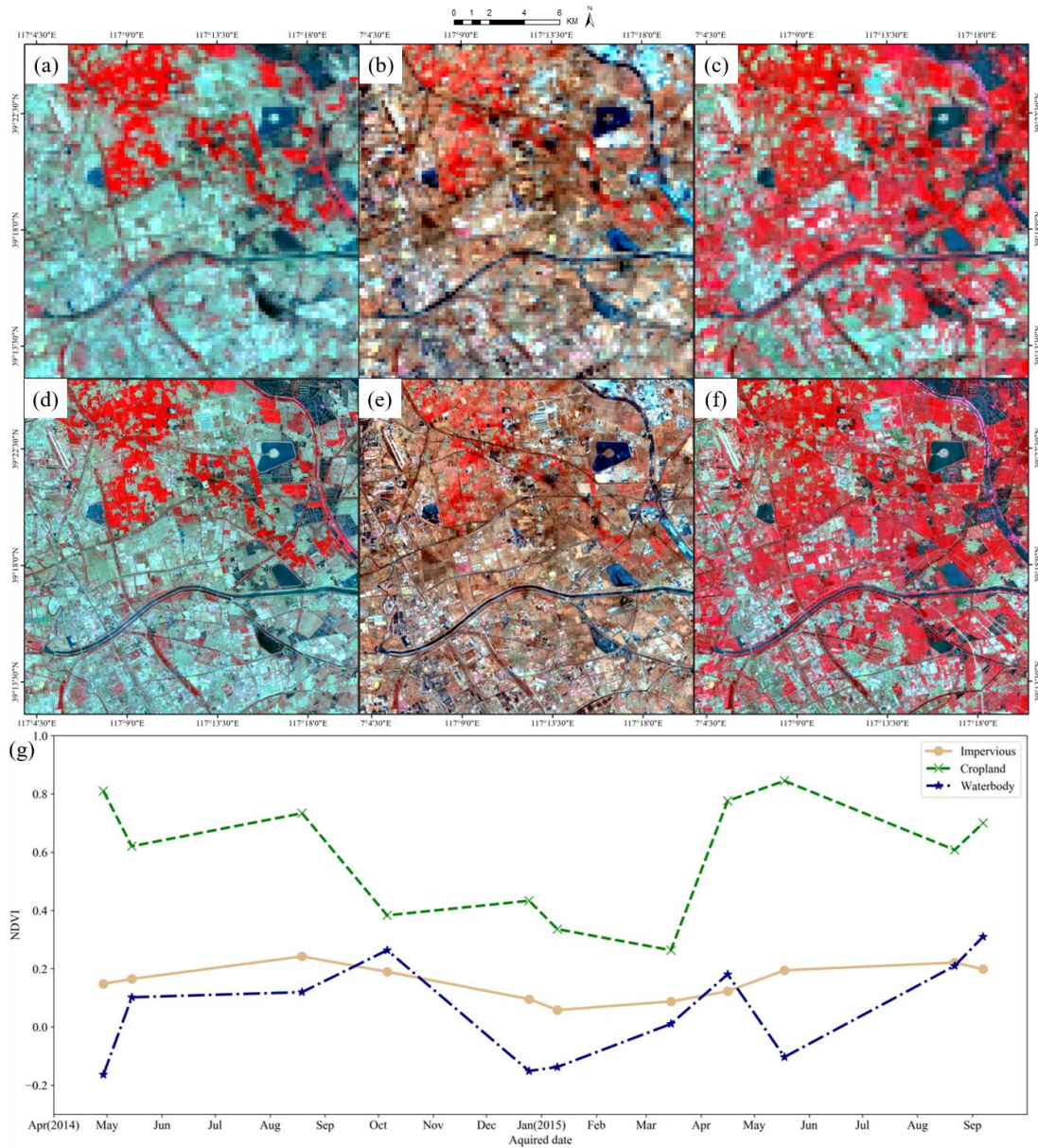


278
 279 **Fig. 5.** Test data in CIA site. 240 m simulated coarse images and corresponding 30 m
 280 fine images acquired on (a) and (d) November 9, 2001, (b) and (e) February 13, 2002,

281 and (c) and (f) May 4, 2002; (g) NDVI time-series of three typical land covers. All
 282 images use NIR-red-green as RGB.



283
 284 **Fig. 6.** Test data in GWY site. 240 m simulated coarse images and corresponding 30 m
 285 fine images acquired on (a) and (d) May 2, 2004, (b) and (e) December 12, 2004, and
 286 (c) and (f) April 4, 2005; (g) NDVI time-series of three typical land covers. All images
 287 use NIR-red-green as RGB.



288

289 **Fig. 7.** Test data in Tianjin site. 240 m simulated coarse images and corresponding 30
 290 m fine images acquired on (a) and (d) April 29, 2014, (b) and (e) December 25, 2014,
 291 and (c) and (f) August 2, 2015; (g) NDVI time-series of three typical land covers. All
 292 images use NIR-red-green as RGB.

293 2.3 Spatiotemporal fusion methods

294 We selected 1~2 typical methods for each category mentioned by Zhu et al. (2018)
 295 for the comparison experiments except for the Bayesian-based methods due to the lack

296 of open-source code. For quantifying the error propagation of fusion results caused by
 297 misregistration and radiometric consistency between sensors, key concepts and equa-
 298 tions of each method were introduced here for the convenience of the later theoretical
 299 analysis in the discussion part.

300 For simplification, the algorithms were reintroduced here based on a consistent
 301 denotation (Table 3).

302 **Table 3** Common variables used in different spatiotemporal fusion methods.

Symbol	Meaning
(x, y)	geolocation of specific pixel
t_1	base time;
t_2	predicted time;
C_1	the input coarse image at t_1 ;
C_2	the input coarse image at t_2 ;
F_1	the input fine image at t_1 ;
\hat{F}_2	the output image at t_2
M	the moving window of pixel (x, y) .

303

304 2.3.1 UBDF

305 As an unmixing-based method, UBDF employ a constrained least square with a
 306 moving to unmix coarse images for appropriate results (Zurita-Milla et al., 2008). Based
 307 on the linear spectral mixing model, NDVI at a coarse pixel is regarded as a linear
 308 combination of the NDVIs of its endmembers. Assuming that the fine pixels are pure
 309 enough to be endmembers, the NDVI at the coarse pixels (x, y) and a predicted time is:

$$C_2(x, y) = \sum_{i=1}^c f_i(x, y) F_2^i(x, y) + \varepsilon(x, y), \quad (2)$$

310 where $F_2^i(x, y)$ is the NDVI of the i th land cover type in the coarse pixel (x, y) ; $f_i(x, y)$ is
 311 the fraction of the i th endmember in the coarse pixel; c is the number of endmembers;
 312 and ε is the residual error. $f_i(x, y)$ is calculated based on the classification result of

313 fine images at t_1 , as land cover is assumed to be unchanged in UBDF. With another
 314 assumption that endmembers are consistent in a moving window of coarse pixels (\mathbf{M}),
 315 $F_2^i(x,y)$ can be solved by the following equations with constrained corresponding to the
 316 mixing models in a moving window ($m \times m$ coarse pixels):

$$\begin{bmatrix} C_2(1,1) \\ \mathbf{M} \\ C_2(x,y) \\ \mathbf{M} \\ C_2(m,m) \end{bmatrix} = \begin{bmatrix} f_1(1,1) & f_2(1,1) & \text{L} & f_c(1,1) \\ \mathbf{M} & \mathbf{M} & & \mathbf{M} \\ f_1(x,y) & f_2(x,y) & \text{L} & f_c(x,y) \\ \mathbf{M} & \mathbf{M} & & \mathbf{M} \\ f_1(m,m) & f_2(m,m) & \text{L} & f_c(m,m) \end{bmatrix} \begin{bmatrix} F_2^1 \\ F_2^2 \\ \mathbf{M} \\ F_2^c \end{bmatrix}. \quad (3)$$

317 Also, Eq. (3) could be written in a matrix form for convenience

$$\mathbf{C}_2(\mathbf{M}) = \mathbf{f}(\mathbf{M})\mathbf{F}_2(\mathbf{M}). \quad (4)$$

318 Thus, F_2^i could be estimated by the least-square method:

$$\hat{\mathbf{F}}_2(\mathbf{M}) = [\mathbf{f}^T(\mathbf{M})\mathbf{f}(\mathbf{M})]^{-1} \mathbf{f}^T(\mathbf{M})\mathbf{C}_2(\mathbf{M}). \quad (5)$$

319 Finally, the fine image at t_2 can be generated by assigning the estimated F_2^i to the
 320 corresponding fine pixels based on the classification result of t_1 .

321 2.3.2 LMGM

322 To further enhance spatial details in the unmixing-based fusion results, LMGM
 323 makes use of F_1 (Rao et al., 2015). It assumes that the growth rate of the same land
 324 cover is constant in a short period. Therefore, LMGM estimates the growth rates of
 325 endmembers ($\Delta F = F_2 - F_1$) by unmixing the growth rate of coarse pixels ($\Delta C = C_2 - C_1$), as
 326 shown in Eq. (6):

$$\begin{bmatrix} \Delta C(1,1) \\ \mathbf{M} \\ \Delta C(x,y) \\ \mathbf{M} \\ \Delta C(m,m) \end{bmatrix} = \begin{bmatrix} f_1(1,1) & f_2(1,1) & \mathbf{L} & f_c(1,1) \\ \mathbf{M} & \mathbf{M} & & \mathbf{M} \\ f_1(x,y) & f_2(x,y) & \mathbf{L} & f_c(x,y) \\ \mathbf{M} & \mathbf{M} & & \mathbf{M} \\ f_1(m,m) & f_2(m,m) & \mathbf{L} & f_c(m,m) \end{bmatrix} \begin{bmatrix} \Delta F^1 \\ \Delta F^2 \\ \mathbf{M} \\ \Delta F^c \end{bmatrix}. \quad (6)$$

327 Then, LMGM calculates \hat{F}_2 by adding the estimated growth rate of class i ($\Delta\hat{F}^i$) to
 328 F_1

$$\hat{F}_2(x,y) = F_1(x,y) + \Delta\hat{F}^i. \quad (7)$$

329 2.3.3 STARFM

330 STARFM is the most typical and popular fusion method based on a weight func-
 331 tion (Gao et al., 2006). It assumes that the systematic bias between two sensors does
 332 not change over time. STARFM firstly resamples the coarse images to the same spatial
 333 resolution as the fine image. Thus, $F_2(x,y)$ can be estimated as:

$$\hat{F}_2(x,y) = F_1(x,y) + \Delta C(x,y). \quad (8)$$

334 Considering the issues of mixed pixel and land cover change, the information of similar
 335 neighboring pixels is introduced for the final estimation of F_2 :

$$\hat{F}_2(x,y) = \sum_{i=1}^{n_s} W_i (F_1(x_i, y_i) + \Delta C(x_i, y_i)), \quad (9)$$

336 where n_s is the number of similar pixels in the moving window and W_i is the weight
 337 of the i th similar pixel. The definition of spectral neighbor similar pixels is that they
 338 belong to the same class. And the calculation of the weight W_i combines the spatial
 339 distance (D_i) and spectral difference between coarse and fine images (S_i) (Gao et al.,
 340 2006; Gao et al., 2015):

$$D_i = \sqrt{(x_{w/2} - x_i)^2 + (y_{w/2} - y_i)^2}, \quad (10)$$

$$S_i = |F_1(x_i, y_i) - C_1(x_i, y_i)|, \quad (11)$$

341 where $(x_{w/2}, y_{w/2})$ and (x_i, y_i) are the central pixel of the moving window and candidate
 342 similar neighboring pixel, respectively. The spatial closer similar pixel with smaller
 343 spectral difference possesses the higher weight.

344 2.3.4 Fit-FC

345 For capturing the temporal changes of fine pixels accurately, Fit-FC introduces a
 346 linear regression model established based on coarse images (Wang and Atkinson, 2018).
 347 A local linear regression model is firstly established between C_2 and C_1 within a
 348 moving window \mathbf{M} :

$$\mathbf{C}_2(\mathbf{M}) = a \times \mathbf{C}_1(\mathbf{M}) + b + \mathbf{R}(\mathbf{M}), \quad (12)$$

349 where $\mathbf{R}(\mathbf{M})$ are the coarse residuals in the moving window and a and b are the regres-
 350 sion coefficients. Then, the regression coefficients are applied to the fine pixels within
 351 a moving window corresponding to the coarse moving window (\mathbf{M}) for the RM (i.e.,
 352 Regression Model) prediction. Finally, unlike STARFM with the spectral difference
 353 between coarse and fine images, an another searching similar neighboring pixels ap-
 354 proach only with spatial distance using threshold is adopted to address the problem of
 355 blocky artifacts while considering the residuals of the regression model:

$$\hat{F}_2(x, y) = \sum_{i=1}^{n_k} W_i (a \times F_1(x_i, y_i) + b + r(x_i, y_i)), \quad (13)$$

356 where $r(x_i, y_i)$ is the residual at the fine pixel (x_i, y_i) , which is resampled from $\mathbf{R}(\mathbf{M})$ by
 357 bicubic interpolation.

358 2.3.5 OPDL

359 Dictionary-learning based methods reconstruct images with an overcomplete dic-
360 tionary and the corresponding coefficients of sparse representation (Huang and Song,
361 2012). Song and Huang (2013) proposed the dictionary-based learning method OPDL,
362 which requires only one image pair. The key idea of OPDL is that coarse image and
363 fine image acquired at the same location share the same sparse representation coeffi-
364 cients, and the overcomplete dictionary trained from images acquired at base time
365 should be time-invariant. Therefore, C_1 and F_1 provide the dictionary and C_2 provides
366 the corresponding coefficients to generate transition image T_2 . And with the same pro-
367 cess, T_1 can be also produced. Finally, the high-pass modulation is introduced to transfer
368 the temporal change from transition images to F_1 for prediction:

$$\hat{F}_2(x, y) = \frac{T_2(x, y)}{T_1(x, y)} F_1(x, y). \quad (14)$$

369 Due to the large spatial resolution difference between the fine image and coarse
370 image, OPDL is implemented in a two-layer framework (Song and Huang, 2013). The
371 first layer produces an image with the intermediate resolution between the coarse and
372 fine image. Subsequently, the second layer generates the final results using the image
373 synthesized by the first layer.

374 2.3.6 FSDAF

375 FSDAF (Zhu et al., 2016) is a hybrid method that combines unmixing, spatial in-
376 terpolation, and similar neighboring pixel smoothing for robust fusion results. Firstly,
377 similar to LMGM, FSDAF estimates the temporal change of a fine pixel (ΔF^{fp}) by an

378 unmixing-based method to produce the temporal prediction (F_2^{tp}), except that the un-
 379 mixing procedure is conducted in the whole image instead of a moving window. Then,
 380 with the TPS interpolation (Dubrule, 1984), the spatial prediction (F_2^{sp}) of F_2 can be
 381 generated. The residuals between the sum of ΔF and ΔC are considered in FSDAF:

$$R(x_i, y_i) = \Delta C(x_i, y_i) - \frac{1}{n} \left[\sum_{j=1}^n F_2^{tp}(x_{ij}, y_{ij}) - \sum_{j=1}^n F_1(x_{ij}, y_{ij}) \right], \quad (15)$$

382 where $R(x_i, y_i)$ is the residual in the coarse pixel at location (x, y) and n is the num-
 383 ber of fine pixels inside a coarse pixel and the fine pixel at location (x_i, y_i) is inside the
 384 coarse pixel at location (x, y) . In a homogenous area, the spatial prediction performs
 385 well, which is applied to calculate a new residual:

$$R_{ho}(x, y) = F_2^{sp}(x, y) - F_2^{tp}(x, y). \quad (16)$$

386 Thus, a weighted function (w_h) integrates two residuals (i.e., R_{ho} and R) using a
 387 homogeneity index for residual compensation. The final prediction of FSDAF can be
 388 expressed as:

$$\hat{F}_2(x, y) = F_1(x, y) + \sum_{i=1}^{n_s} W_i \left(\Delta F^{tp}(x_i, y_i) + n \times R(x_i, y_i) \times w_h(x_i, y_i) \right). \quad (17)$$

389 where W_i is the weight of similar pixel as same as Fit-FC.

390 2.4 Parameter settings of six spatiotemporal fusion methods

391 Referring to previous studies (Gao et al., 2006; Song and Huang, 2013; Rao et al.,
 392 2015; Wang et al., 2018; Zhu et al., 2016; Zurita-Milla et al., 2008), parameters of six
 393 spatiotemporal fusion methods were carefully tuned for different experimental sites and
 394 different resolution ratios. Table 4 shows the key parameters of UBDF, LMGM,
 395 STARFM, Fit-FC, and FSDAF. It is noted that we set the same values for the parameters

396 with similar functions in different methods to achieve a fair comparison (i.e., similar
 397 neighboring pixel smoothing in STARFM, Fit-FC and FSDAF). The key parameters of
 398 OPDL are separately shown in

399 Table 5 as they are very different from those of other five fusion methods. The
 400 patch size of dictionary representation in two layers was consistently set as 3 and 4 for
 401 all of the experimental sites and resolution ratios.

402

403 **Table 4** Key parameters of five fusion methods (c : class number, m : moving window
 404 size, n_s : number of similar neighboring pixels, m_s : moving window size for searching
 405 similar neighboring pixels, R : Spatial resolution ratio of coarse and fine images).

	c			m	n_s	m_s
	CIA	GWY	Tianjin			
UBDF	6	5	7	5×5	N/A	N/A
LMGM	6	5	7	5×5	N/A	N/A
STARFM	6	5	7	N/A	N/A	1.5× R +1
Fit-FC	N/A	N/A	N/A	3×3	1.5× R	1.5× R +1
FSDAF	6	5	7	N/A	1.5× R	1.5× R +1

406

407 **Table 5** Key parameters of OPDL method (resolution ratio is equal to the product of
 408 scale factors of two layers)

Dictionary size (Layer 1, Layer 2)		Resolution ratio (scale factors of two layers)			
		4 (2×2)	8 (2×4)	16 (4×4)	32 (4×8)
Experi- mental site	CIA	(1500,1500)	(700,1500)	(700,1500)	(50,1500)
	GWY	(900,1500)	(200,1500)	(200,1500)	(100,1500)
	Tianjin	(600,1500)	(200,1500)	(200,1500)	(50,1200)

409

410 **3. Results**

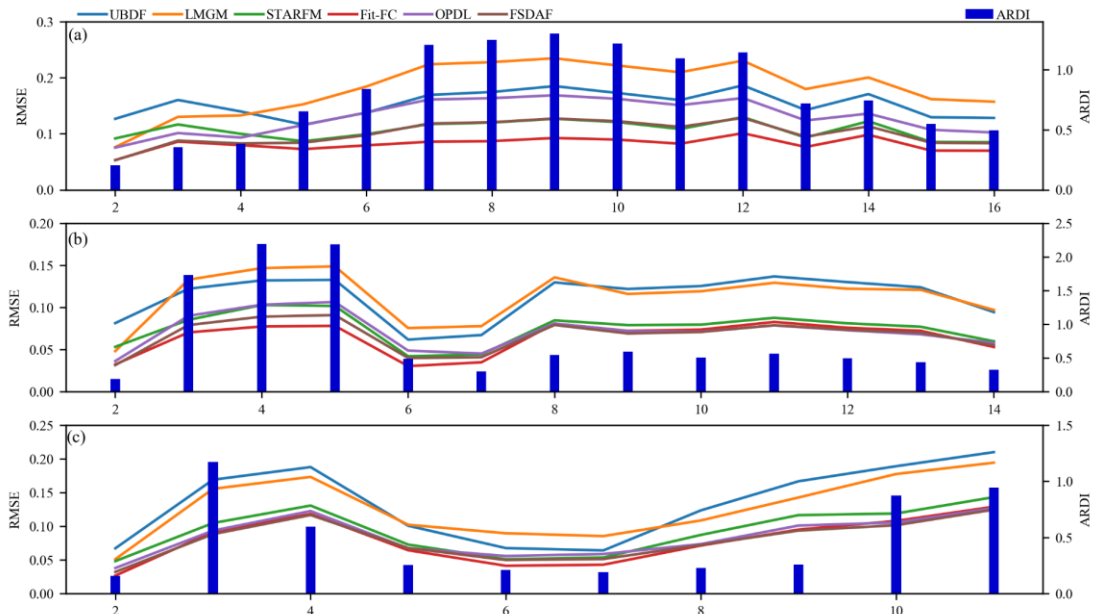
411 **3.1 Standard comparison**

412 The performances of the six methods at the two sites were evaluated with the ideal
 413 simulation data. Table 6 shows the averaged RMSE and r for each method. In general,
 414 Fit-FC performed best, followed by FSDAF. Among the other four methods, STARFM
 415 performed better than UBDF, LMGM, and OPDL. As shown in Fig. 8, all the methods
 416 performed worse when ADRI increased, while Fit-FC and FSDAF always generated
 417 better results for all the images in the time series than the other four methods.

418 **Table 6** Standard comparison evaluated by averaged RMSE, AD and r at the three sites.

		UBDF	LMGM	STARFM	Fit-FC	OPDL	FSDAF
CIA	RMSE	0.1533	0.1816	0.1292	0.0816	0.131	0.1006
	r	0.7606	0.7717	0.8883	0.8979	0.8555	0.8758
GWY	RMSE	0.1125	0.1133	0.0754	0.0643	0.0718	0.0669
	r	0.8681	0.8726	0.9196	0.9226	0.9072	0.9175
Tianjin	RMSE	0.1346	0.1296	0.0926	0.0788	0.0843	0.0797
	r	0.8769	0.8748	0.9376	0.9409	0.9329	0.9385

419



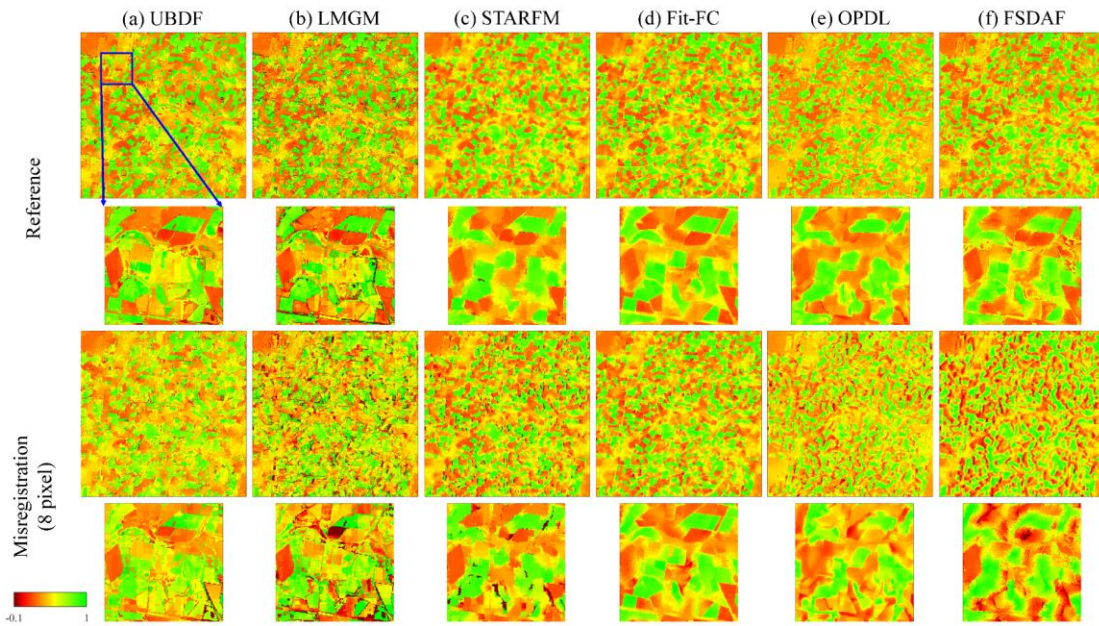
420

421 **Fig. 8.** The relationship between the prediction accuracy and the temporal variation at
422 the two sites: (a) CIA; (b) GWY; (c) Tianjin. Image number is the number of the pre-
423 dicted image in the image time series.

424 **3.2 Geometric misregistration**

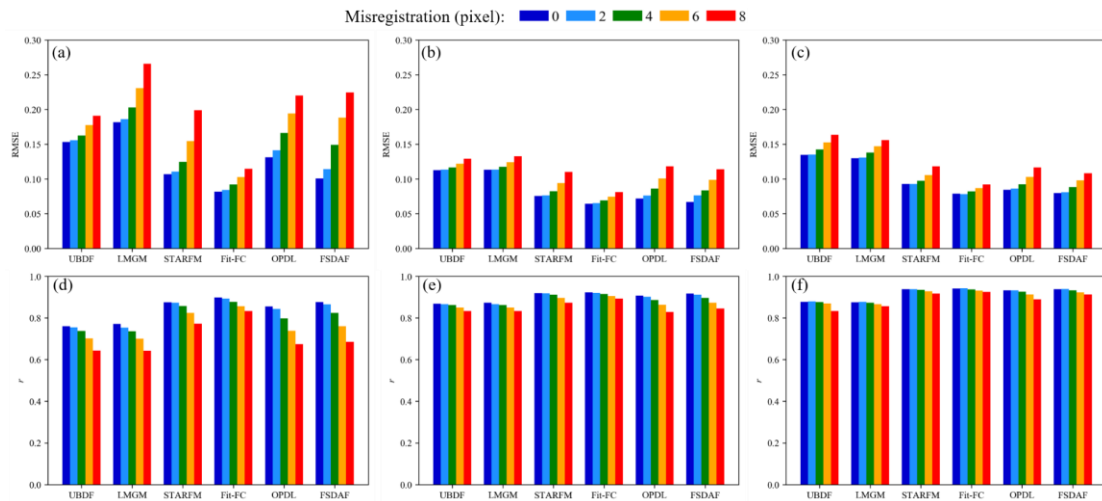
425 The performances of the six methods with the simulated misregistration errors be-
426 tween coarse and fine images are showed in this section. The extent of misregistration
427 was quantitatively measured as the shifting pixel distance. From visual comparison of
428 NDVI fusion results (Fig. 9), there are little distortions of the results fused by Fit-FC
429 under eight pixel shifting. However, the results fused by other five fusion methods are
430 obviously different from the reference results. When evaluated by quantitative indices
431 (Fig. 10), it is apparent that the accuracy of each method generally decreases as the
432 shifting distance increases. Fit-FC is the most robust method for misregistration fol-
433 lowed by UBDF, as their evaluation index values vary the most slowly. The other four
434 methods, LMGM, STARFM, OPDL, and FSDAF, are more sensitive to the geometric
435 error, as shown in Fig. 10 (a) and (b) where they all have sheer accuracy drops along in
436 addition to the shifting distance increases. The results of reflectance (Fig. A1, Fig. A4,
437 and Fig. A7) are similar to those of NDVI. [ZJ1]

438



439

440 **Fig. 9.** Using fusion results of the CIA site on February 13, 2002, visual comparison of
 441 NDVI results without geometric errors (e.g., reference) and with misregistration (eight
 442 pixel shifting) by six methods: (a) UBDF; (b) LMGM; (c) STARFM; (d) Fit-FC; (e)
 443 OPDL; (f) FSDAF.

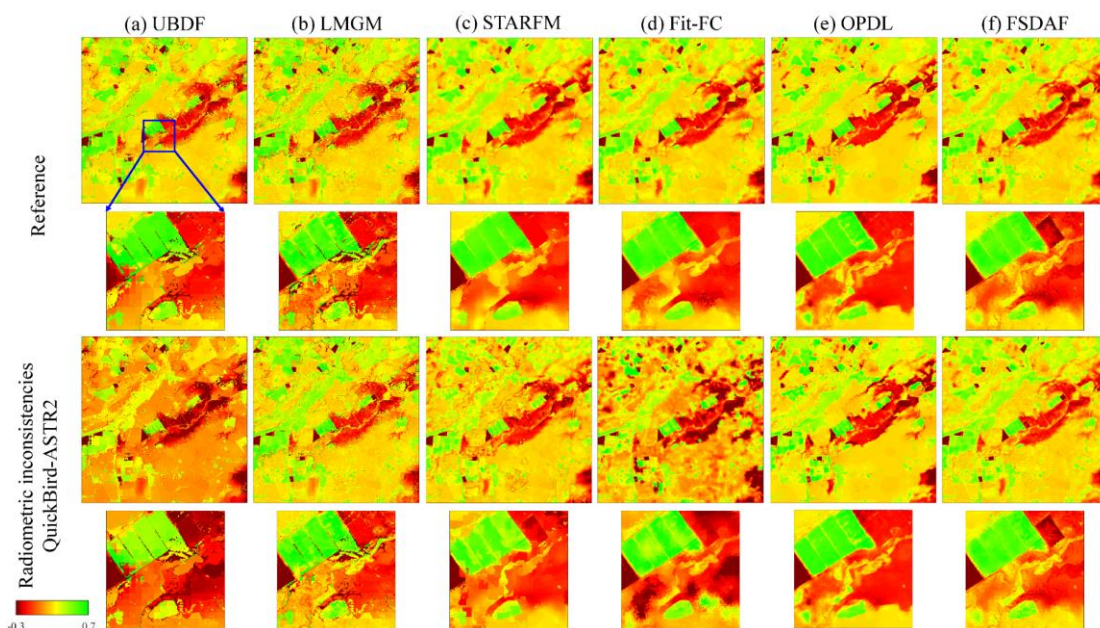


444

445 **Fig. 10.** Quantitative comparison of NDVI fusion results under different levels of geo-
 446 metric errors from 0 to 8 (misregistration pixel). (a) RMSE in CIA; (b) RMSE in GWY;
 447 (c) RMSE in Tianjin; (d) r in CIA; (e) r in GWY; (f) r in Tianjin.

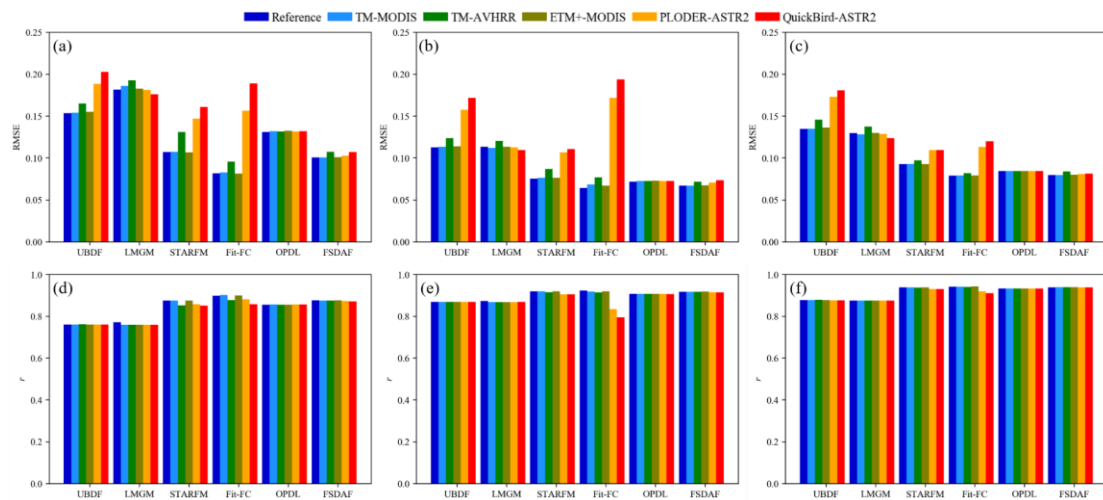
448 **3.3 Radiometric inconsistency**

449 Fig. 11 shows the robustness of different methods to the radiometric inconsistency
 450 between sensors (i.e., linear stretches of QuickBird-ASTR2). There are significant dis-
 451 tortions in the results fused by UBDF, STARFM, and Fit-FC. And the results evaluated
 452 by quantitative indices are in good agreement with that of visual comparison. When the
 453 fusion methods were applied to the datasets with small radiometric inconsistency (Table
 454 1), such as TM-MODIS, TM-AVHRR, and ETM+-MODIS, they all produced accurate
 455 results. However, when there were larger radiometric inconsistencies, like POLDER-
 456 ASTR2, and QuickBird-ASTR2, UBDF, STARFM, and Fit-FC showed larger errors
 457 than the other methods. In contrast, LMGM, OPDL and FSDAF are more robust to the
 458 radiometric inconsistency between two sensors. As for the results of reflectance (Fig.
 459 A2, Fig. A5, and Fig. A8), the sensitivity of fusion methods to radiometric errors is
 460 consistent with that of NDVI.



461
 462 **Fig. 11.** Using fusion results of the GWY site on December 12, 2004, visual comparison

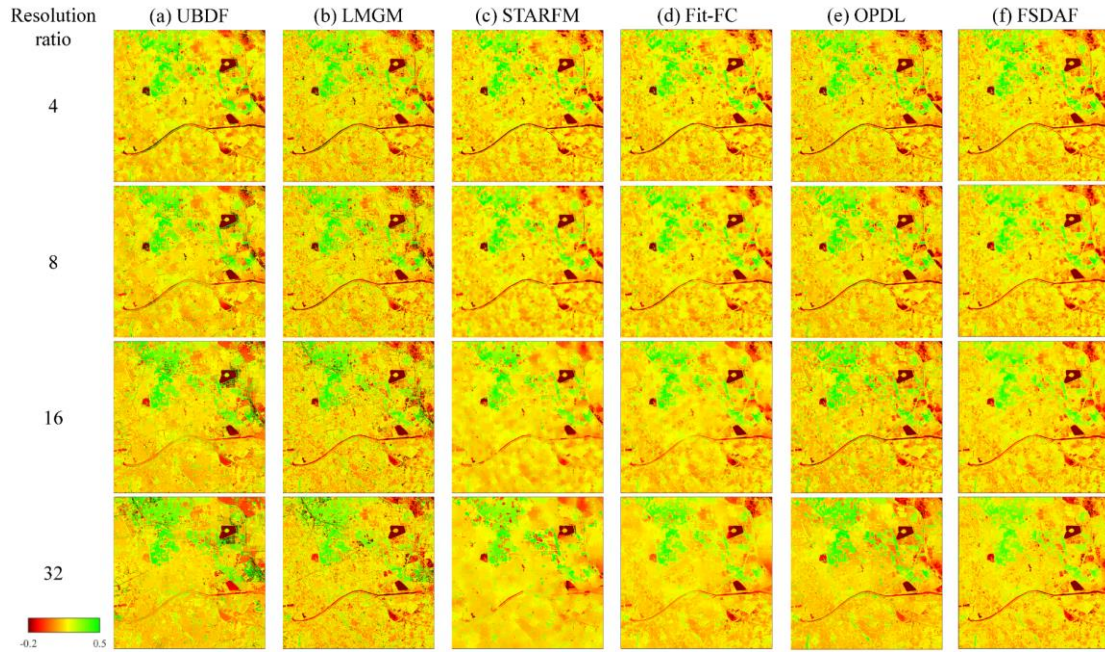
463 of NDVI results without any errors (e.g., reference) and with radiometric inconsistency
 464 by six methods: (a) UBDF; (b) LMGM; (c) STARFM; (d) Fit-FC; (e) OPDL; (f) FSDAF.



465
 466 **Fig. 12.** Quantitative comparison of NDVI fusion results under different levels of radi-
 467 ometric inconsistencies. (a) RMSE in CIA; (b) RMSE in GWY; (c) RMSE in Tianjin;
 468 (d) r in CIA; (e) r in GWY; (f) r in Tianjin. Reference means that there is no radiometric
 469 inconsistency.

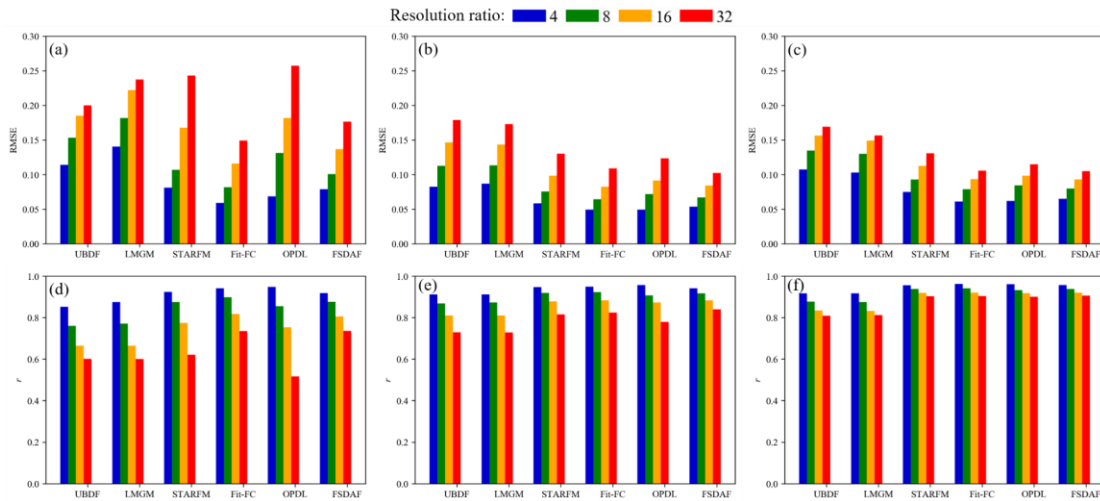
470 3.4 Spatial resolution ratio

471 Fig. 13 presents the accuracies of the six fusion methods in the scenarios of differ-
 472 ent spatial resolution ratios between coarse and fine images. In general, all the methods
 473 perform worse when the spatial resolution ratio increases. Among these methods,
 474 OPDL is the most sensitive to the spatial resolution ratio. The accuracy of the OPDL
 475 fusion results decreases the fastest as the spatial resolution ratio increases (Fig. 14).
 476 STARFM is also highly sensitive to the spatial resolution ratio, especially in heteroge-
 477 neous sites like CIA. In contrast, UBDF, LMGM, FSDAF, and Fit-FC are somehow
 478 less sensitive to the spatial resolution ratio. And the results of reflectance is similar to
 479 those of NDVI (Fig. A3, Fig. A6, and Fig. A9).



480

481 **Fig. 13.** Using NDVI fusion results of the Tianjin site on December 25, 2014, visual
 482 comparison under different levels of spatial resolution ratio from 4 to 32. by six meth-
 483 ods: (a) UBDF; (b) LMGM; (c) STARFM; (d) Fit-FC; (e) OPDL; (f) FSDAF.

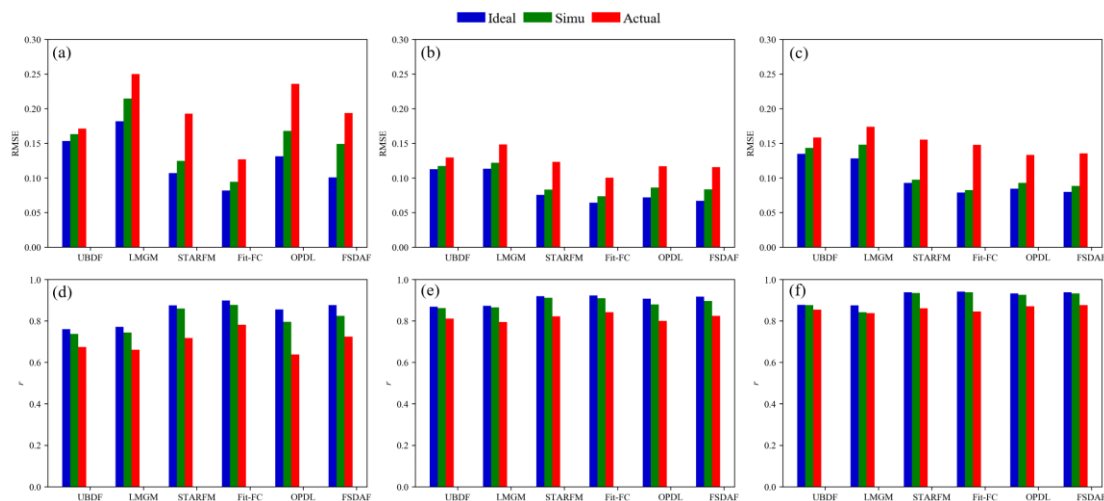


484

485 **Fig. 14.** Quantitative comparison of NDVI fusion results under different levels of spa-
 486 tial resolution ratio from 4 to 32. (a) RMSE in CIA; (b) RMSE in GWY; (c) RMSE in
 487 Tianjin; (d) r in CIA; (e) r in GWY; (f) r in Tianjin.

488 **3.5 Actual MODIS data**

489 The performances of the six fusion methods based on the actual MODIS images
 490 were compared with those based on the simulated MODIS images (Fig. 15). Generally,
 491 the former results were worse than those based on the simulated images with only par-
 492 tial errors, and further worse than those based on ideal images without any errors. This
 493 indicates that the simulated geometric and radiometric errors, to some extent, account
 494 for the performance differences between the experiments based on the ideal simulated
 495 MODIS data and those based on the actual MODIS data. [ZJ2]For the experiment based
 496 on actual MODIS data, Fit-FC performed best in CIA and GWY sites whereas OPDL
 497 performed best in Tianjin site. FSDAF ranked second in GWY and Tianjin sites,
 498 whereas UBDF ranked second in CIA site. These results imply the complexity of the
 499 fusion experiments based on actual data. However, in overall, Fit-FC and FSDAF still
 500 worth recommendation considering their relative stability.



501 **Fig. 15.** Quantitative comparison of NDVI fusion results based on actual MODIS im-
 502 ages. (a) RMSE in CIA; (b) RMSE in GWY; (c) RMSE in Tianjin; (d) r in CIA; (e) r in
 503 GWY; (f) r in Tianjin. Ideal, Simu, and Actual imply that the input coarse images are
 504

505 simulated ideally without any errors, simulated with geometric and radiometric errors,
 506 and are the actual MODIS images, respectively.

507 **4. Discussions**

508 To further explore the sensitivity of the six fusion methods to various factors, the-
 509 oretical derivations were conducted to analyze the geometric and radiometric error
 510 propagation from the input data to the results. As for the spatial resolution ratio, its
 511 influence was similar to the influence of the spatial heterogeneity of input data.

512 For the convenience of comparison, the different fusion methods (except OPDL)
 513 were grouped into three types, origin weighting (Eq. (18)), increment weighting (Eq.
 514 (19)), and regression weighting methods here (Eq. (20)):

$$\hat{F}_2(x, y) = \sum_i w_i C_{2i}, \quad (18)$$

$$\hat{F}_2(x, y) = F_1(x, y) + \sum_i w_i (C_{2i} - C_{1i}), \quad (19)$$

$$\hat{F}_2(x, y) = \sum_i w_i (a \times F_{1i} + b + r_i), \quad (20)$$

515 where $F_1(x_i, y_i)$, $C_1(x_i, y_i)$, $C_2(x_i, y_i)$, and $r(x_i, y_i)$ are denoted as F_{1i} , C_{1i} ,
 516 C_{2i} , and r_i for simplification, respectively. UBDF is a typical origin weighting
 517 method. As shown in Eq. (18), the fused result is calculated by weighting different
 518 coarse pixels acquired at t_2 ; the w_i is calculated by $[\mathbf{f}^T(\mathbf{M})\mathbf{f}(\mathbf{M})]^{-1}\mathbf{f}^T(\mathbf{M})$ for UBDF.
 519 LMGM, STARFM, and FSDAF belong to increment weighting methods. As shown in
 520 Eq. (19), the fused result is calculated by weighting the temporal increments from t_1 to
 521 t_2 of different coarse pixels; the w_i is calculated in different ways for different algo-
 522 rithms. Fit-FC is a novel developed regression weighting method. As shown in Eq. (20),

523 the fused result is calculated by weighing the linear transformation of fine pixels ac-
 524 quired at t_1 ; w_i is calculated based on a similar pixel smoothing strategy. For convenient
 525 theoretical analysis, Eq. (28) was further simplified by replacing r_i with $C_{2i} - aC_{1i} - b$

$$\begin{aligned}\hat{F}_2(x, y) &= \sum_i w_i (a \times F_{1i} + b + C_{2i} - aC_{1i} - b) \\ &= \sum_i w_i (a(F_{1i} - C_{1i}) + C_{2i})\end{aligned}\quad (21)$$

526 4.1 Propagation of geometric errors to fusion results

527 As the fine images were considered as the reference, the NDVI error induced by
 528 geometric errors could be expressed only in coarse images. Although the geometric
 529 error is a kind of systematic error, the induced NDVI error is random. Thus, the NDVI
 530 error on coarse pixel induced at t_1 and t_2 are assumed as random variables of δC_1 and
 531 δC_2 . Subsequently, although w_i is calculated in different ways by the five fusion meth-
 532 ods, it is mainly determined by the information of the fine pixels that are not affected
 533 by geometric errors. Thus, w_i could be considered as a constant in the error propagation
 534 procedure. The fusion errors ($\delta \hat{F}_2^G$) induced by geometric misregistration could be es-
 535 timated based on the error propagation equation. For UBDF, the fusion uncertainty of
 536 UBDF induced by geometric errors could be derived as:

$$\text{std}(\delta \hat{F}_2^G(x, y)) = \sqrt{\sum_i w_i^2 \text{var}(\delta C_2)}, \quad (22)$$

537 where std and var are the standard deviation and variance, respectively. Similarly, the
 538 standard deviation of fusion errors of LMGM, STARFM, and FSDAF could be derived
 539 as:

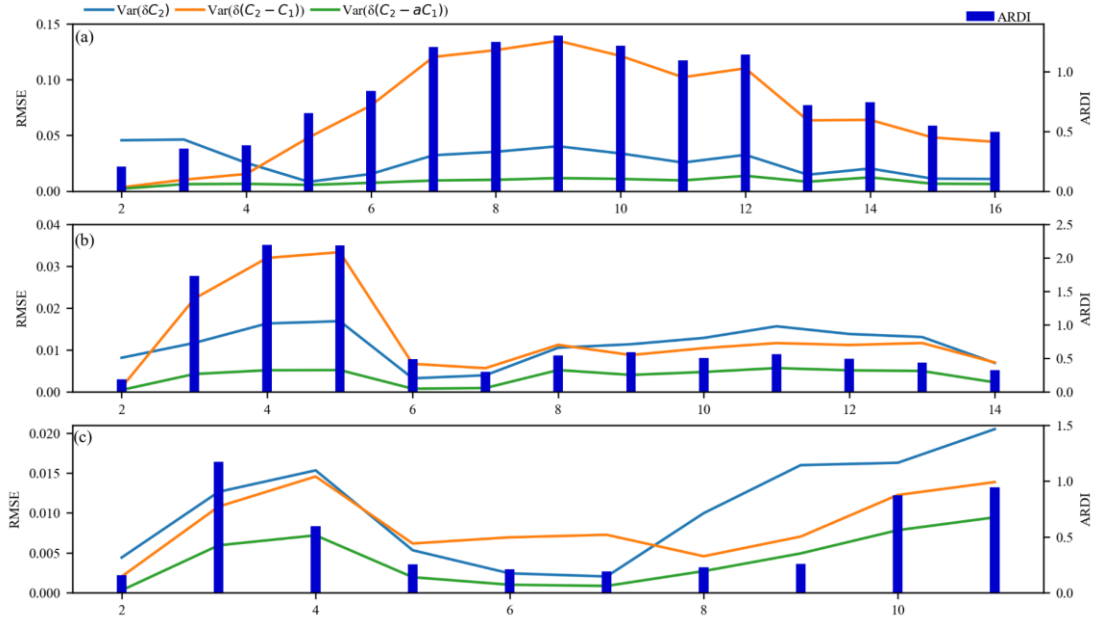
$$\begin{aligned}\text{std}(\delta\hat{F}_2^G(x, y)) &= \sqrt{\sum_i \text{var}(w_i\delta(C_2 - C_1))} \\ &= \sqrt{\sum_i w_i^2 (\text{var}(\delta C_2) + \text{var}(\delta C_1) - 2\text{cov}(\delta C_1, \delta C_2))},\end{aligned}\quad (23)$$

540 where cov is the covariance. If δC_1 and δC_2 are independent (i.e., the temporal
541 change between t_1 and t_2 is significant), the term $\text{cov}(\delta C_1, \delta C_2)$ approaches zero. Thus,
542 $\text{var}(\delta(C_2 - C_1))$ is larger than $\text{var}(\delta C_2)$ because of error accumulation, which is
543 also confirmed in the simulated data in most cases (Fig. 16). Therefore, LMGM,
544 STARFM and FSDAF are more sensitive to geometric errors than UBDF in general.
545 The standard deviation of fusion error of Fit-FC could be also derived as:

$$\begin{aligned}\text{std}(\delta F_2^G(x, y)) &= \sqrt{\sum_i \text{var}(w_i\delta(C_{2i} - aC_{1i}))} \\ &= \sqrt{\sum_i w_i^2 (\text{var}(\delta C_{2i}) + \text{var}(\delta(aC_{1i})) - 2\text{cov}(\delta C_{2i}, \delta(aC_{1i})))}.\end{aligned}\quad (24)$$

546 As a is the regression coefficient between C_1 and C_2 , $\text{var}(\delta(aC_{1i}))$ is strongly corre-
547 lated with δC_2 . Thus, $\text{var}(\delta(C_{2i} - aC_{1i}))$ is smaller than $\text{var}(\delta C_2)$ because of er-
548 ror compensation, which could also be shown in Fig. 16. Therefore, Fit-FC is the most
549 robust method for geometric error.

550 Analysis of the error propagation of OPDL is difficult due to the nonlinear opti-
551 mization in the dictionary learning procedure. The sensitivity to geometric error could
552 depend on different learned features, thus varied case by case.



553

554 **Fig. 16.** Comparison of variances of three weighting terms in the two sites: (a) CIA; (b)

555 GWY; (c) Tianjin, image number is the number of the predicted image in the image

556 time series. Fine images were shifted 8 pixels before aggregation.

557 4.2 Propagation of radiometric error to fused result

558 As radiometric inconsistency is usually a systematic error, linear stretch was used

559 to express radiometric inconsistency. Thus, the fusion error of UBDF induced by radi-

560 ometric inconsistency could be derived as:

$$\begin{aligned} \Delta \hat{F}_2^R(x, y) &= \sum_i w_i (\alpha C_{2i} + \beta) - \sum_i w_i C_{2i} \\ &= \sum_i w_i ((\alpha - 1) C_{2i} + \beta) \end{aligned} \quad (25)$$

561 where α and β are the coefficients for simulating radiometric inconsistency (i.e., slope

562 and intercept in Table 1). The fusion error of UBDF induced by radiometric incon-

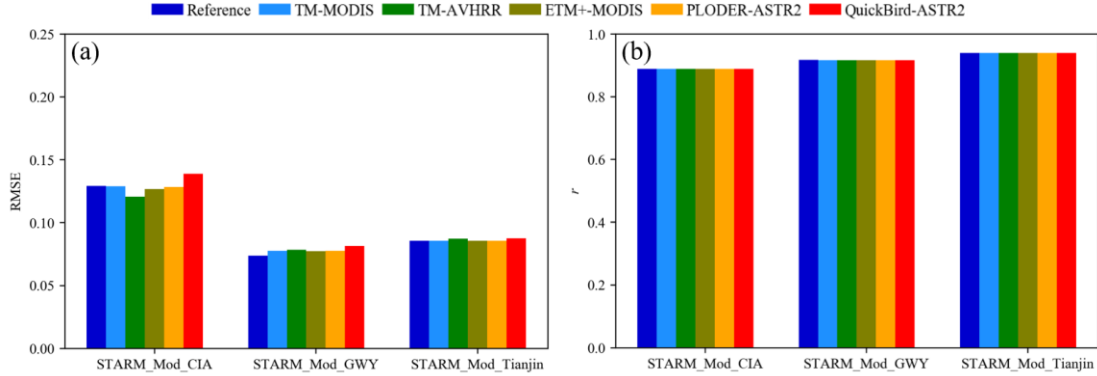
563 sistency depends linearly on two stretching parameters.

564 For STARFM, LMGM, and FSDAF, the fusion error induced by radiometric in-

565 consistency can be derived as:

$$\begin{aligned}\Delta \hat{F}_2^R(x, y) &= \sum_i w_i [(\alpha C_{2i} + \beta) - (\alpha C_{1i} + \beta)] - \sum_i w_i (C_{2i} - C_{1i}) \\ &= \sum_i w_i (\alpha - 1)(C_{2i} - C_{1i})\end{aligned}\quad (26)$$

566 Therefore, the intercept term (β) is removed in the term of ΔC . Theoretically, these
567 three methods are less sensitive to radiometric inconsistency compared to UBDF. How-
568 ever, STARFM shows high sensitivity to radiometric inconsistency in the experiments
569 (Fig. 11 and Fig. 12), which is somehow inconsistent with above theoretical analysis.
570 It is because the weight (w_i) calculation of the similar pixel smoothing in STARFM
571 includes a term of absolute NDVI difference between coarse and fine pixels (Eq. (11)),
572 which is sensitive to radiometric inconsistency. If the weight calculation in original
573 STARFM is modified as that in Fig-FC, the modified STARFM will be also robust to
574 radiometric inconsistency as the theoretical analysis (Fig. 17).



575 **Fig. 17.** Quantitative comparison of NDVI fusion results of STARFM_Mod_Site (i.e.,
576 results of modified STARFM in the different site) under different levels of radiometric
577 inconsistencies. (a) RMSE; (b) r . Reference means that there is no radiometric incon-
578 sistencies. (a) RMSE; (b) r . Reference means that there is no radiometric incon-
579 sistency.

580 Similarly, for Fit-FC, the fusion error induced by radiometric inconsistency can be
581 expressed as:

$$\begin{aligned}\Delta\hat{F}_2^R(x, y) &= \sum_i w_i (a(F_{1i} - \alpha C_{1i} - \beta) + \alpha C_{2i} + \beta) - \sum_i w_i (a(F_{1i} - C_{1i}) + C_{2i}) \\ &= \sum_i W_i ((\alpha - 1)(C_{2i} - aC_{1i}) + (1 - a)\beta)\end{aligned}\quad (27)$$

582 Compared with the second group (LMGM, STARFM, and FSDAF), Fit-FC is more
 583 sensitive to radiometric inconsistency because α and β both influence the fusion result.
 584 Subsequently, OPDL is robust to radiometric inconsistency because it employs a linear
 585 regression model for intercalibration of coarse and fine images.

586 4.3 Influence of spatial resolution ratio on spatiotemporal fusion

587 The spatial resolution ratio of sensors determines the information gap between
 588 coarse and fine images acquired at the same time. In other words, with the spatial res-
 589 olution ratio increasing, coarse pixels contain more fine pixels and, thus, become more
 590 mixed; this is a similar effect as the increase of spatial heterogeneity. Thus, those meth-
 591 ods that perform relatively better in heterogonous images should also be less sensitive
 592 to the spatial resolution ratio. As the unmixing module employed in fusion methods can
 593 better capture the spatial heterogeneity, UBDF, LMGM, and FSDAF, which employ the
 594 unmixing module, are less sensitive to the spatial resolution ratio than STARFM and
 595 OPDL. Fit-FC is also relatively less sensitive to the spatial resolution ratio although the
 596 unmixing module is not employed in this method. It is because only two land cover
 597 types (i.e., vegetation and non-vegetation) need to be considered in NDVI fusion; there-
 598 fore, the linear regression model in Fit-FC with two degrees of freedom (i.e., two coef-
 599 ficients a and b) plays a similar unmixing role, which is adequate in capturing the tem-
 600 poral changes of the two land cover types. Furthermore, it implies that Fit-FC is partic-
 601 ularly more suitable for the spatiotemporal fusion of NDVI data than reflectance data.

602 4.4 Method selection and guidance for future method design

603 The above comparison and analyses can guide the selection of suitable spatiotem-
604 poral fusion methods in applications. Other than the influential factors of geometric
605 misregistration, radiometric inconsistency, and spatial resolution ratio, the perfor-
606 mances of the fusion methods mainly depend on the spatiotemporal variations of input
607 datasets. The selection of a suitable method should, therefore, consider the influence
608 extent of all the factors and balance the pros and cons of each method according to the
609 characteristics of their data and applications. Similar to a recent comparative study (Liu
610 et al., 2019), Fit-FC and FSDAF were shown to have better performances than the other
611 three methods (i.e., UBDF, STARFM, and OPDL) for the actual MODIS data, indicat-
612 ing that Fit-FC and FSDAF are robust to different spatiotemporal variations. For a com-
613 prehensive comparison, the advantages and disadvantages of the six fusion methods are
614 summed up in Table 7. The most recommended algorithm is Fit-FC, which can produce
615 accurate results with high efficiency for NDVI fusion. However, it should be noted that
616 Fit-FC needs to be implemented with radiometric normalization (Gao et al, 2010;
617 Gevaert and García-Haro, 2015) considering its sensitivity to systematic radiometric
618 error. FSDAF is another favorable method with high accuracy if geometric misregistra-
619 tion can be well corrected.

620 **Table 7** The pros and cons of six typical fusion methods under comparison of different
621 influential factors (worst: 1, good: 2~4, best: 5). Due to the dominant of spatiotemporal
622 variations in the fusion method performances, a triple weight has been used in the cal-
623 culations of the total scores. (Variations = Spatiotemporal variations, Ratio = Spatial

624 Resolution ratio)

Method	Variations	Geometric	Radiometric	Ratio	Total
UBDF	2	3	2	3	14
LMGM	1	2	5	3	13
STARFM	3	2	3	1	15
Fit-FC	5	5	1	5	26
OPDL	3	3	5	1	18
FSDAF	4	2	5	3	22

625 This study can also give guidance for the future development of spatiotemporal
626 fusion methods. Previous developments of spatiotemporal fusion methods were gener-
627 ally designed without the consideration of inevitable geometric and radiometric errors.
628 For example, increment weighting (Eq. (19)) is commonly used in a large group of
629 fusion methods (e.g., STARFM, LMGM, and FSDAF) as it can keep good spatial de-
630 tails and reduce the radiometric inconsistency of sensors to some extent. However, the
631 above analysis indicates that it would be highly sensitive to geometric error. In contrast,
632 the regression model employed in Fit-FC is resistant to geometric errors, whereas, it is
633 sensitive to radiometric inconsistency. Therefore, combining the strength of Fit-FC and
634 increment weighting might be a promising strategy in the future development of novel
635 methods; other techniques that can mitigate these errors should also be taken into con-
636 sideration.

637 It should be noted that this study has not completely considered all the influential
638 factors. The geometric and radiometric errors were simply simulated by pixel shifting
639 and linear transformation in this study. However, there are more complicated errors

640 between sensors, including complex geometric errors from imagery scaling, rotation,
641 and skewing (Dai and Khorram, 1998; Toutin, 2004) and radiometric inconsistency
642 caused by nonlinear distortion, such as analogous bands between sensors with different
643 spectral response functions (SRFs), radiometric resolution difference and the angle ef-
644 fect that solar-sensor geometry bidirectional reflectance distribution function (BRDF)
645 changes over time (Chander et al., 2013b; Gao et al., 2006; Roy et al., 2008). These
646 errors could cause large uncertainties in the fusion results. This is shown in the actual
647 MODIS experiments and should be considered carefully in the future. The selection of
648 typical fusion methods might be another issue. It is impossible in this study to compare
649 all of the spatiotemporal fusion methods due to limitations of the source code availa-
650 bility and heavy works. Notwithstanding the representative methods as much as possi-
651 ble that we selected, the better methods are probably missed. An organization of pro-
652 gramming contest with a standard dataset and assessment protocol could be a solution
653 to engaging more algorithm developers and a fair comparison of different spatiotem-
654 poral fusion methods in near future.

655 **5. Conclusions**

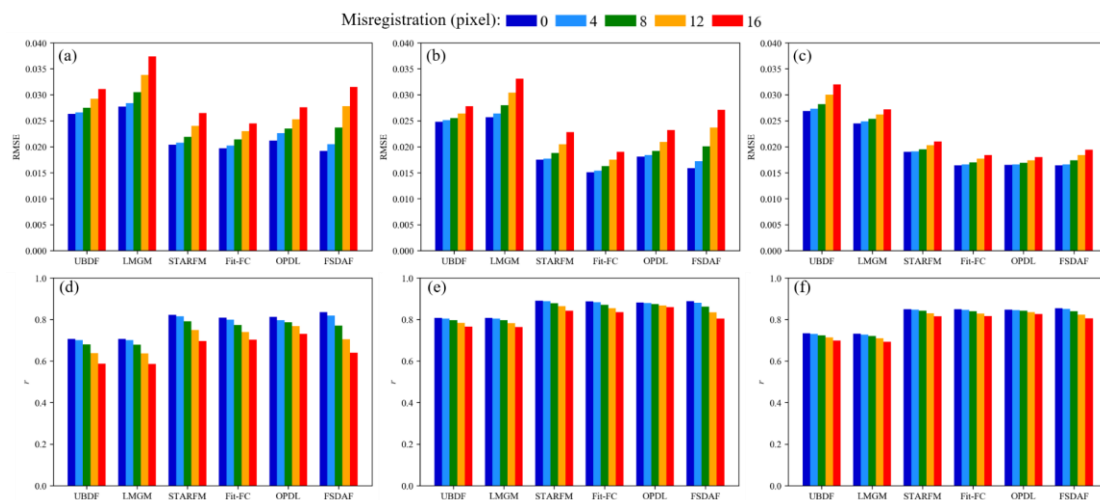
656 Besides the spatiotemporal variations of input datasets, this study presents the ne-
657 cessity of considering the sensitivity of fusion methods to three influential factors (i.e.,
658 geometric misregistration, radiometric inconsistency, and spatial resolution ratio) when
659 they are employed in real applications. These influencing factors could affect different
660 fusion methods to different degrees. The simulation experiment and the theoretical anal-
661 ysis showed that Fit-FC achieved the best performances for both sites with the best

662 resistance to geometric errors among the six typical spatiotemporal fusion methods
 663 when the radiometric inconsistency between sensors was negligible, suggesting it is the
 664 first recommended algorithm for NDVI time-series reconstruction. However, Fit-FC is
 665 sensitive to systematic radiometric error and, thus, performs poorly if there is a signif-
 666 icant radiometric inconsistency between the two sensors. FSDAF could also generate
 667 satisfactory results through its ability to reduce radiometric inconsistency; however, it
 668 is sensitive to geometric errors. Therefore, precise geometric registration is required
 669 when using FSDAF. These findings could not only help users to select suitable methods
 670 according to the characteristics of their data and applications but could also provide
 671 guidance for developers in designing novel algorithms more robust to different influ-
 672 ential factors in the future.

673 Acknowledgement

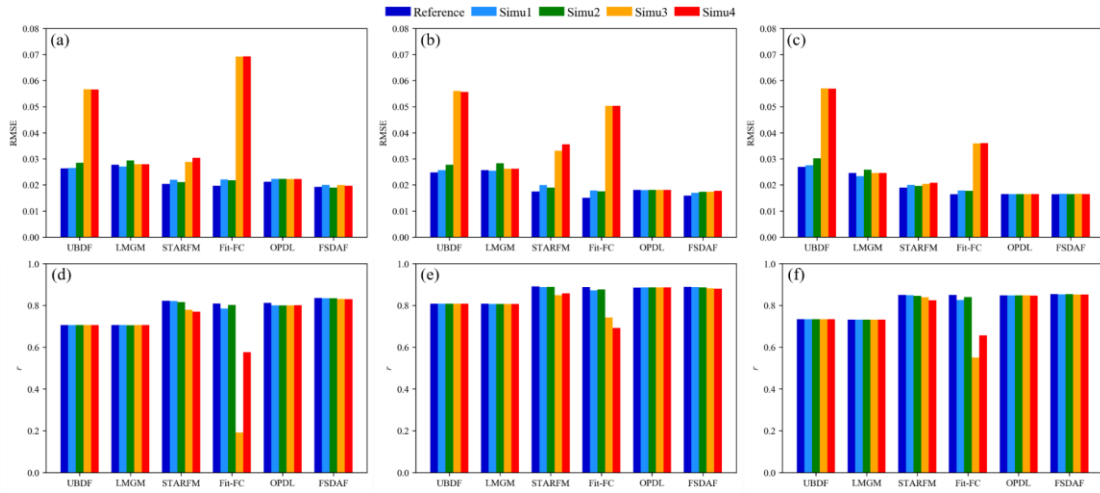
674 This study is supported by National Natural Science Foundation of China (No.
 675 67441830108 and No. 41871224). [ZJ3]

676 Appendix

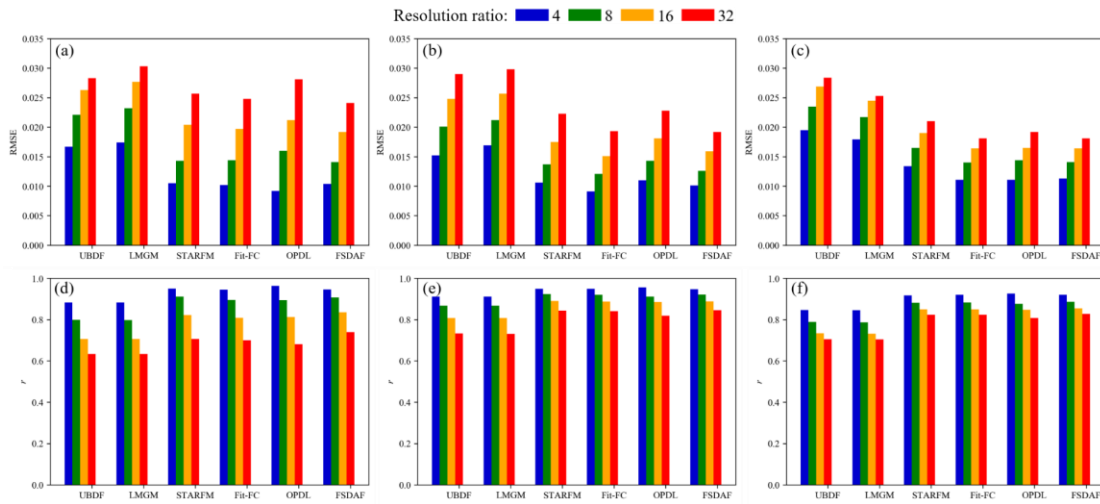


677
 678 **Fig. A1.** Quantitative comparison of green band fusion results under different levels of

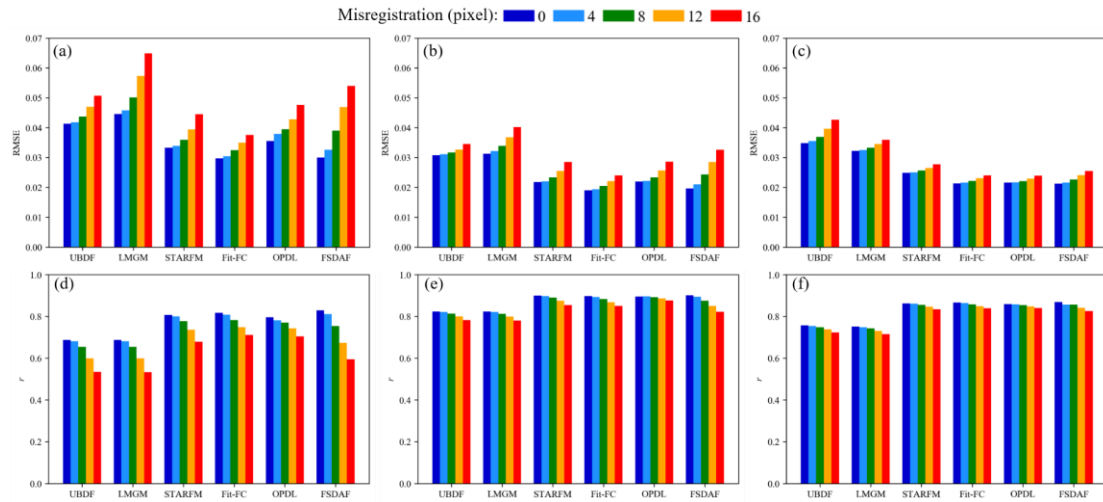
679 geometric errors from 0 to 16 (misregistration pixel). (a) RMSE in CIA; (b) RMSE in
 680 GWY; (c) r in CIA; (d) r in GWY.



681
 682 **Fig. A2.** Quantitative comparison of green band fusion results under different levels of
 683 radiometric inconsistencies. (a) RMSE in CIA; (b) RMSE in GWY; (c) r in CIA; (d) r
 684 in GWY. Reference means that there is no radiometric inconsistency.

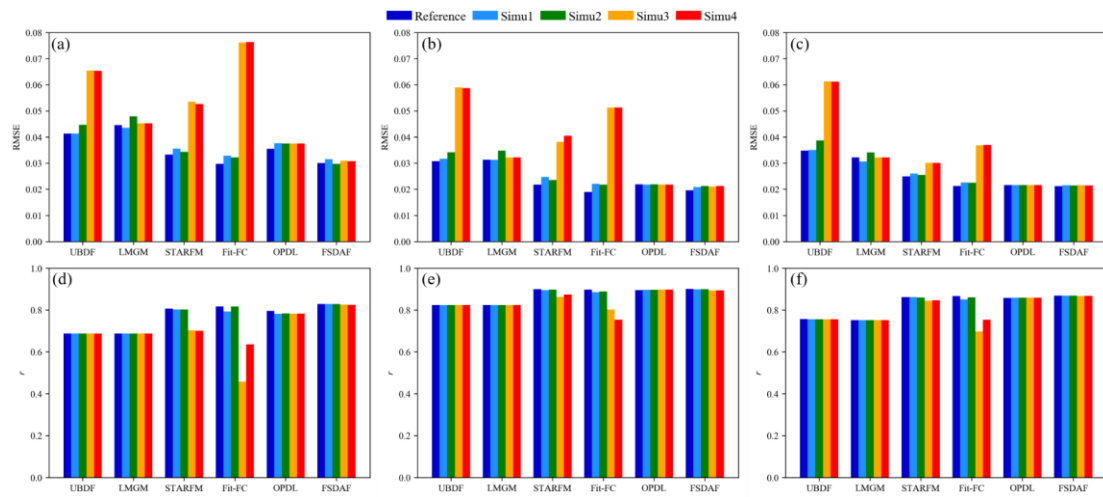


685
 686 **Fig. A3.** Quantitative comparison of green band fusion results under different levels of
 687 spatial resolution ratio from 4 to 32. (a) RMSE in CIA; (b) RMSE in GWY; (c) r in
 688 CIA; (d) r in GWY.



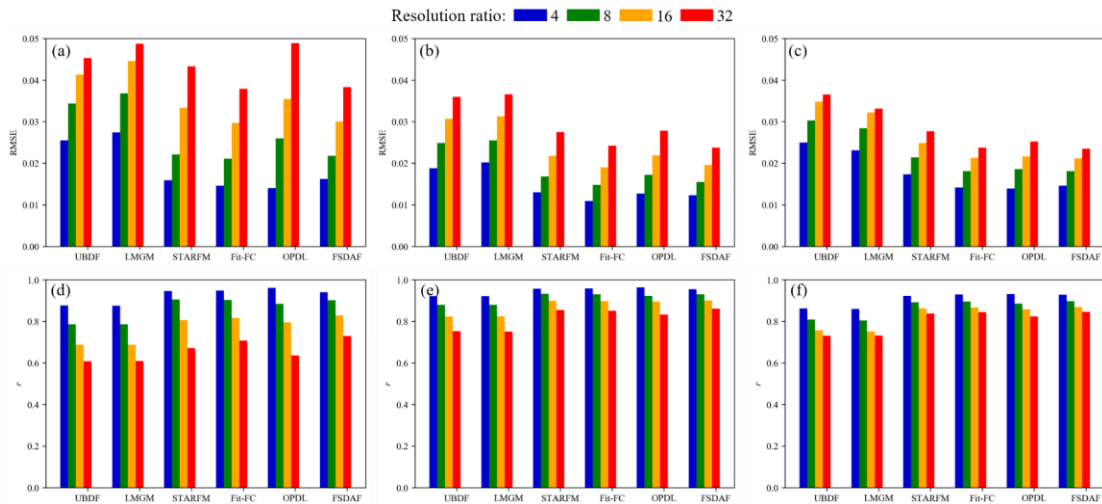
689

690 **Fig. A4.** Quantitative comparison of red band fusion results under different levels of
 691 geometric errors from 0 to 16 (misregistration pixel). (a) RMSE in CIA; (b) RMSE in
 692 GWY; (c) r in CIA; (d) r in GWY.



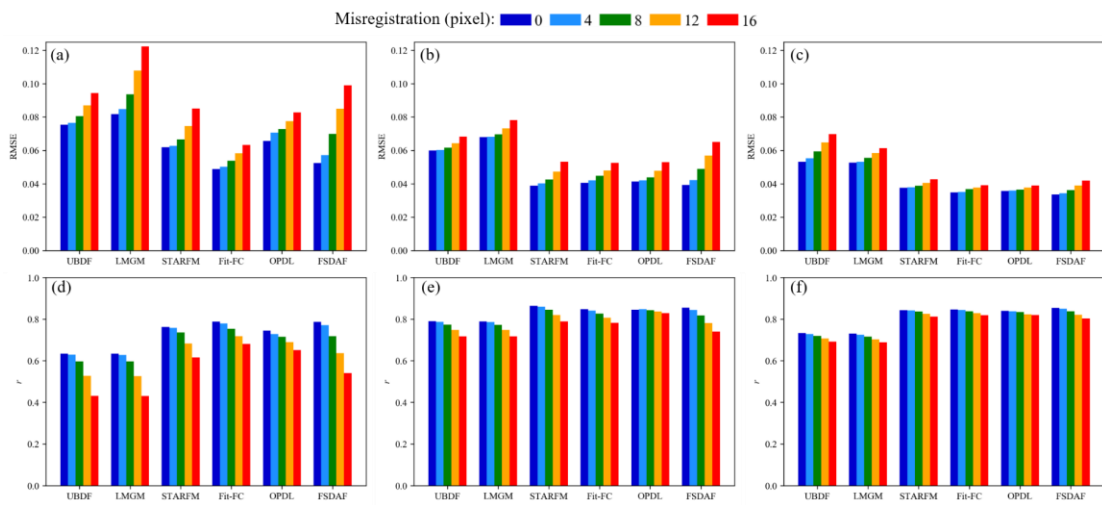
693

694 **Fig. A5.** Quantitative comparison of red band fusion results under different levels of
 695 radiometric inconsistencies. (a) RMSE in CIA; (b) RMSE in GWY; (c) r in CIA; (d) r
 696 in GWY. Reference means that there is no radiometric inconsistency.



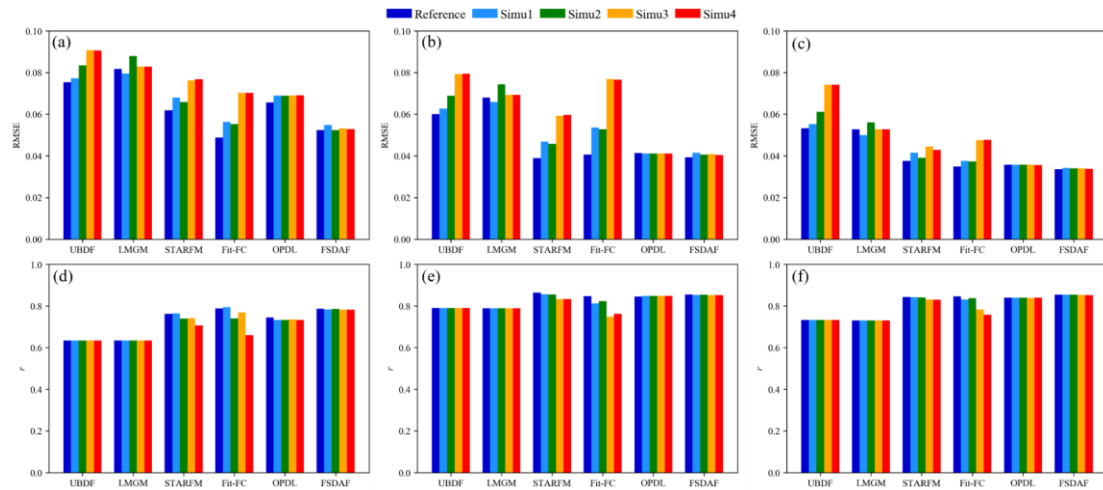
697

698 **Fig. A6.** Quantitative comparison of red band fusion results under different levels of
 699 spatial resolution ratio from 4 to 32. (a) RMSE in CIA; (b) RMSE in GWY; (c) r in
 700 CIA; (d) r in GWY.



701

702 **Fig. A7.** Quantitative comparison of NIR band fusion results under different levels of
 703 geometric errors from 0 to 16 (misregistration pixel). (a) RMSE in CIA; (b) RMSE in
 704 GWY; (c) r in CIA; (d) r in GWY.

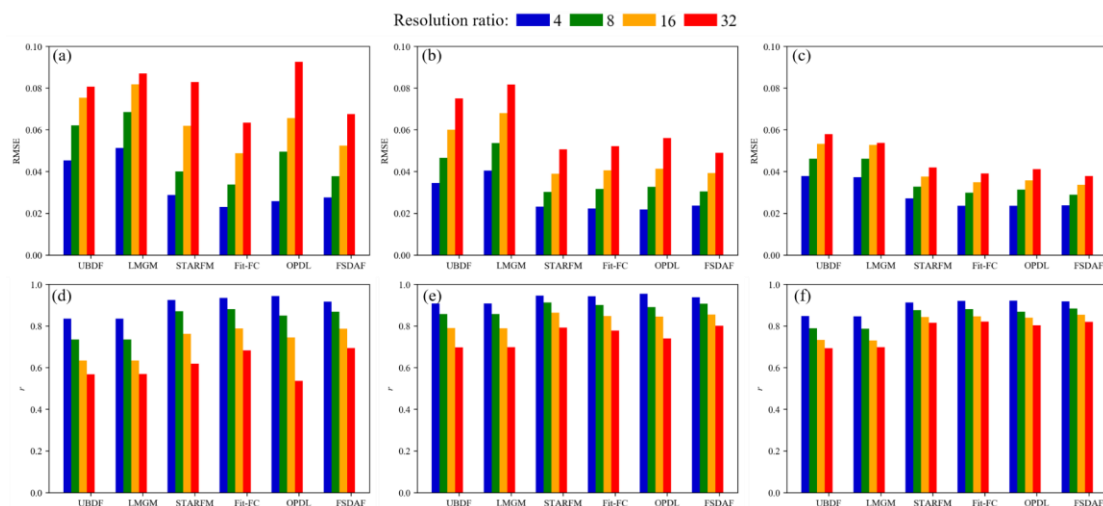


705

706 **Fig. A8.** Quantitative comparison of NIR band fusion results under different levels of

707 radiometric inconsistencies. (a) RMSE in CIA; (b) RMSE in GWY; (c) r in CIA; (d) r

708 in GWY. Reference means that there is no radiometric inconsistency.



709

710 **Fig. A9.** Quantitative comparison of NIR band fusion results under different levels of

711 spatial resolution ratio from 4 to 32. (a) RMSE in CIA; (b) RMSE in GWY; (c) r in

712 CIA; (d) r in GWY.

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